

A STATISTICAL STUDY OF LOW FREQUENCY ULF WAVES AS MEASURED BY SPACE-BORNE AND GROUND-BASED MAGNETOMETERS

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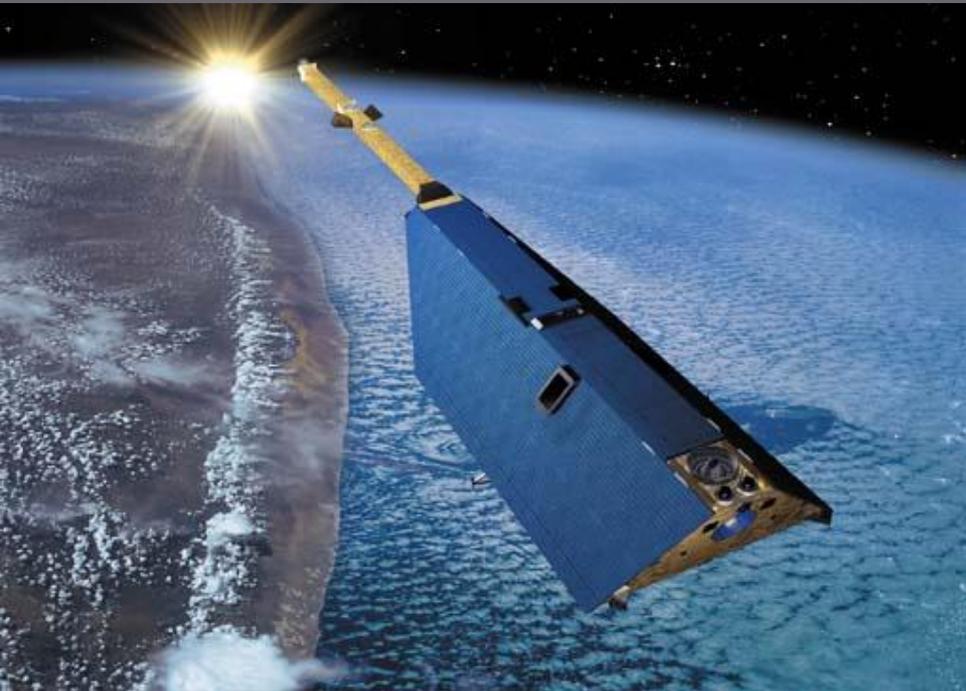
2 Department of Physics, University of Athens, Athens, Greece



Geospace Revisited, 15-20 September 2014,
Rhodes, Greece



CHAMP (CHALLENGING MINISATELLITE PAYLOAD)



Launch: July 15, 2000

Design Lifetime: 5 years

End of Mission: September 19
2010

*(ten years, two month, four days)
(58277 orbits)*

Orbit Characteristics:

- Low Earth Orbit (initial altitude of 454 km)
- Almost circular
- Near polar (87° inclination)
- Approx. 90' per Orbit

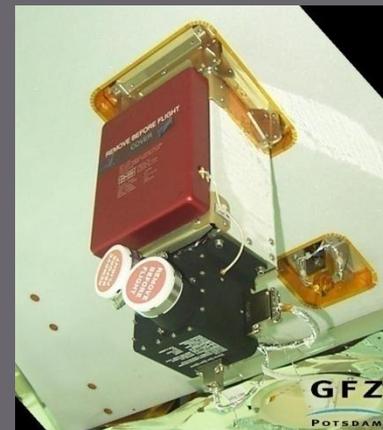


Overhauser Magnetometer



Range	18 000 - 65 000 nT
Resolution	10 pT
Noise level	< 50 pT (rms)
-3 dB bandwidth	0.28 Hz
Sample rate	1 Hz
Absolute accuracy	< 0.5 nT

Digital Ion Drift Meter



Range of ion density	$10^8 - 10^{12}$ ions/m ³
Range of ion temperature	200 - 55 000 K
Range of drift velocity	0 - 6 km/s
Resolution of ion velocity	< 1° direction, < 130 m/s speed
Sample rates	
DM mode	0, 1, 2, 4, 8, 16 Hz
RPA mode	0, 8, 16 Hz
PLP mode	0, 1/15 Hz



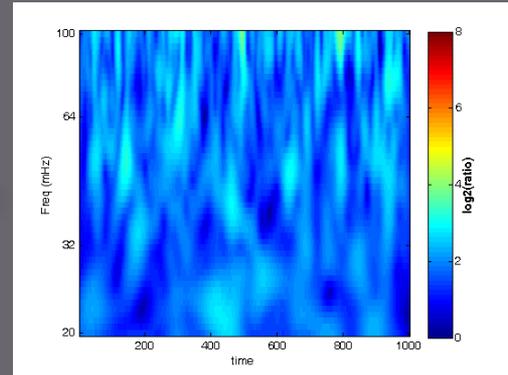
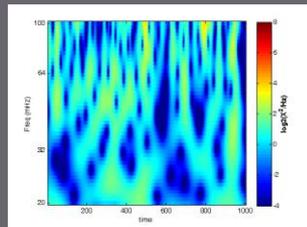
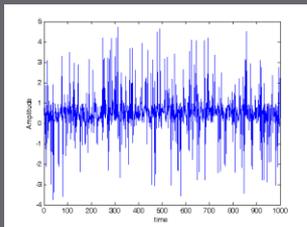
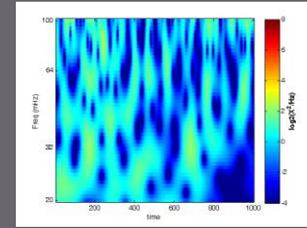
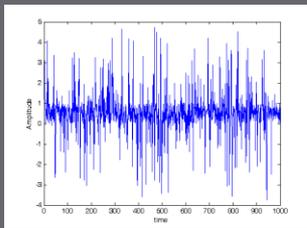
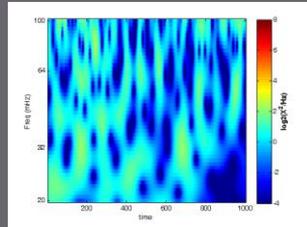
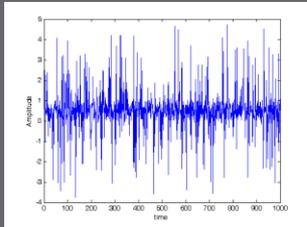
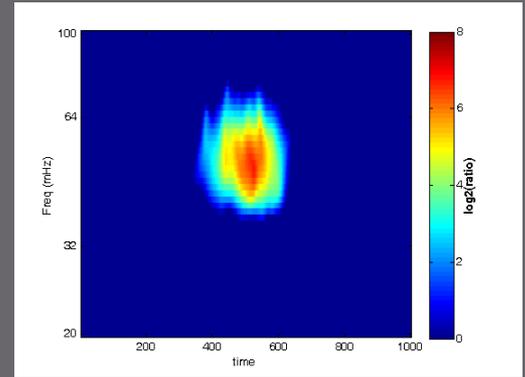
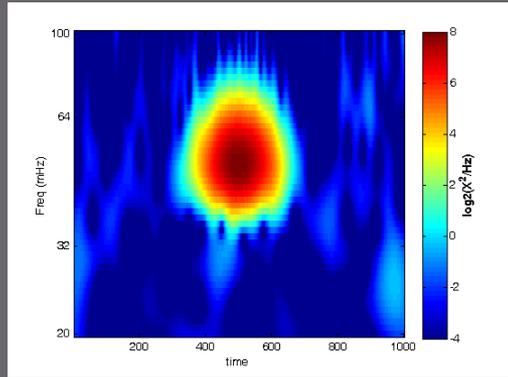
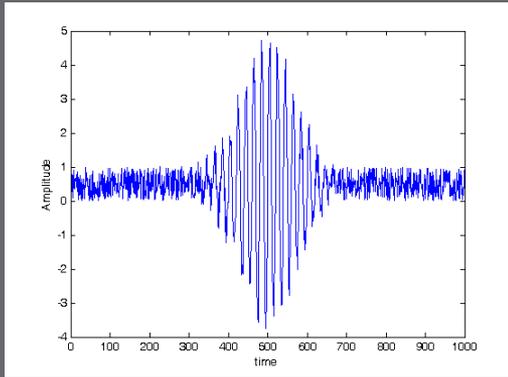
Methodology

- Read Data (daily files) in a day-to-day basis
(+/- a few minutes to avoid edge effects)
- Interpolate (small) gaps
- Filter (cutoff at 20 mHz)
- Wavelet Transform
(logarithmically spaced frequencies
from 20 to 100 mHz)
- Track Segmentation
[-55, +55] degrees in Mag. Lat.
- Event Detection
(statistical thresholding)
- Parameter Extraction
- Database Creation
- **Post Processing**

Balasis et al. 2013; Earth, Planets and Space



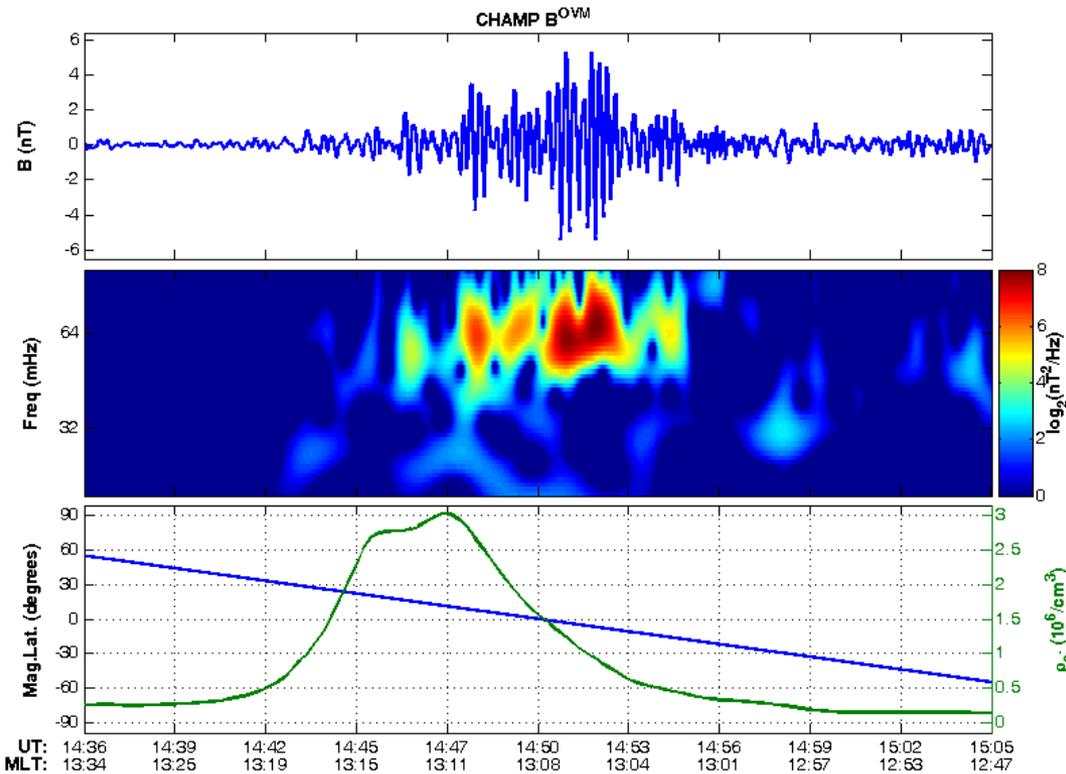
SHUFFLED SERIES



MEAN + 3 ST.DEV

Ignoring largest 10% points to minimize spike influence

Date: 31 Oct 2003



$$t_{\text{peak}} = 14:42$$
$$f_{\text{peak}} = 79 \text{ mHz}$$

$$\Delta t \sim 23'$$

$$\text{Power}_{\text{peak}} = 2^8 \text{ nT}^2/\text{Hz}$$

$$\text{Max}(|B|) = 5.45 \text{ nT}$$

$$\text{Max}(|\Delta B|) = 2.73 \text{ nT/sec}$$

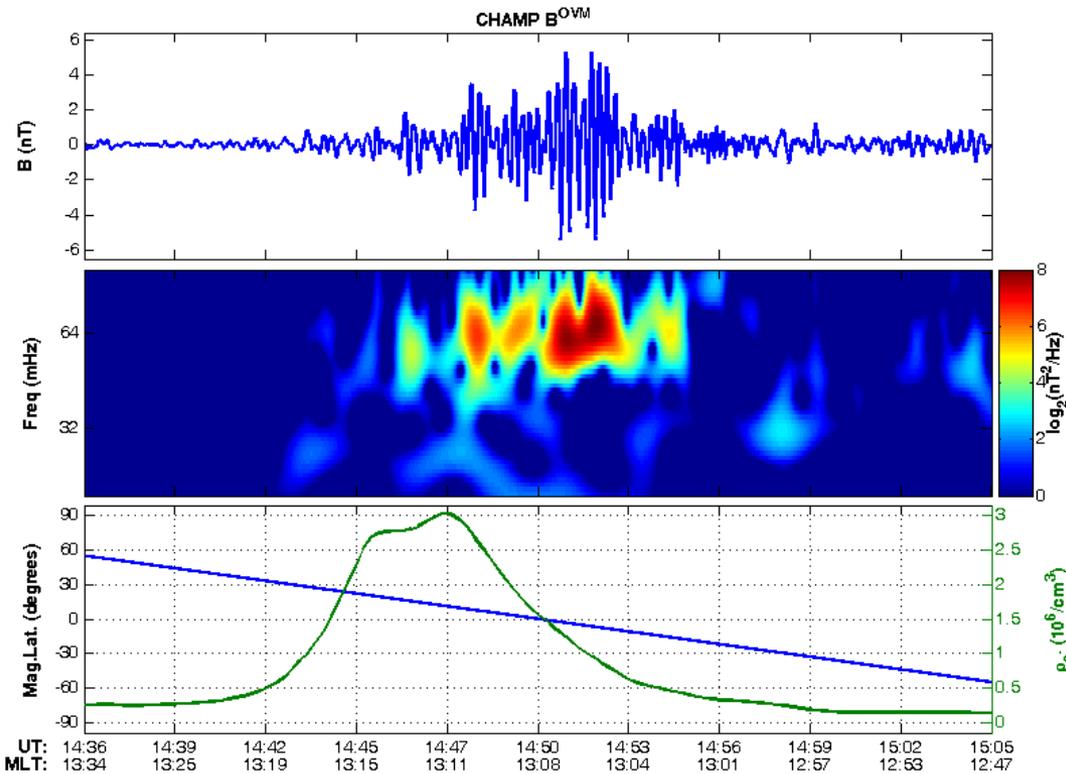
$r_{\text{MAG}} = (-7^\circ, 40^\circ, 401 \text{ km}), \text{MLT} = 13:06, \text{Lshell}$
(ONERA IRBEM LIB)

$$\text{roughness}(n) \sim 5 \cdot 10^3 \text{ cm}^{-3}$$

Balasis et al. ULFwave activity during the 2003 Halloween superstorm: multipoint observations from CHAMP, Cluster and Geotail missions, *Ann. Geophys.*, 30, 1751-1768, doi:10.5194/angeo-30-1751-2012



Date: 31 Oct 2003



$r_{MAG} = (-7^\circ, 40^\circ, 401 \text{ km})$
 MLT = 13:06

$t_{peak} = 14:42$
 $f_{peak} = 79 \text{ mHz}$

NOT AT THE EDGES

$\Delta t \sim 23'$

$\Delta t > 2 (1/f_{peak})$

$Power_{peak} = 2^8 \text{ nT}^2/\text{Hz}$

EXCLUDE WEAK EVENTS

$Max(|B|) = 5.45 \text{ nT}$

$Max(|\Delta B|) = 2.73 \text{ nT/sec}$

REMOVE SPIKES

$Max(|\Delta B|) > 0.6 * Max(|B|) + 0.7$

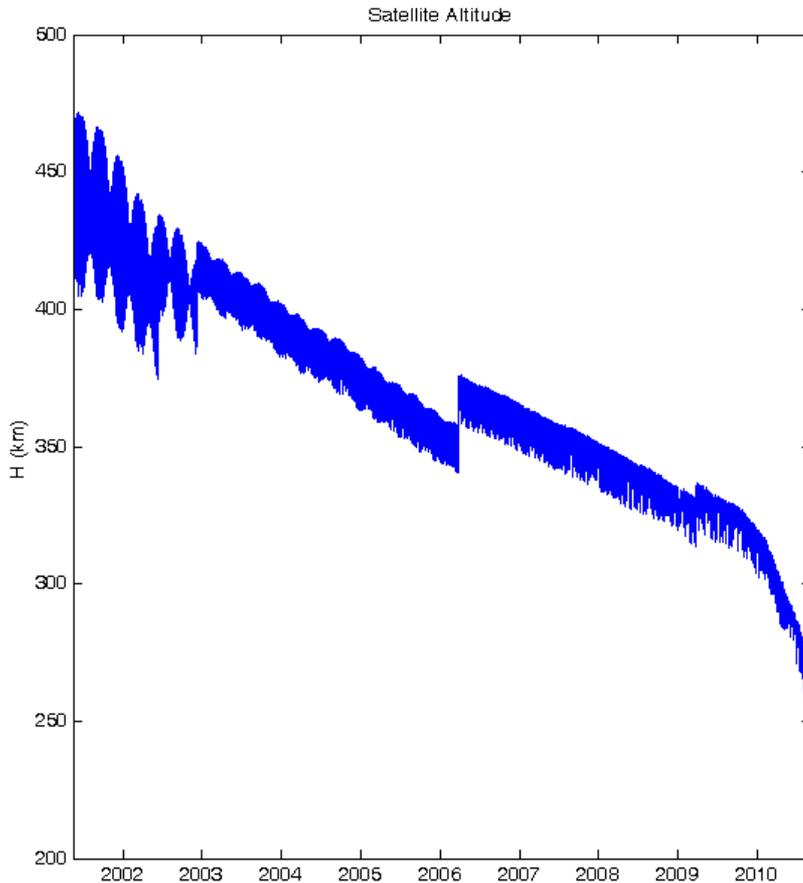
$roughness(n) \sim 5 \cdot 10^3 \text{ cm}^{-3}$

PLASMA INSTABILITIES

Balasis et al. ULFwave activity during the 2003 Halloween superstorm: multipoint observations from CHAMP, Cluster and Geotail missions, Ann. Geophys., 30, 1751-1768, doi:10.5194/angeo-30-1751-2012



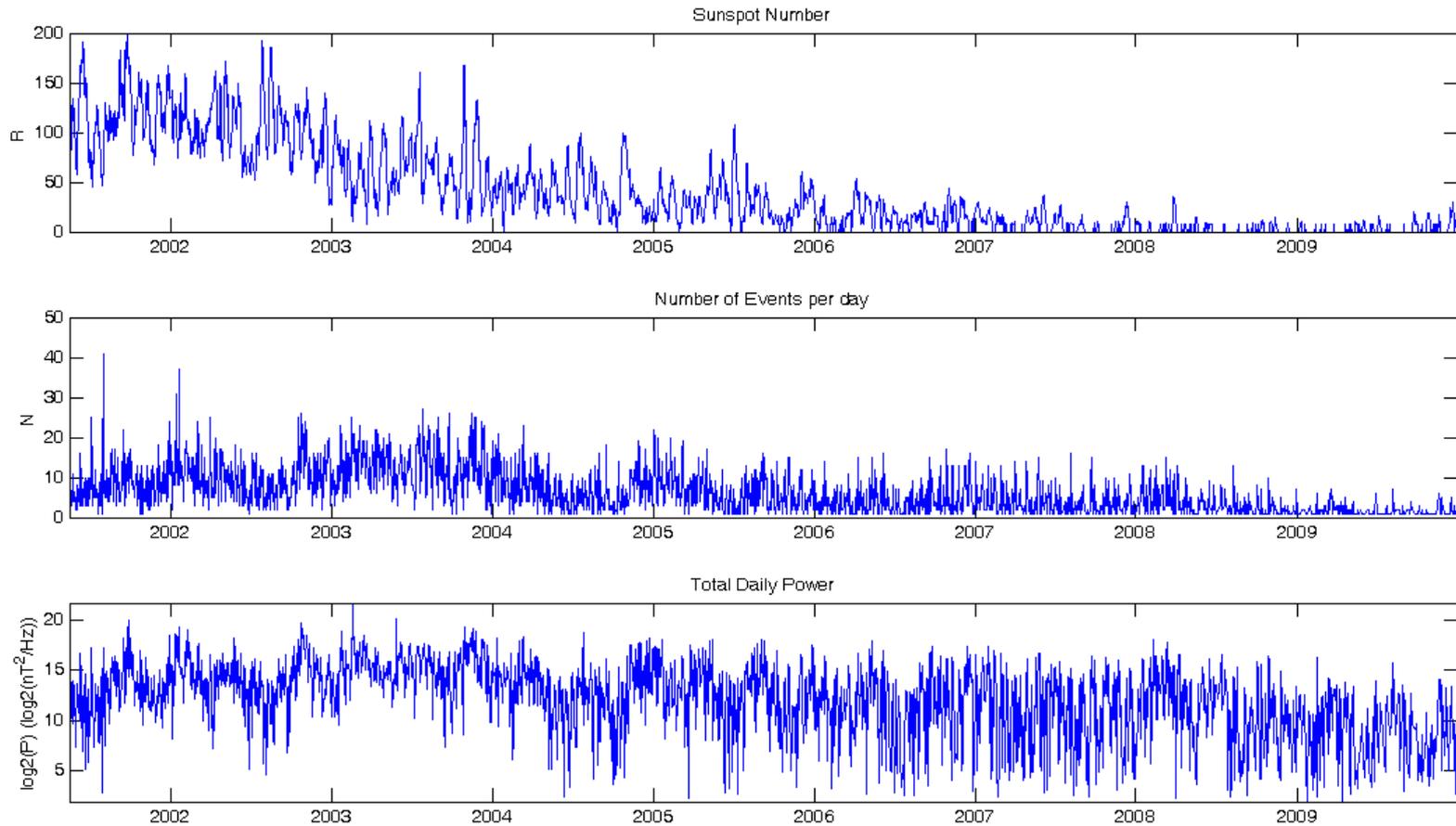
Dataset Description



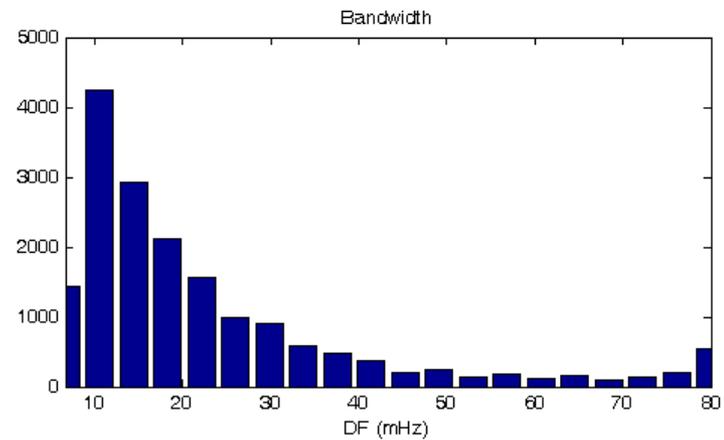
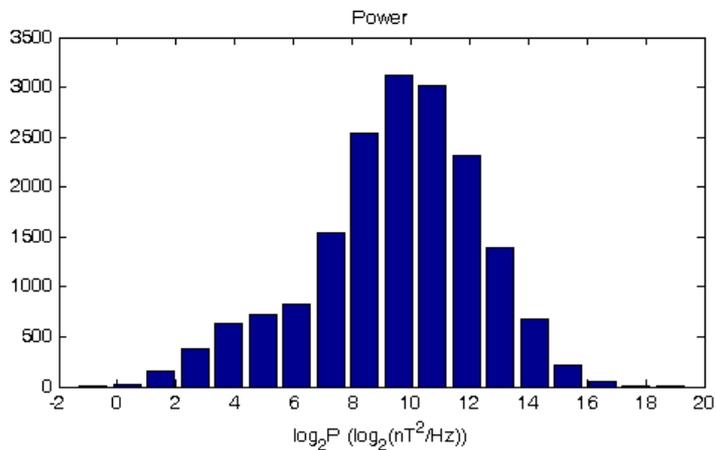
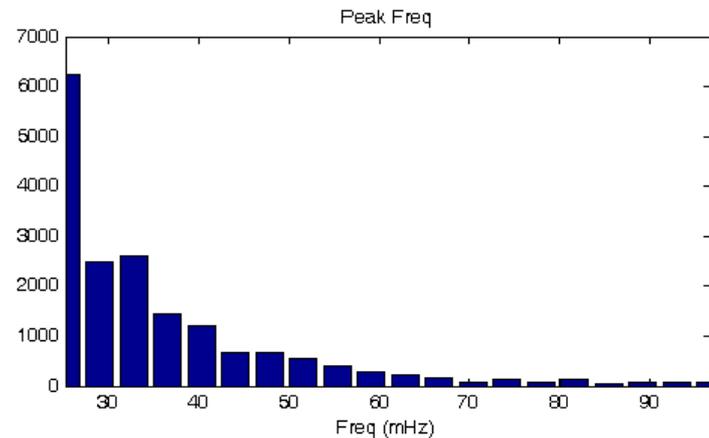
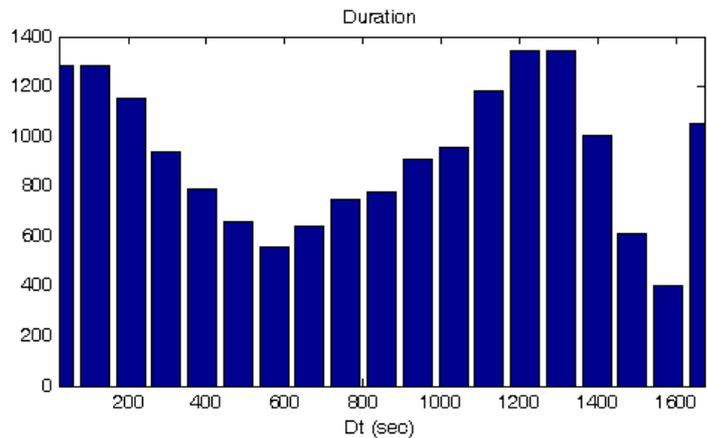
- 1 Hz magnetic field data (scalar)
- 15 sec ion density data
- Time span: May 2001 – Dec 2009
- > 90,000 Tracks
- Limited to [-55, +55] Mag. Lat.
- > 17,000 Wave Events Selected
- Additional Data: OmniWeb Service
<http://omniweb.gsfc.nasa.gov/>



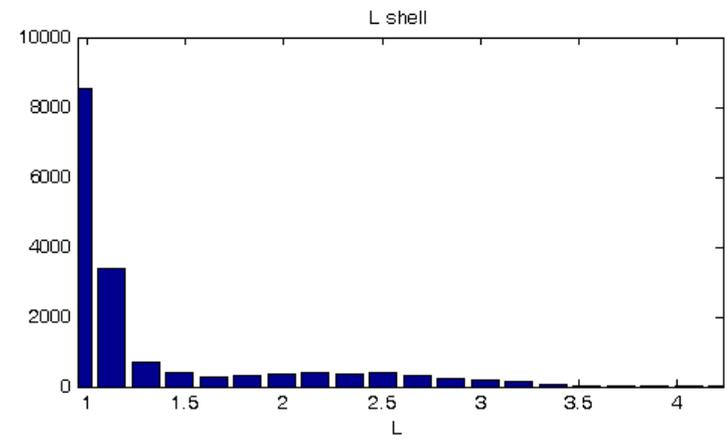
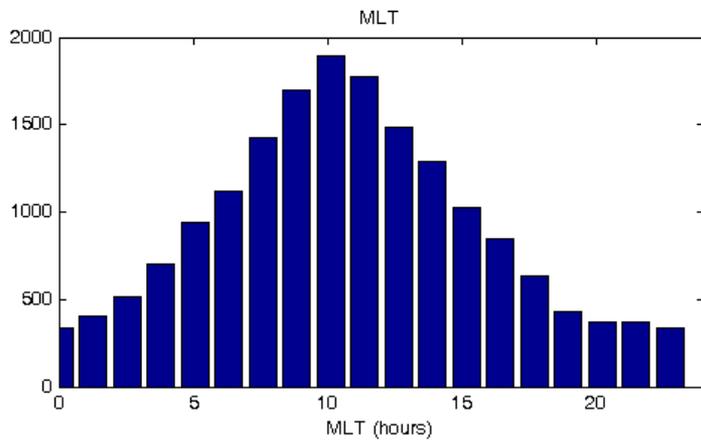
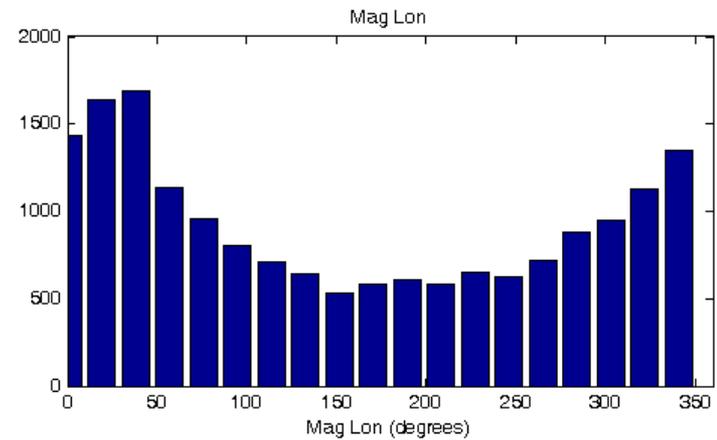
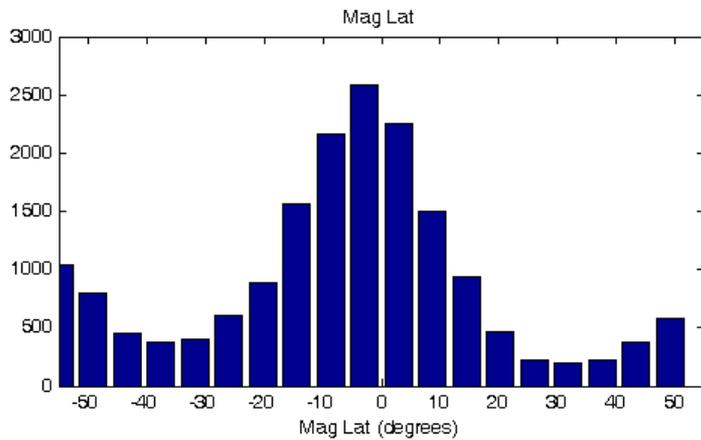
Overview (2001 - 2009)



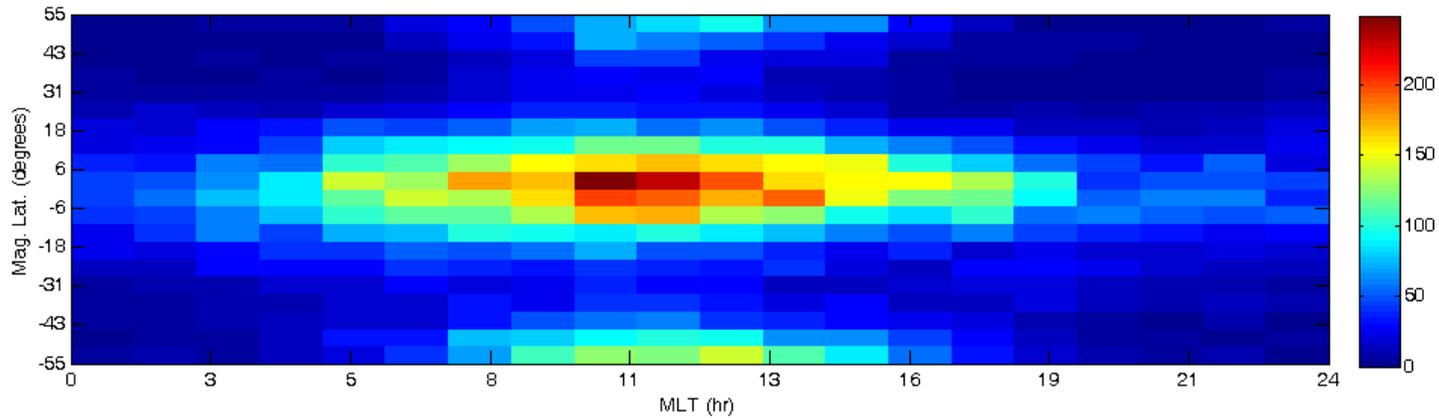
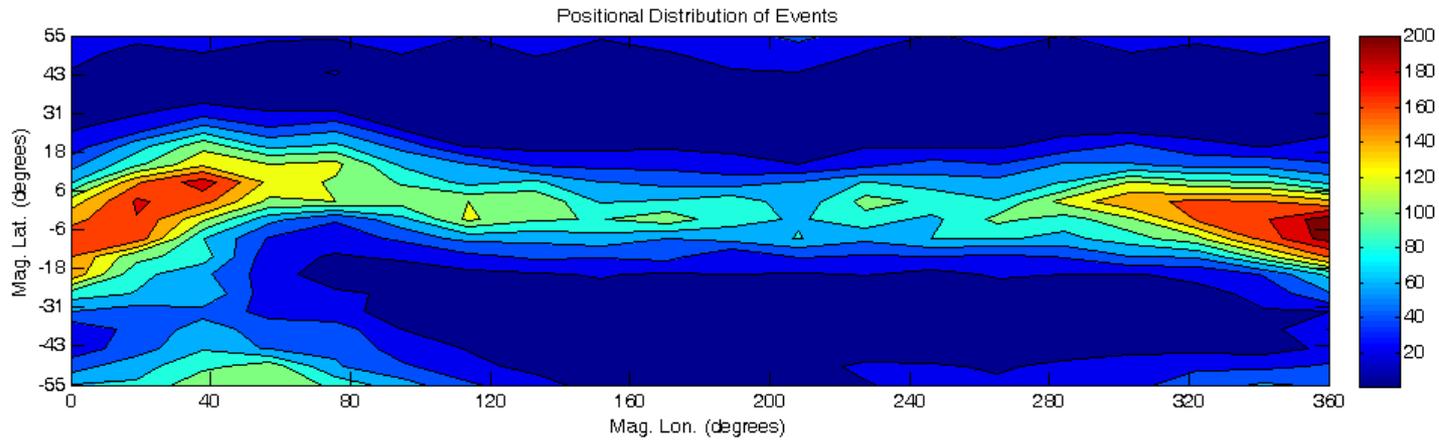
Statistical Distributions of Wave Parameters



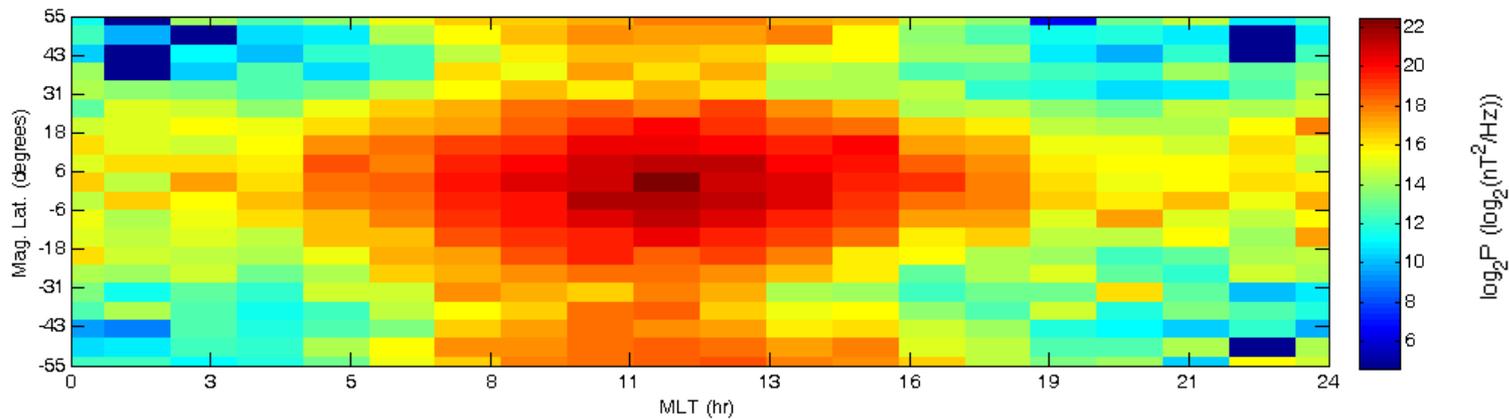
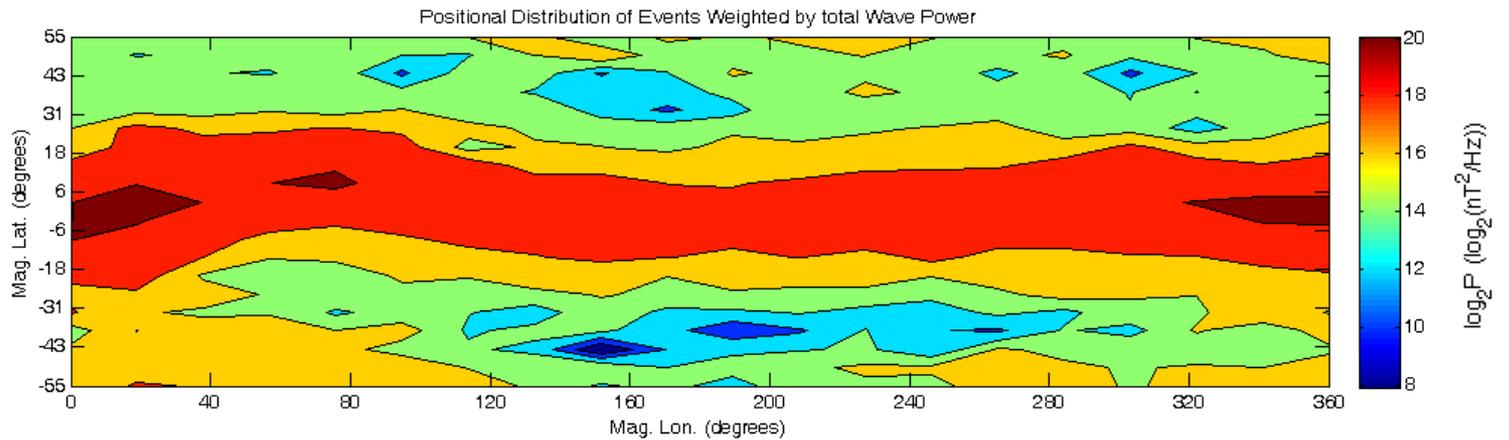
Statistical Distributions of Wave Parameters



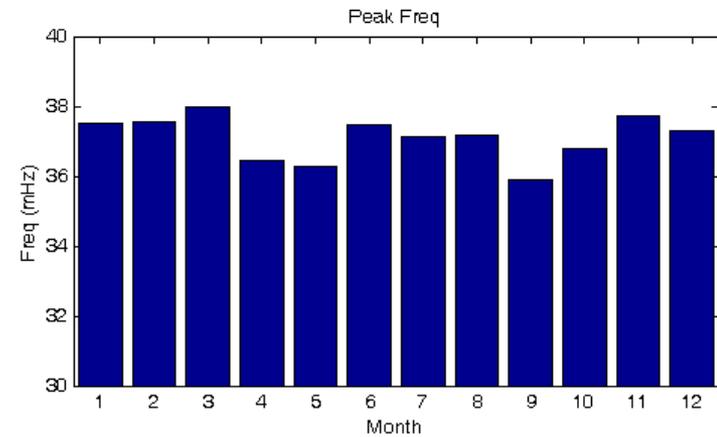
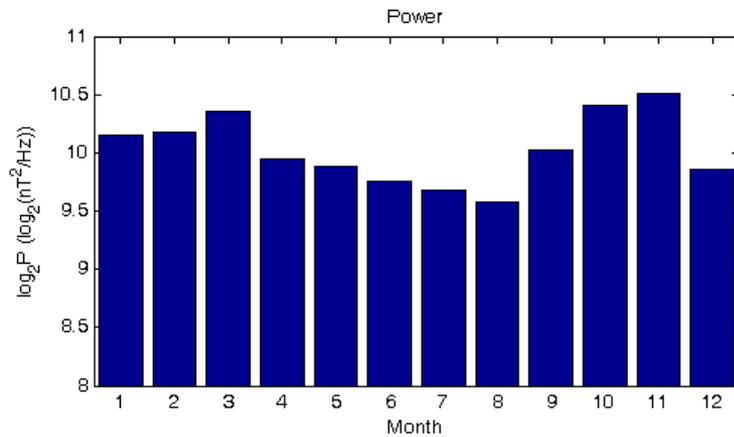
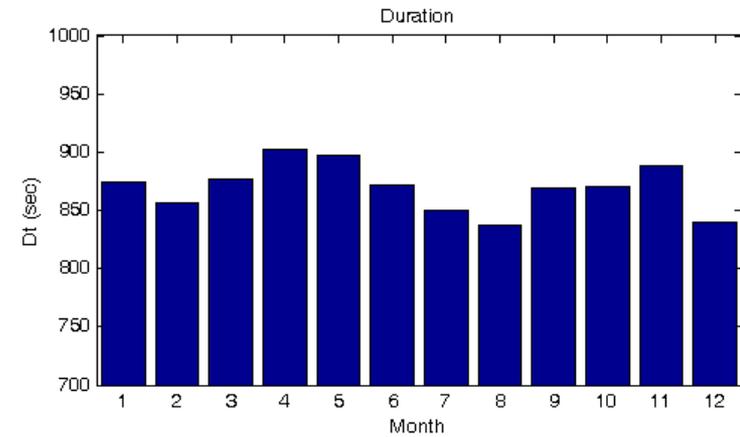
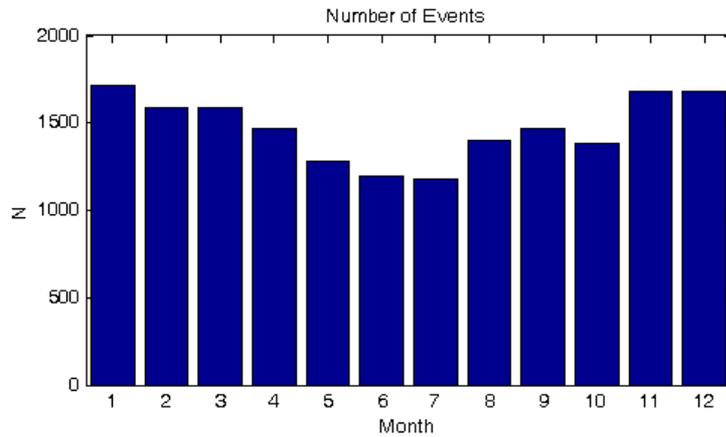
Positional Distribution



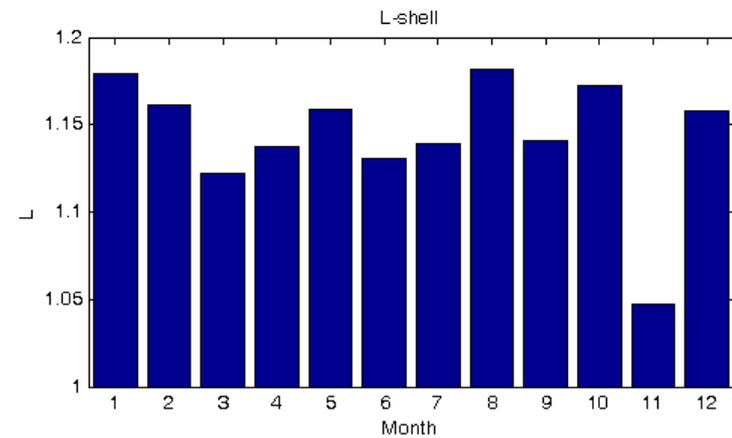
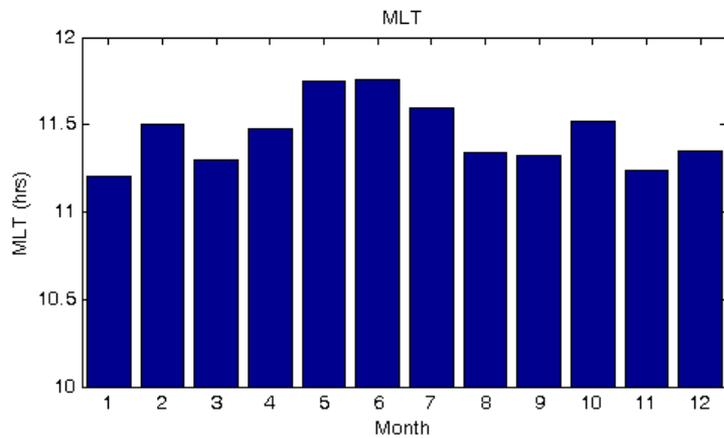
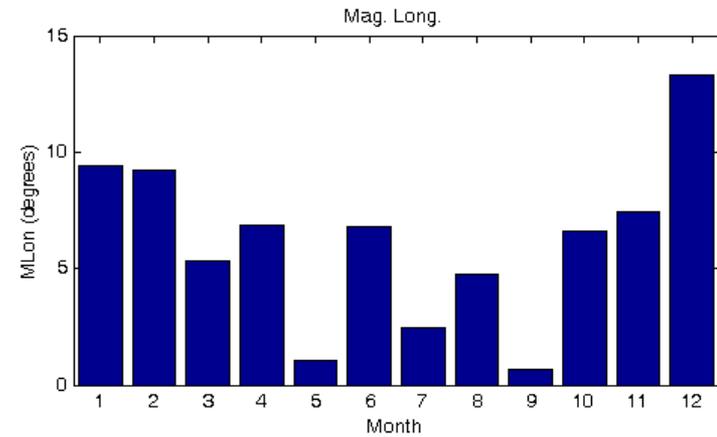
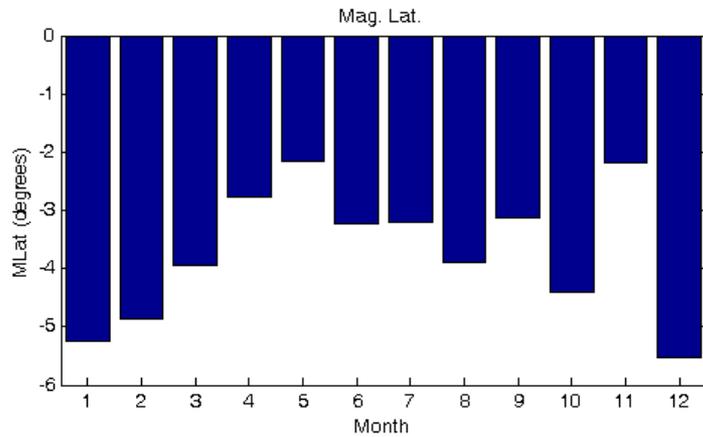
Weighted Distribution (log2)



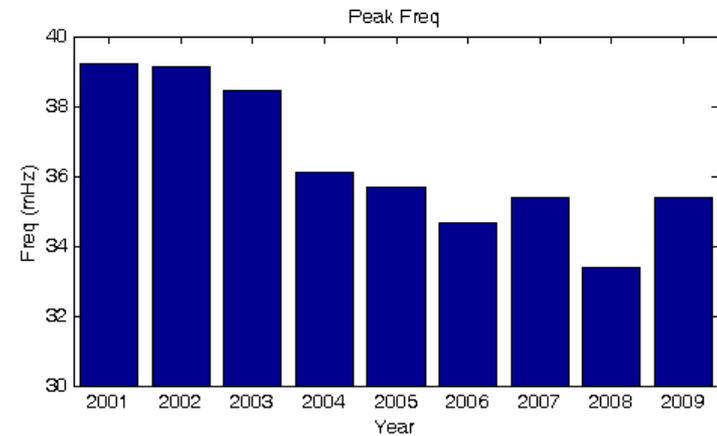
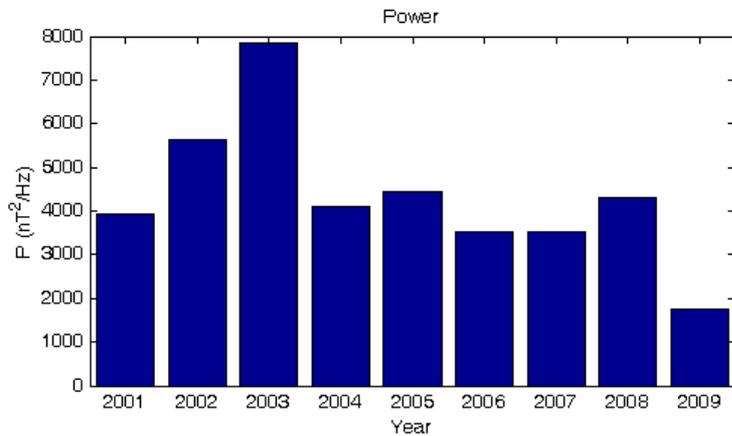
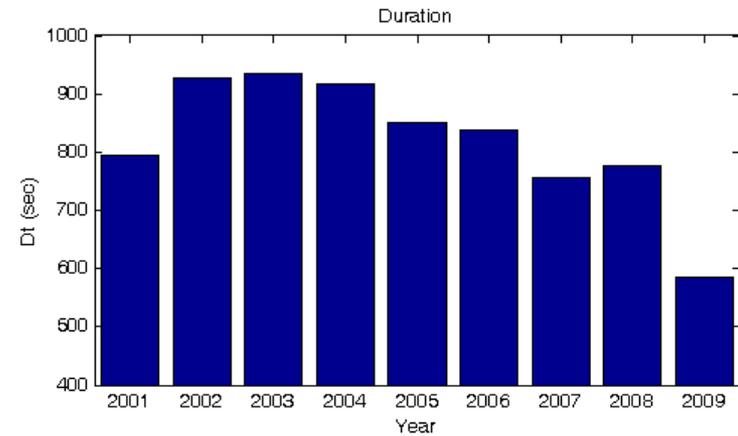
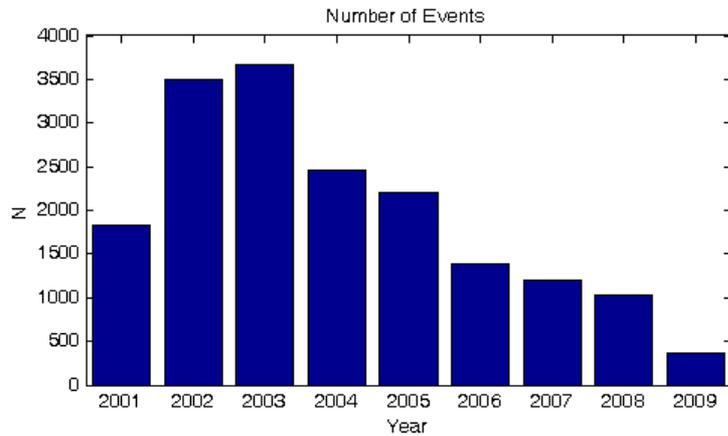
Seasonal Variations



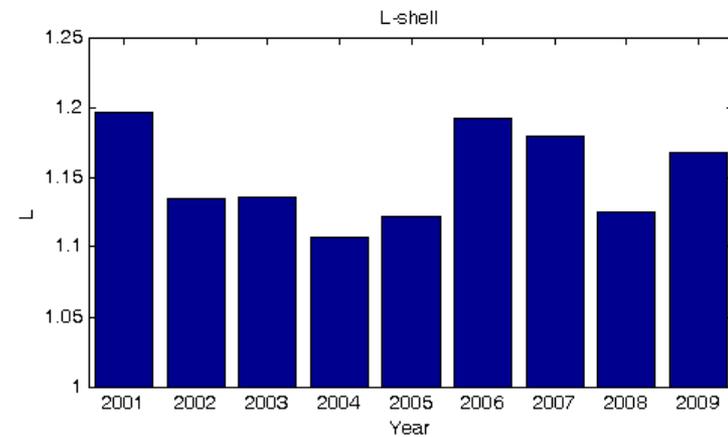
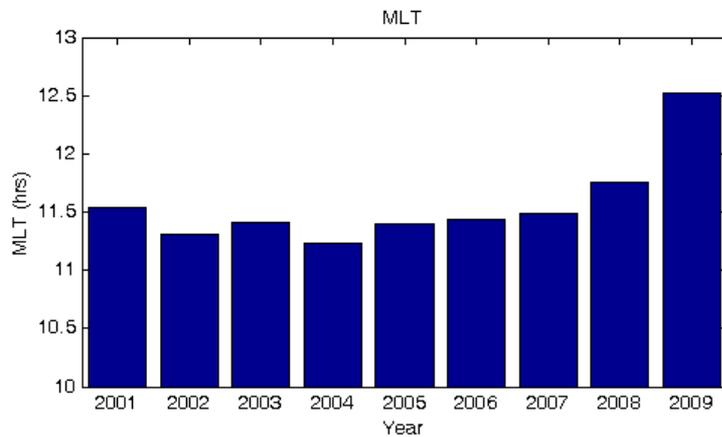
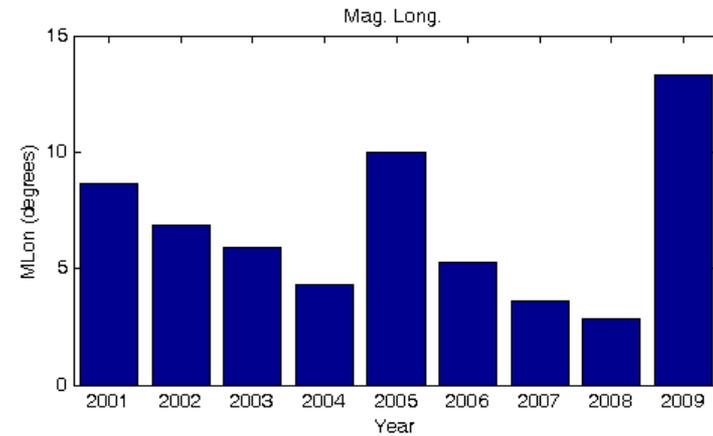
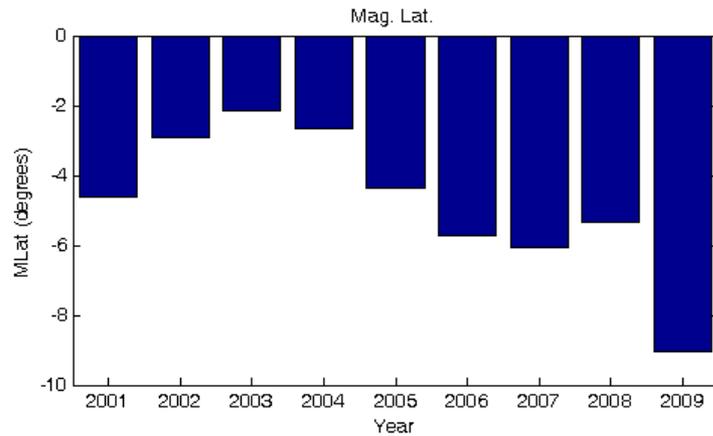
Seasonal Variations



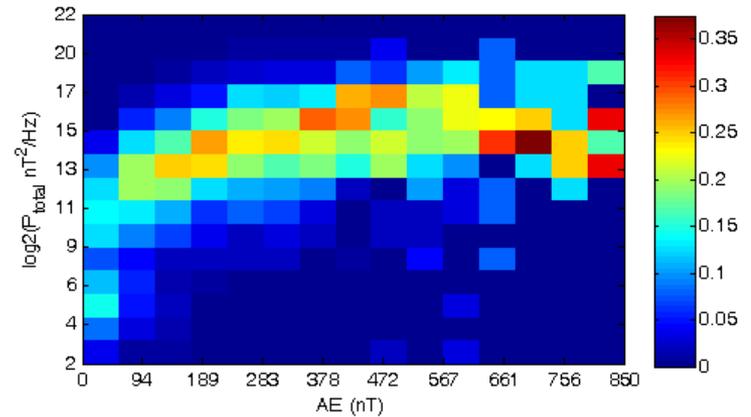
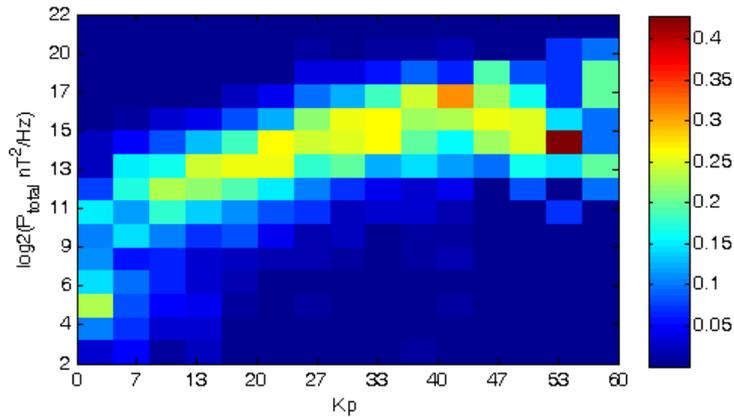
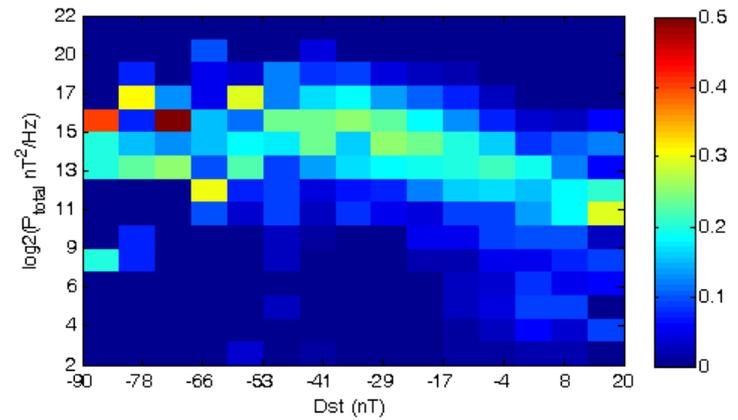
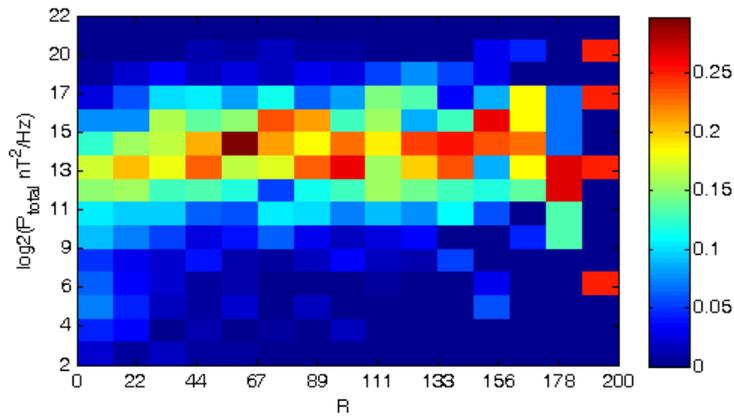
Yearly Variations



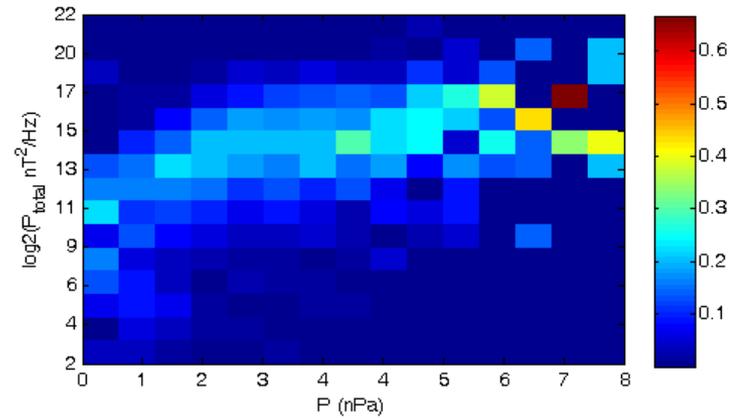
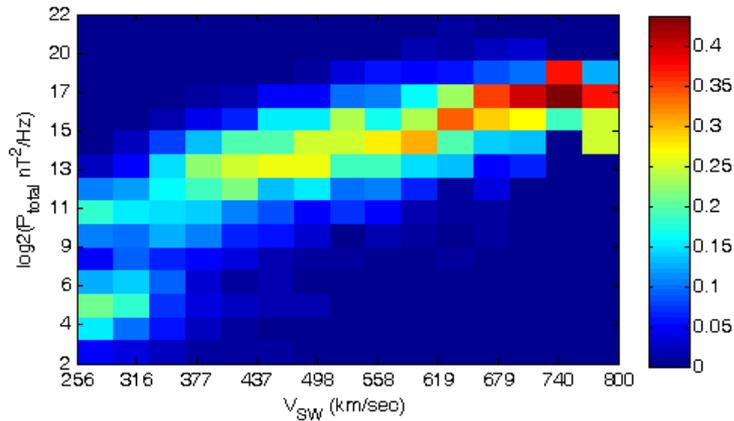
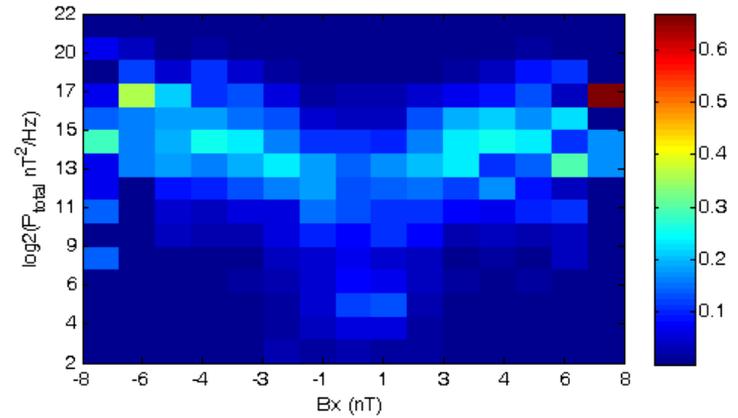
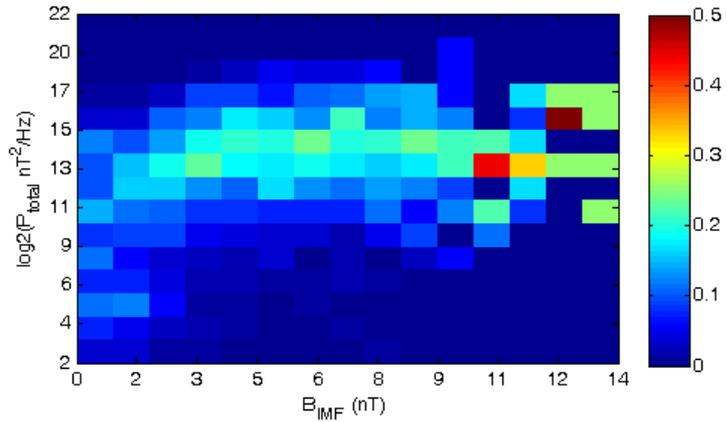
Yearly Variations



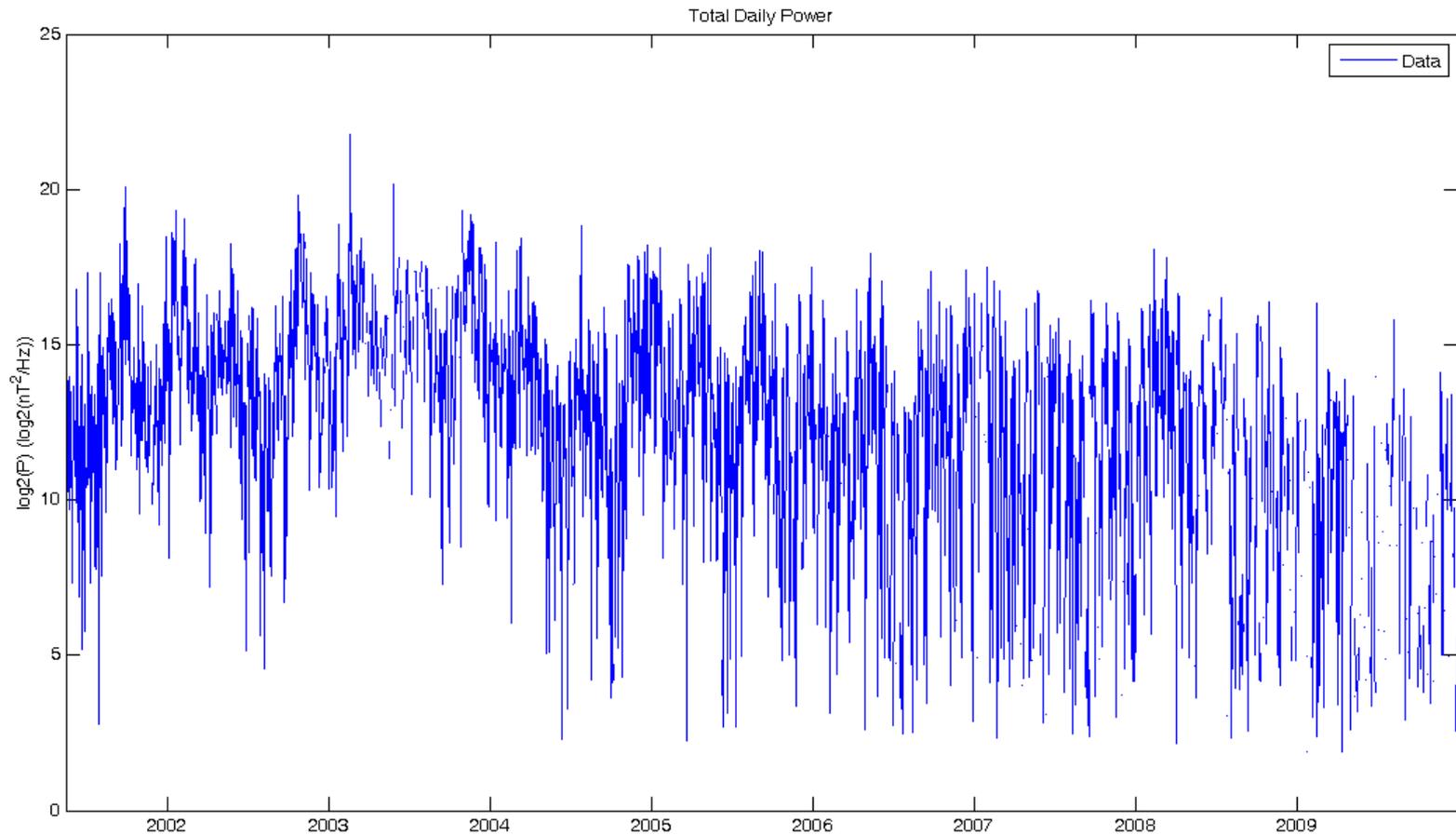
Correlations with Indices



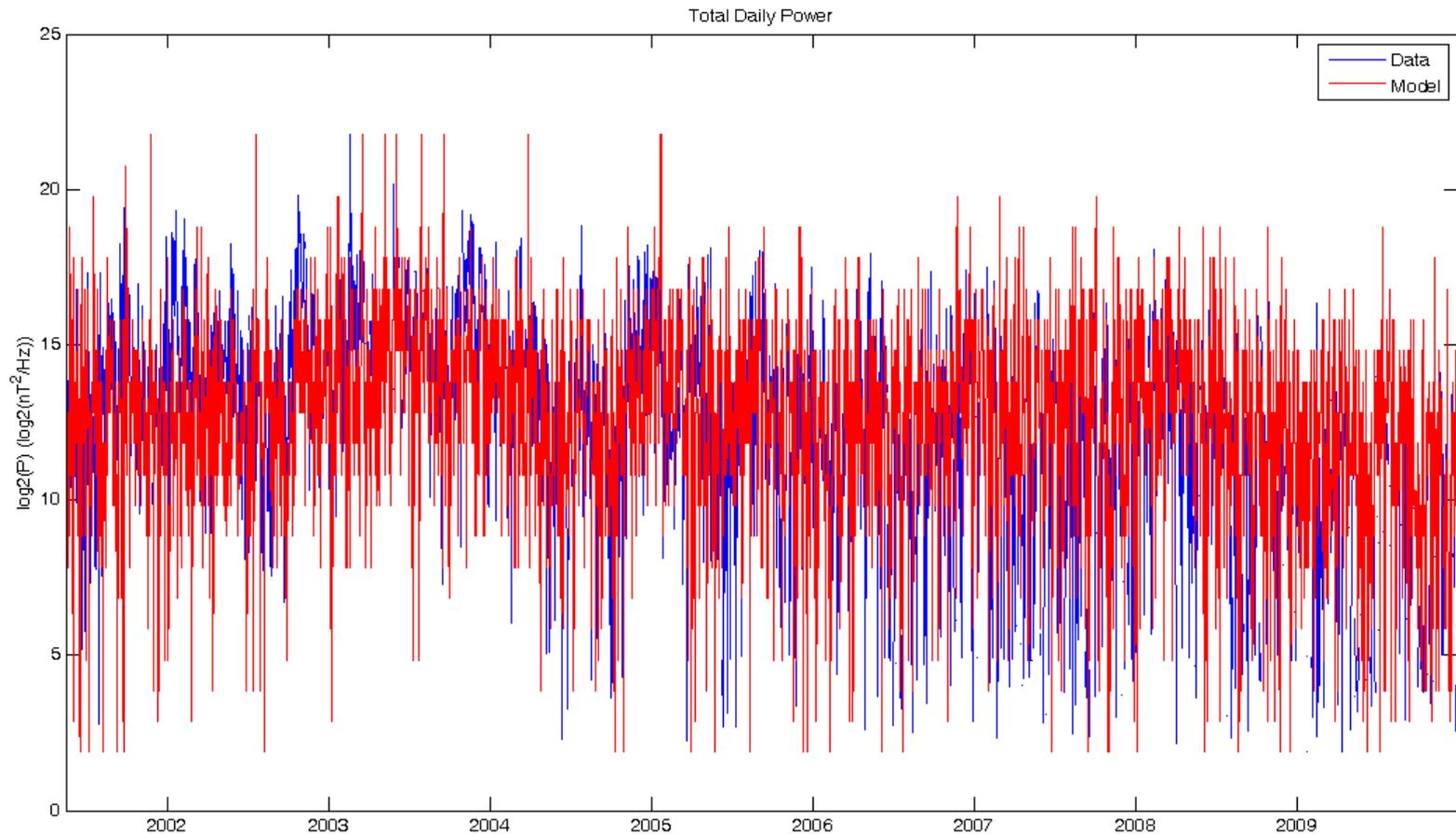
Correlations with SW parameters



V_{SW} -driven Probabilistic Modelling



V_{SW} -driven Probabilistic Modelling



Summary

- ✓ Automated Detection Methodology
 - ✓ Statistics of Wave Parameters
 - ✓ Variations of Wave Parameters vs Time
 - ✓ Correlations with Indices of Magnetospheric activity
 - ✓ Correlations with SW parameters
 - ✓ Modelling (?)
-
- ❑ Smaller time scales (~ hr)
 - ❑ Vector Data of the Magnetic Field
 - ❑ Propagation/Polarization parameters
 - ❑ New datasets (SWARM Mission)



