

Ground-based optical and ULF/ELF/VLF wave measurements at subauroral latitudes prepared for the ERG project

K. Shiokawa, C.-W. Jun, C. Martinez, N. Sunagawa, Y. Miyoshi

Solar-Terrestrial Environment Laboratory (STEL), Nagoya University, Japan

T. Nagatsuma, M. Ishii

National Institute of Information and Communications Technology (NICT), Japan

M. Ozaki

Kanazawa University, Japan

M. Connors, I. Schofield

Athabasca University, Canada

P. T. Jayachandran

University of New Brunswick, Canada

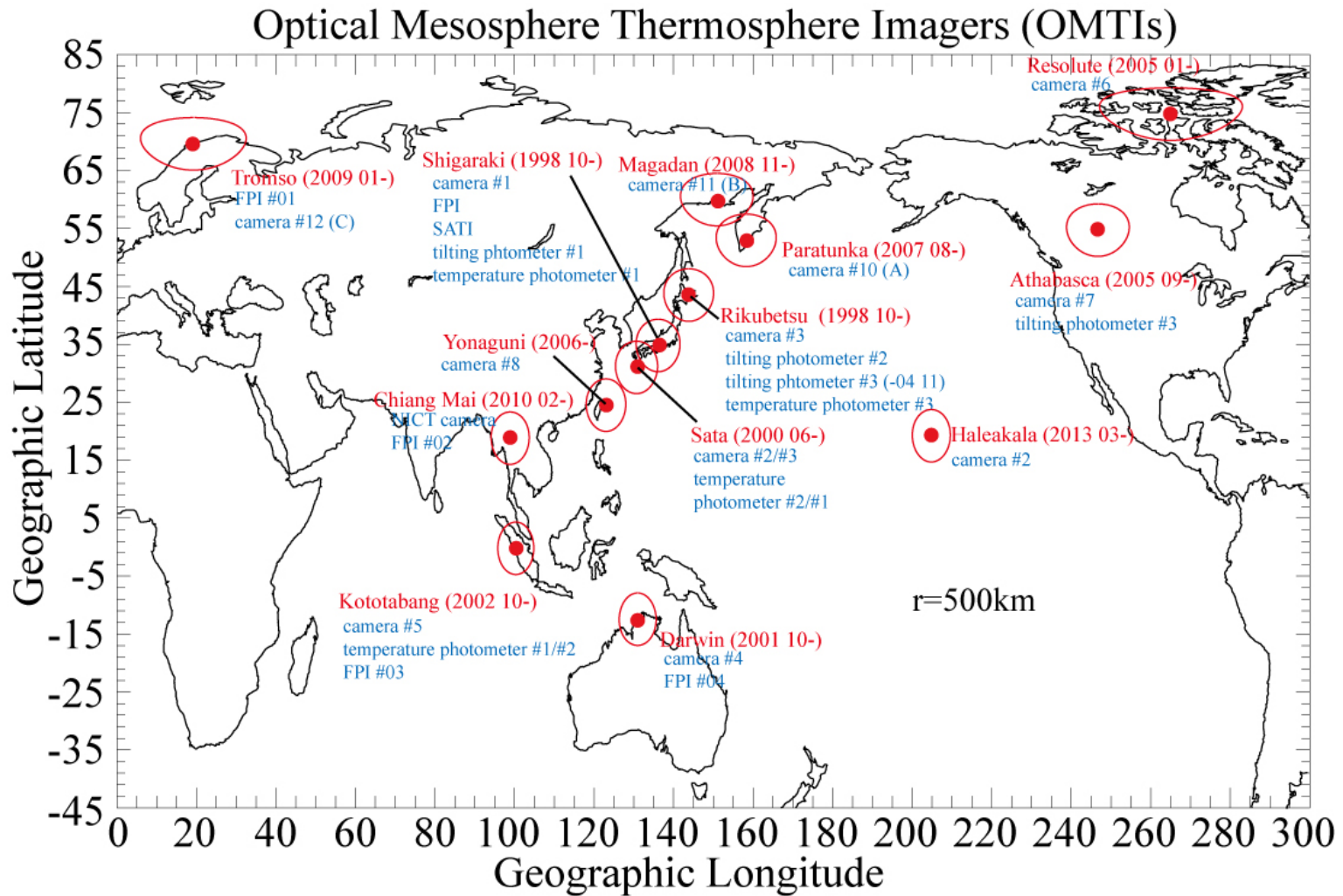
I. Poddelsky, B. Shevtsov

Institute of Cosmophysical Research and Radio Wave Propagation (IKIR), Russia

D. Baishev

Yu.G.Shafer Institute of Cosmophysical Research and Aeronomy (IKFIA), Russia

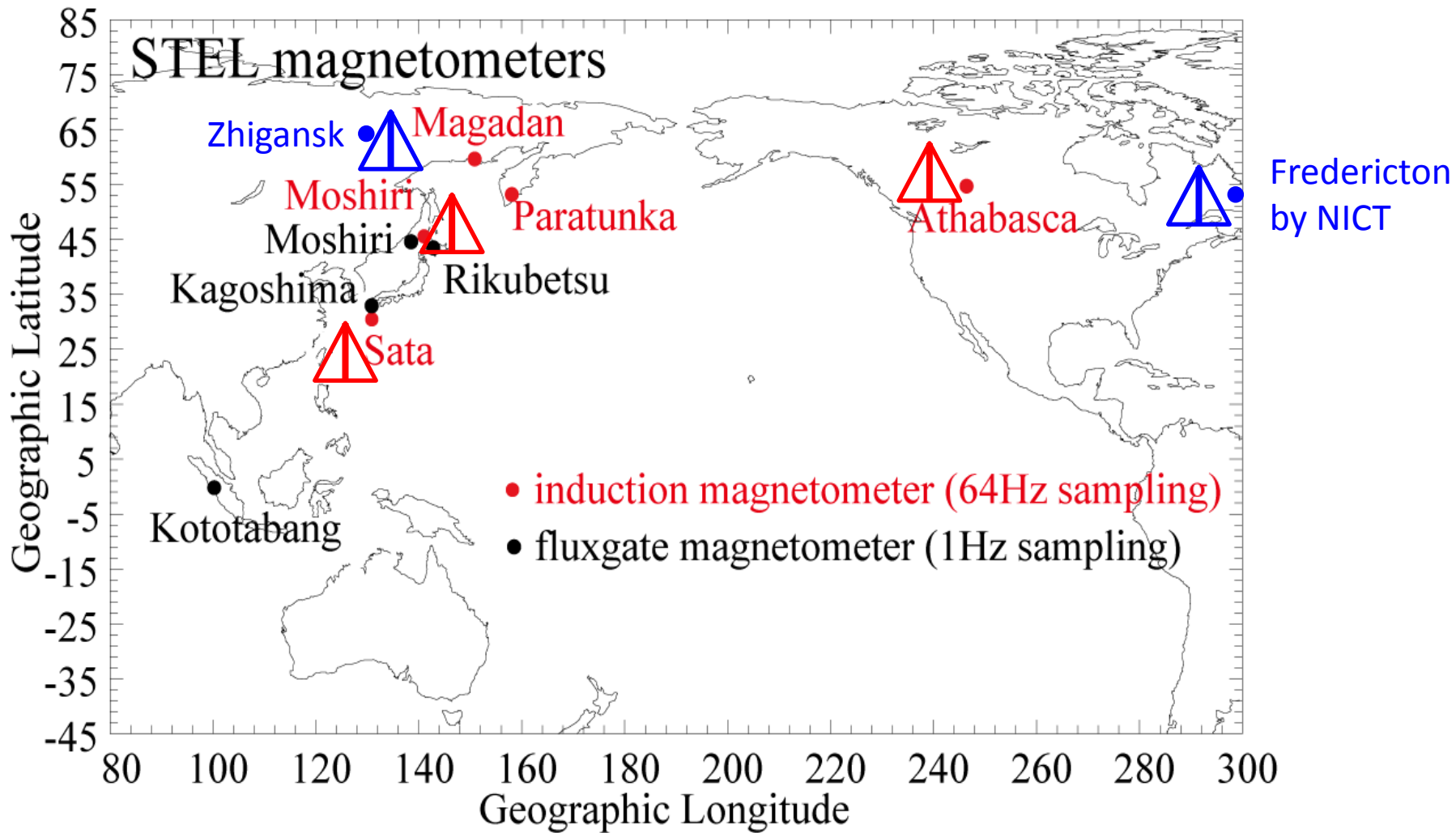
Optical Mesosphere Thermosphere Imagers (OMTIs)



OMTIs optical imager network operated by STEL, Nagoya University: airglow/aurora imagers, Fabry-Perot interferometers, meridian-scanning photometers, and airglow temperature photometers

Contact: Kazuo Shiokawa (PI, [shiokawa at stelab.nagoya-u.ac.jp](mailto:shiokawa@stelab.nagoya-u.ac.jp))

Web page: <http://stdb2.stelab.nagoya-u.ac.jp/omti/index.html>



STEL magnetometer network operated by STEL, Nagoya University: fluxgate and induction magnetometers

Contact: Kazuo Shiokawa (PI, [shiokawa at stelab.nagoya-u.ac.jp](mailto:shiokawa@stelab.nagoya-u.ac.jp))

Web page: <http://stdb2.stelab.nagoya-u.ac.jp/magne/index.html>

CDF database construction usable by TDAS is underway.

Routine measurements at Athabasca (L=4.4), Canada

by Athabasca Univ., STEL, Nagoya Univ., Kanazawa Univ., and Tohoku Univ.

- * **all-sky airglow imager**: 5577, 6300, 4861, 5893, 8446
time resolution 1.5-10 min (STEL)
- * **meridian-scanning photometer**: 5 points in the sky, 4861, 4278 (STEL)
- * **induction magnetometer**: 64Hz sampling, peak at 5Hz (STEL)
- * **all-sky EMCCD camera**: 5577, 6300, BG3,
time resolution: a few sec (Athabasca Univ.)
- * **LF standard radio wave receiver** (Tohoku Univ.)
- * **2ch VLF Antenna** 100kHz sampling (STEL/Kanazawa Univ.)
- * **1ch riometer** (STEL)
- * **EMCCD Camera**: BG3, ~100Hz sampling (STEL): from 2014 winter



Contents

(1) Aurora/Airglow Imagers

MSTID motion → monitoring of penetrating electric field
auroral fragmentation into patches → ballooning/interchange?
SAR arc → plasmasphere-ring current interaction

(2) Induction magnetometers

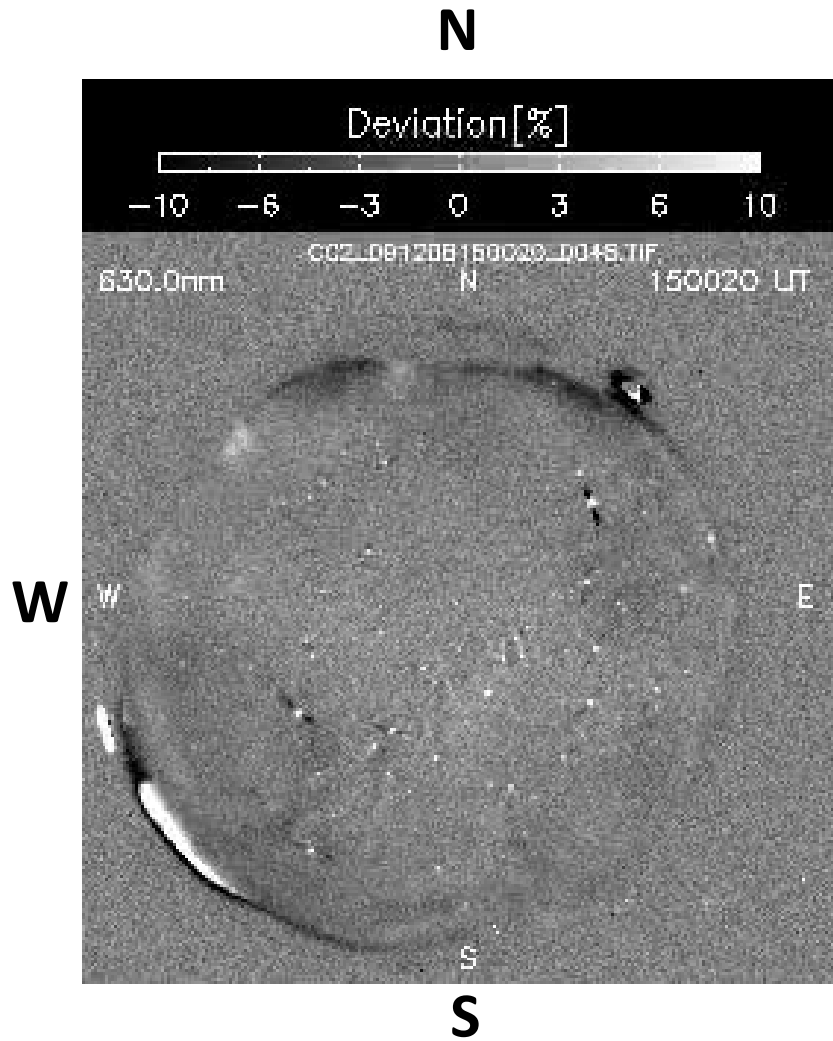
isolated proton aurora → monitoring of wave-particle interaction
Pc1 polarization characteristics (dependence on frequency and distance from the source)
Pc1 pearl structure (amplitude modulation of EMIC waves)
→ Ionospheric beating versus magnetospheric processes

(3) ELF/VLF receivers

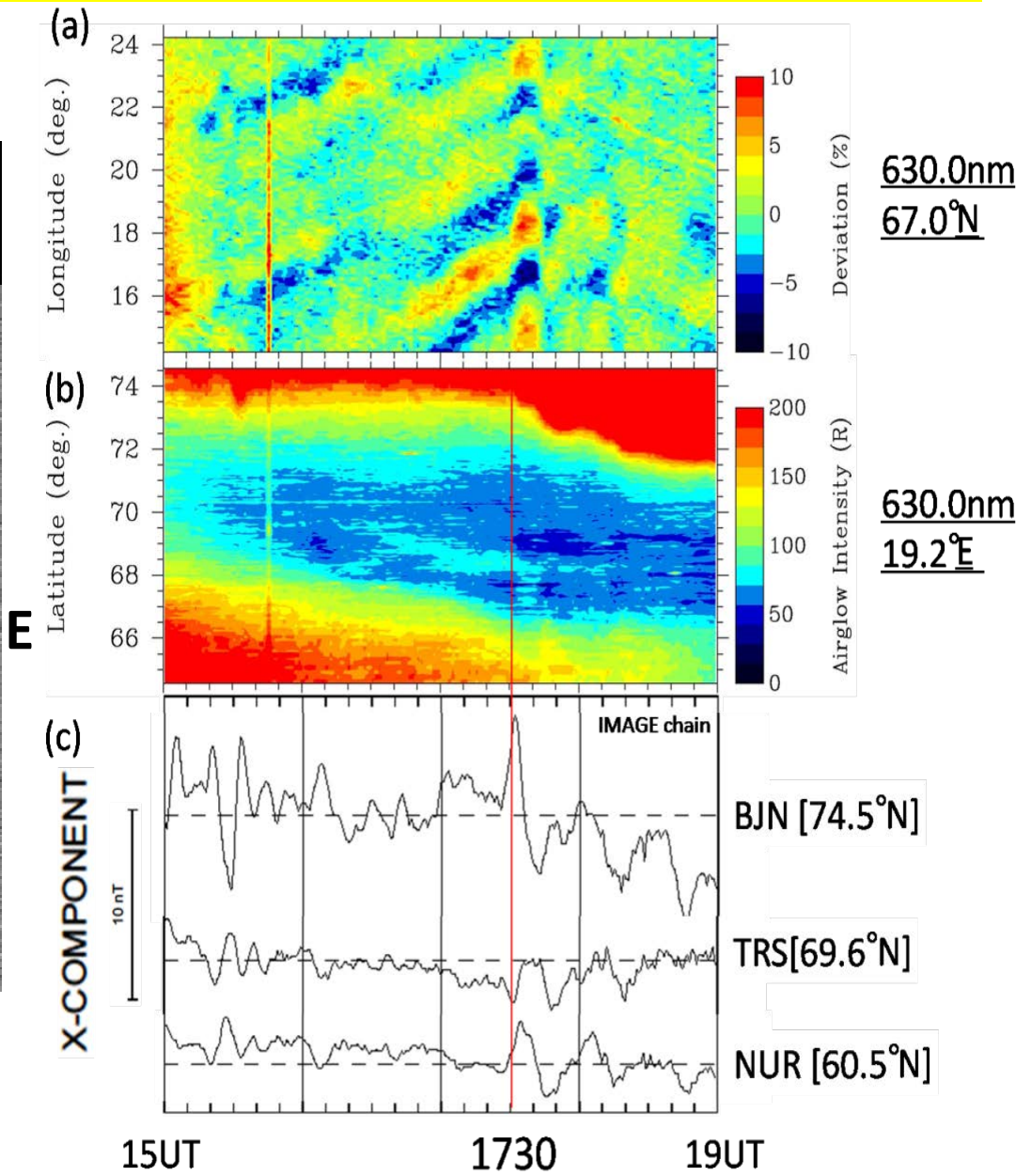
temporal variation of **QP** emission and **falling tone** emission
positive correlation of **frequency sweep rate and intensity** for QP
correlation between **EMIC and chorus** wave intensity
bursty patch emission and **ssc-triggered** emission
frequency and temporal dependence of **chorus polarization**
correlation with pulsating aurora – **30-40s switching** of interaction

Tromsø, Dec.8, 2009
15-19 UT (16-20 LT)

Penetration of auroral E-field (MSTID motion)



Shiokawa et al. (JGR, 2012)

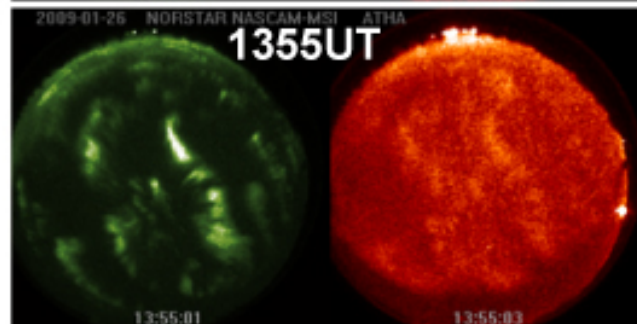
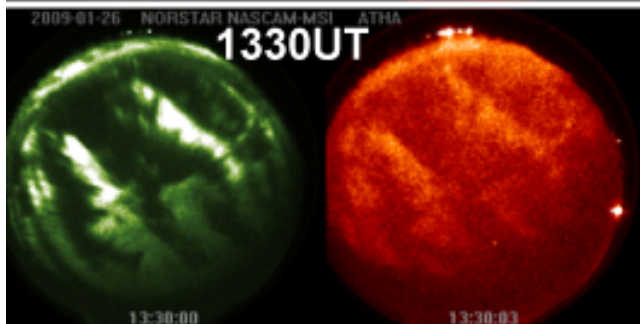
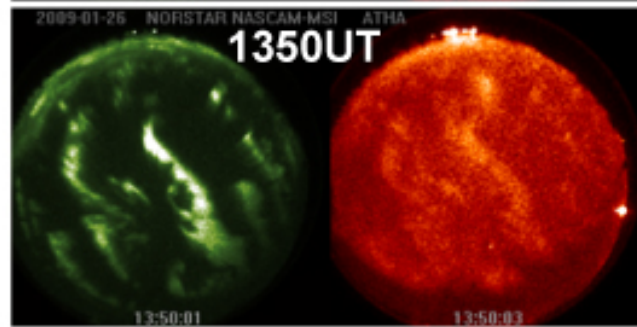
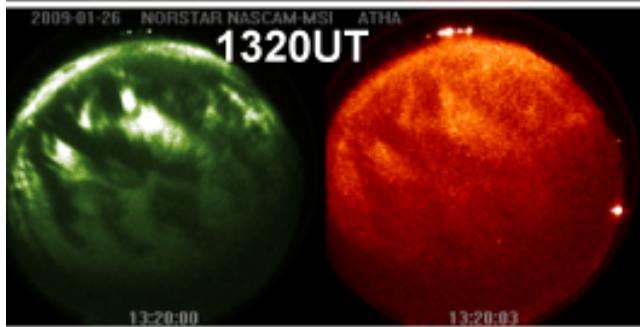
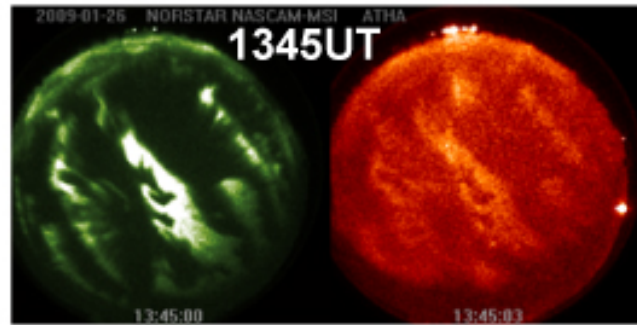
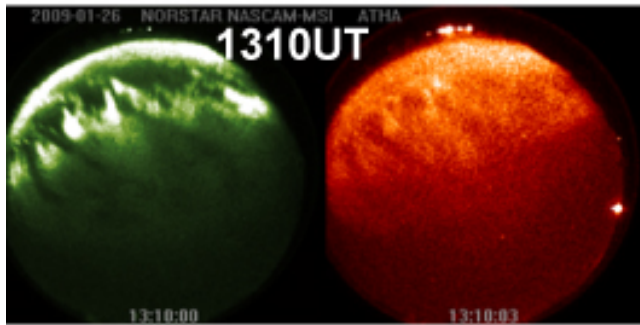
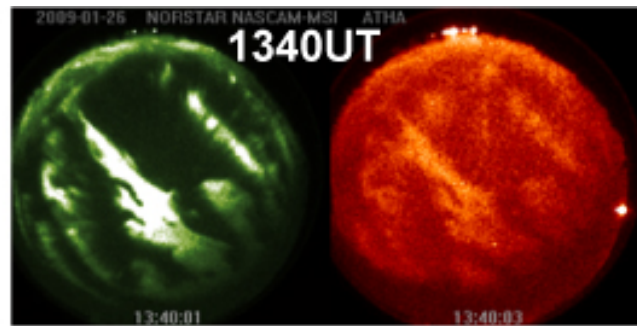
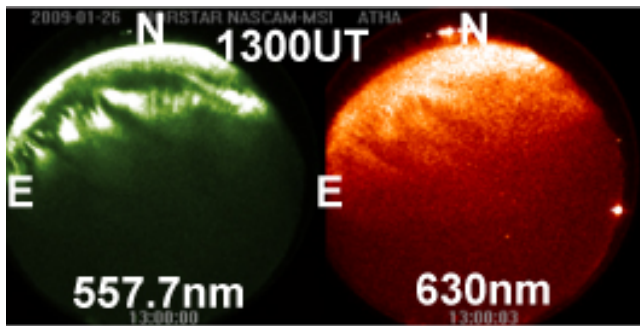


Auroral fragmentation into patches

equatorward
expansion speed
~150 m/s
at 1300-1330UT

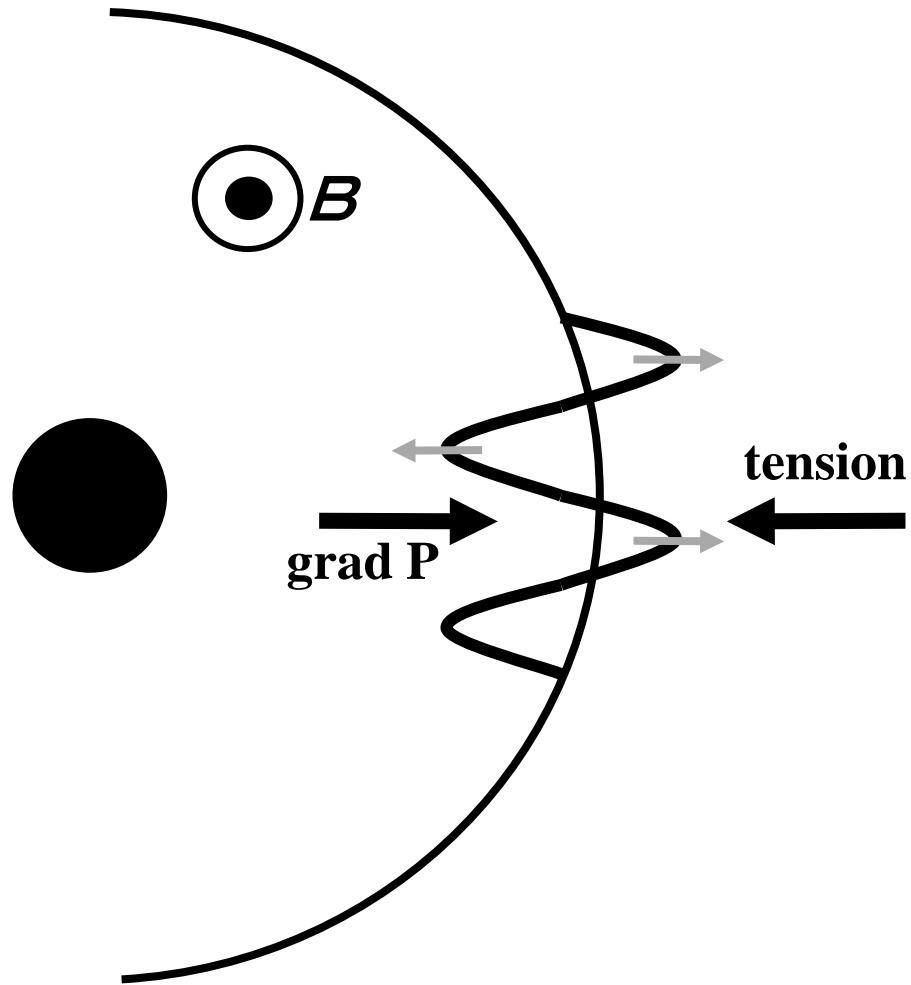
east-west scale size
~50-100km

azimuthal m
number ~180-360



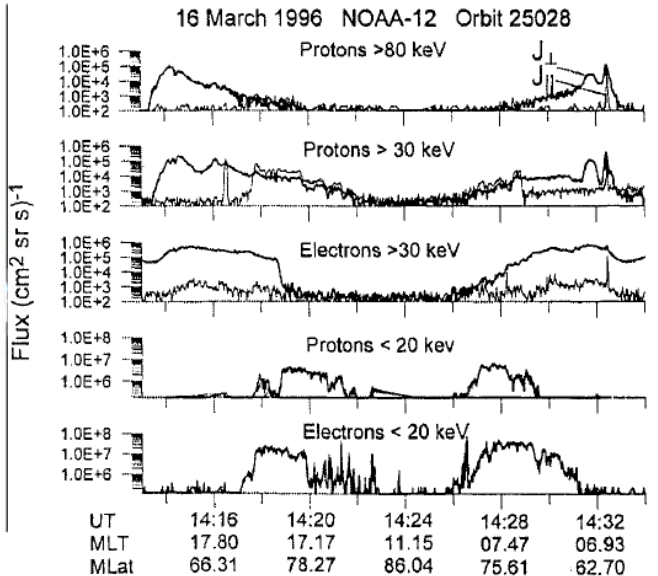
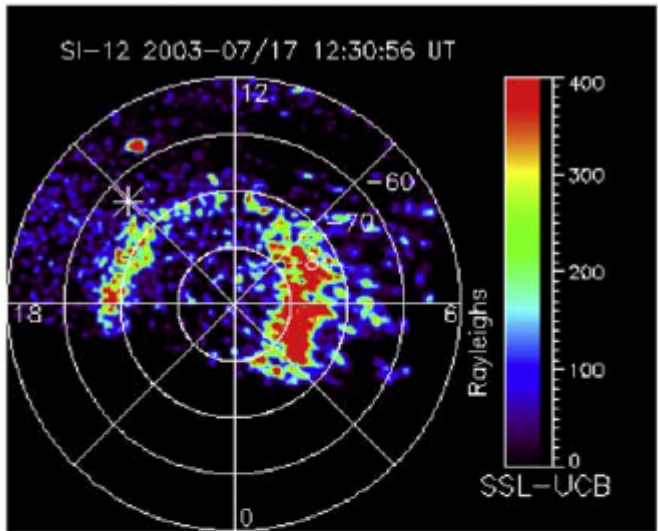
Shiokawa et al.
(JGR, in press, 2014)

Auroral fragmentation into patches



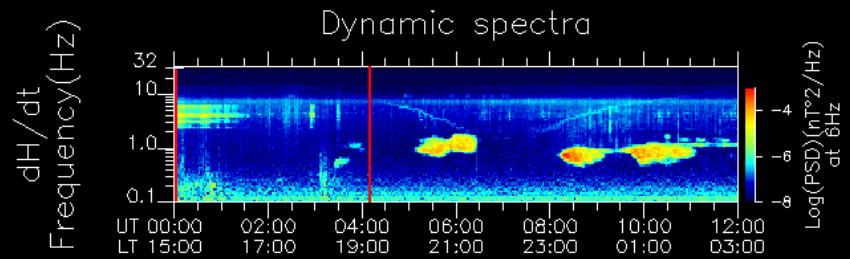
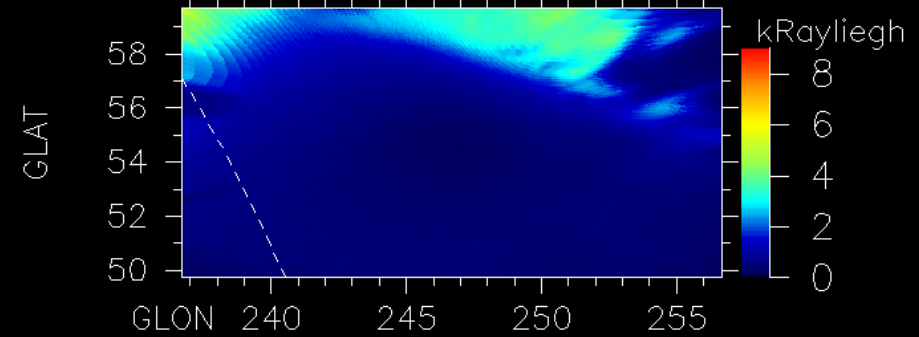
Interchange/ballooning instability

Isolated proton aurora and Pc1/EMIC waves

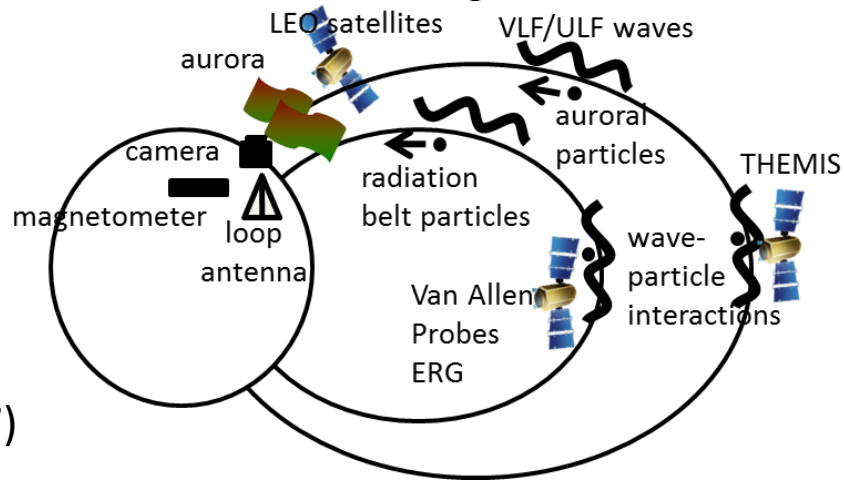


Yahnin et al. (GRL, 2000; JGR, 2007)

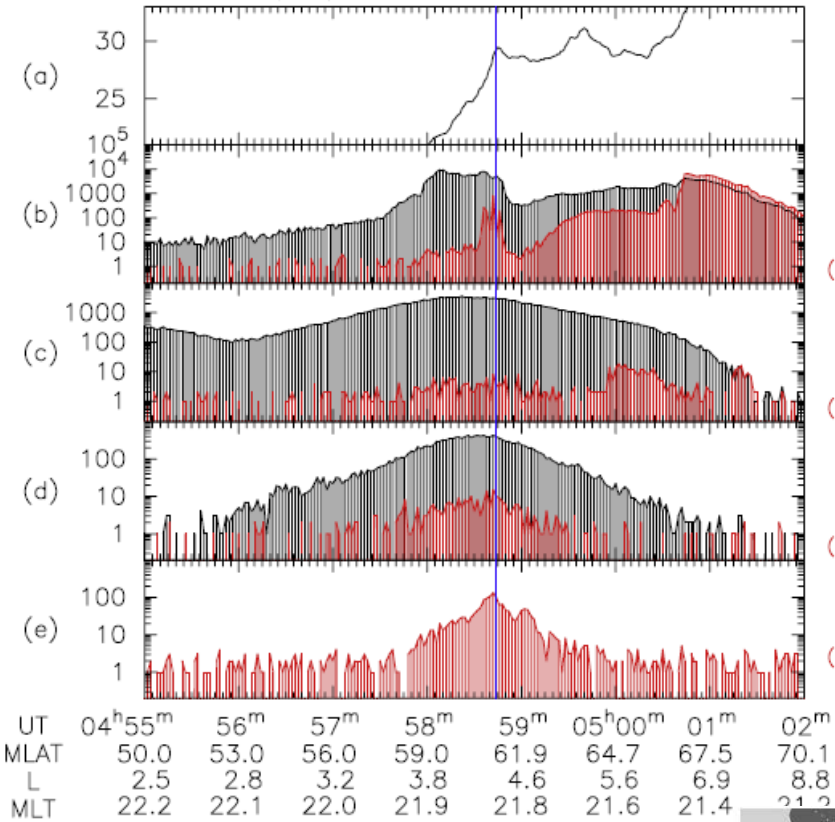
18 April 2006 04:10:00 UT
All-sky Image wavelength 557.7 nm at Athabasca



Sakaguchi et al. [JGR, 2007; 2008]



Isolated proton aurora and Pc1/EMIC waves



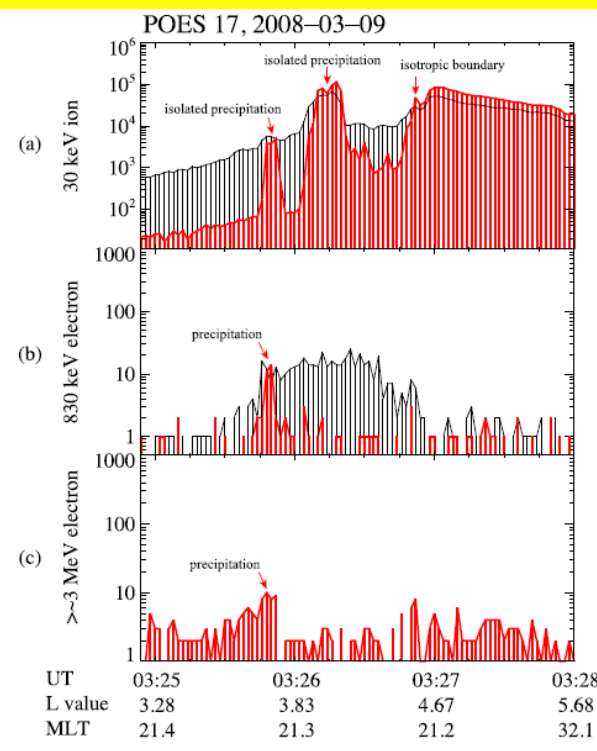
Hbeta
(Rayleigh)

30–80 keV ions
(trapped)
(precipitation)

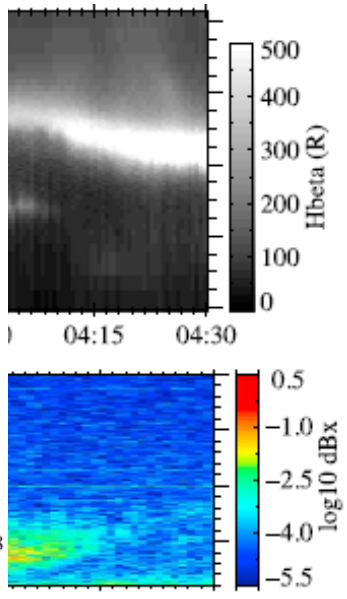
>300 keV electrons
(trapped)
(precipitation)

>800 keV electrons
(trapped)
(precipitation)

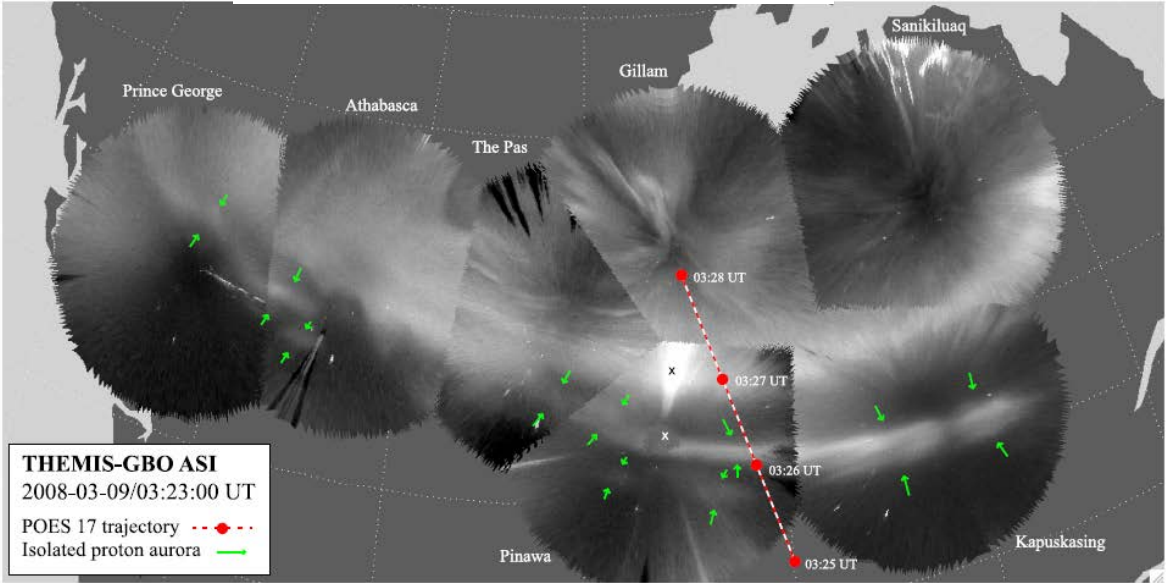
>3 MeV electrons
(precipitation)

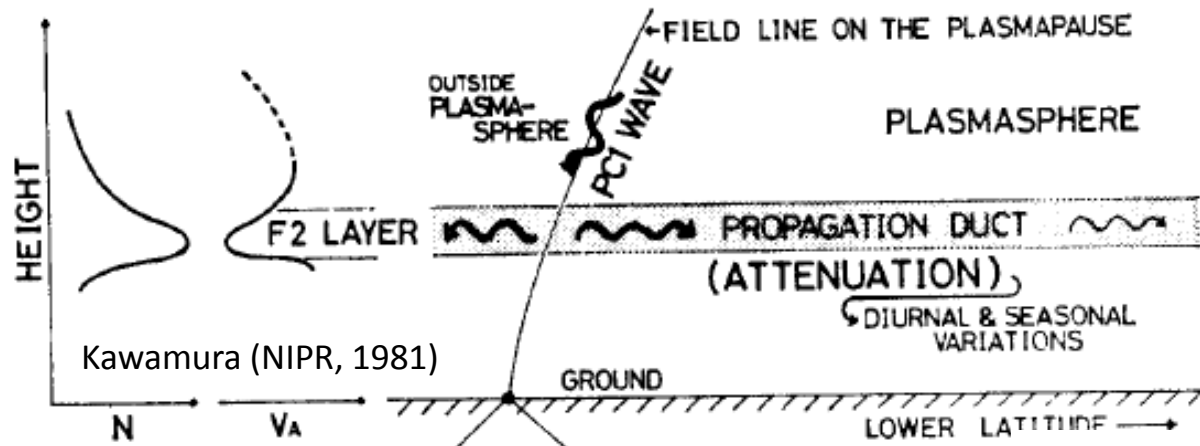


Sakaguchi et al.
(JGR, 2012)



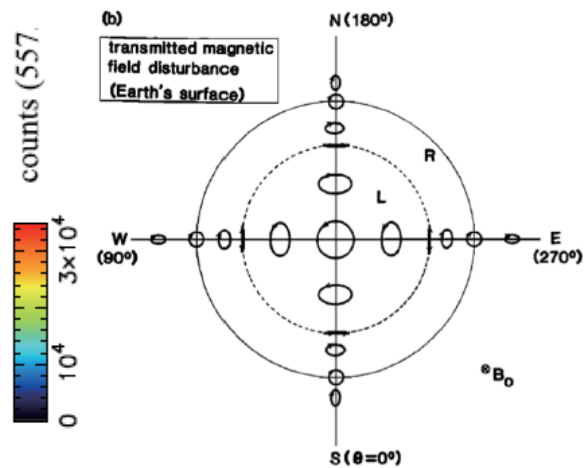
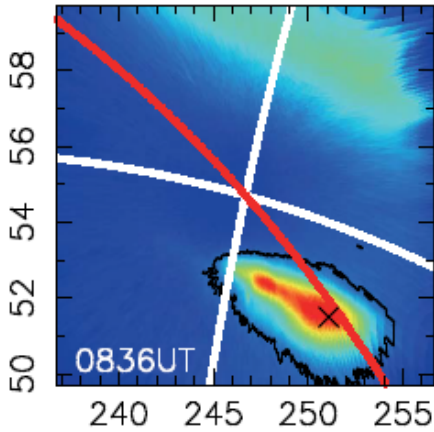
Miyoshi et al. (GRL, 2008)



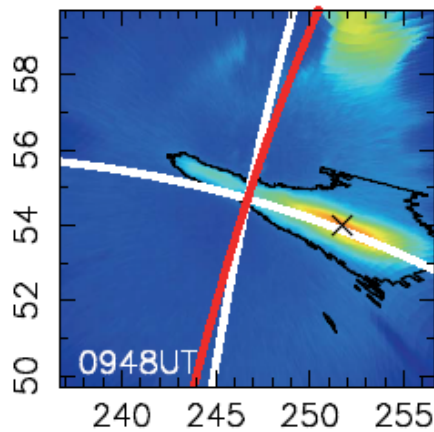


Isolated proton aurora and Pc1/EMIC wave polarization

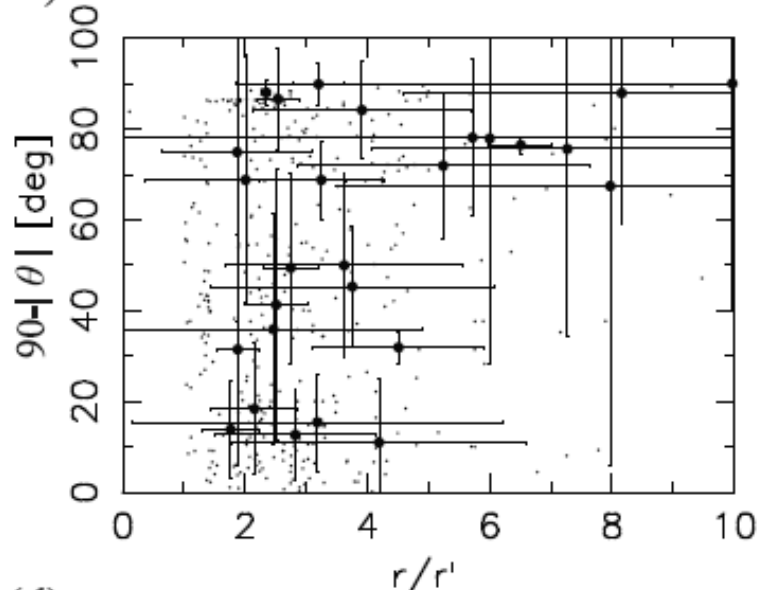
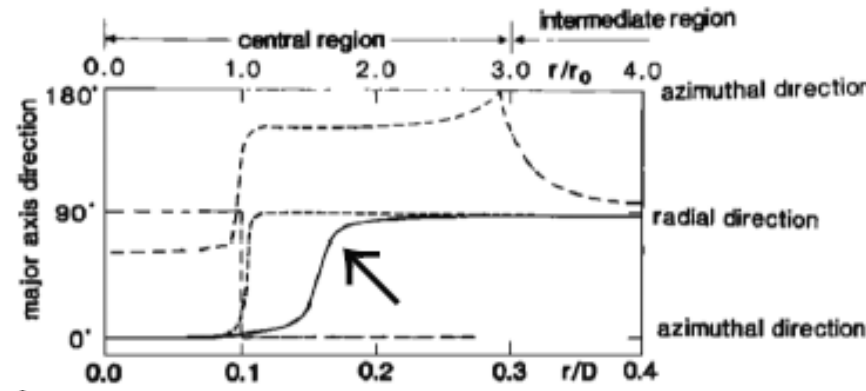
Athabasca



Fujita and Tamao (1988)

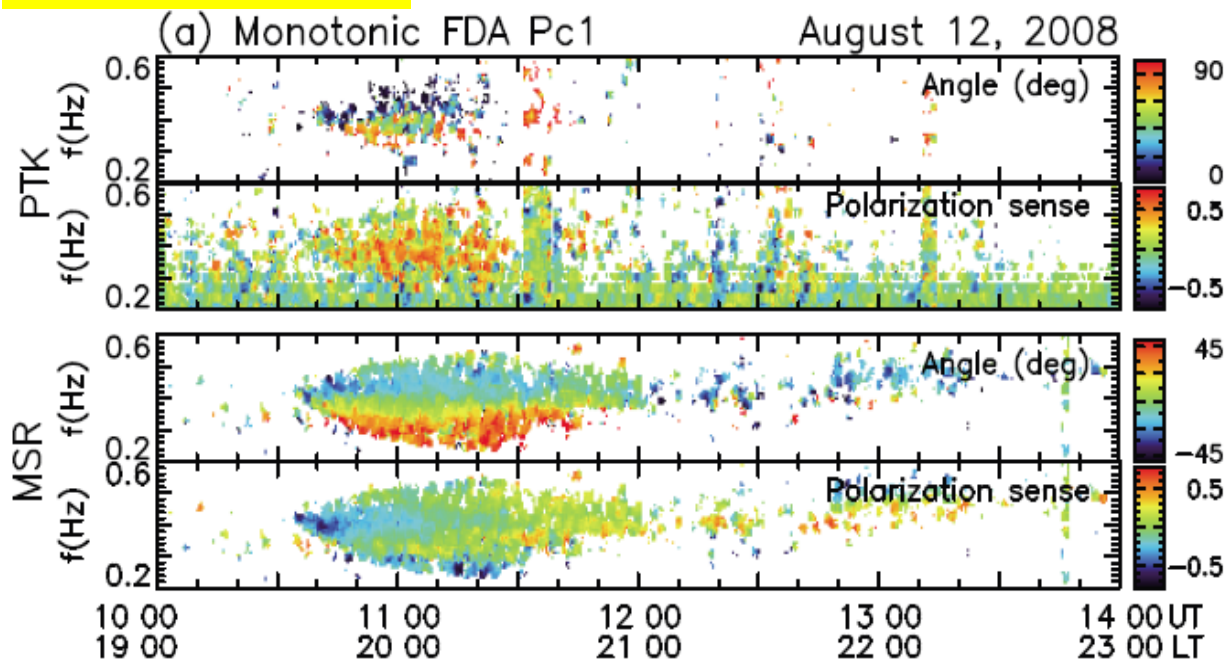


Nomura et al. (JGR, 2012)

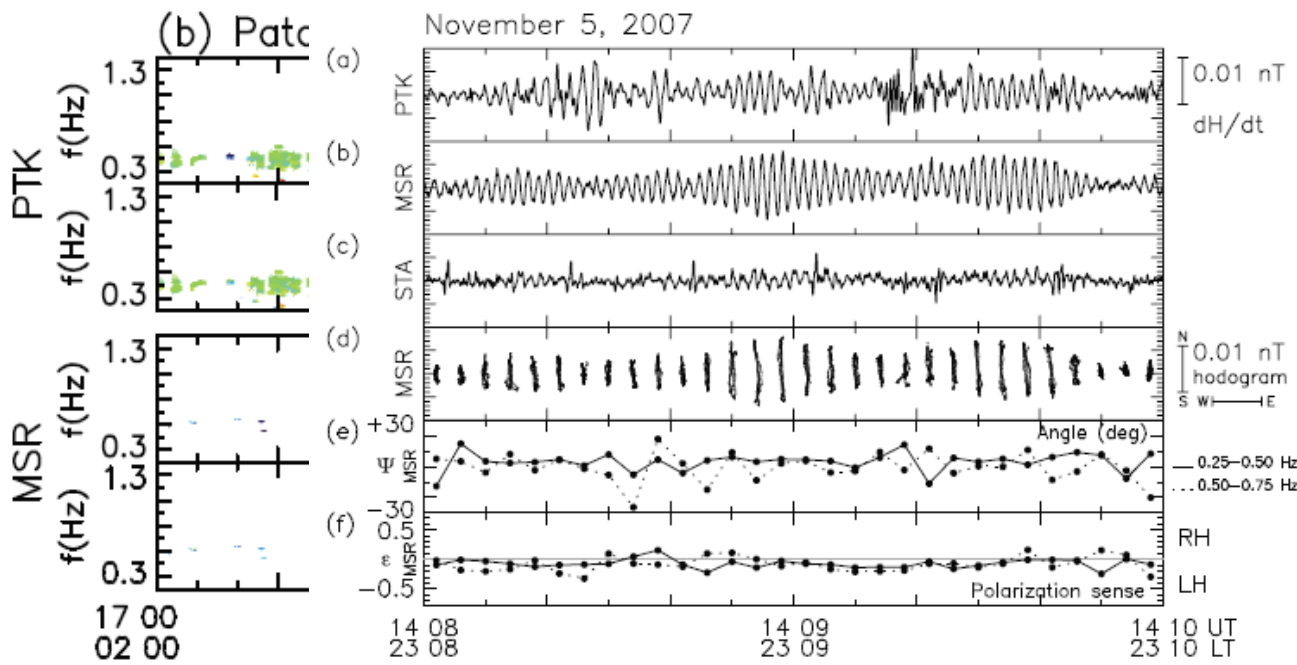
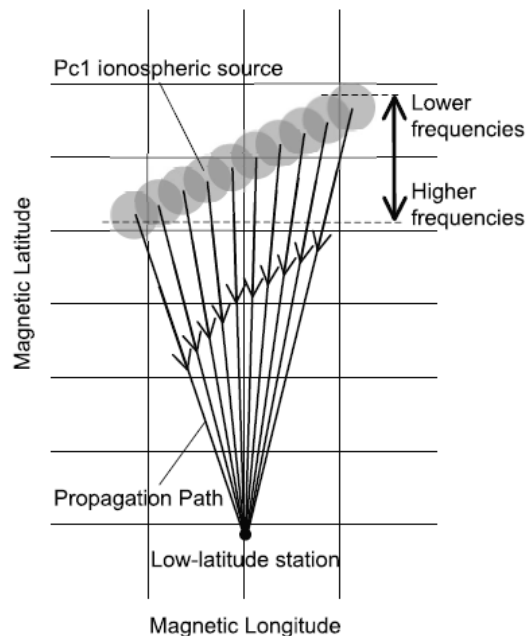


Pc1 polarization

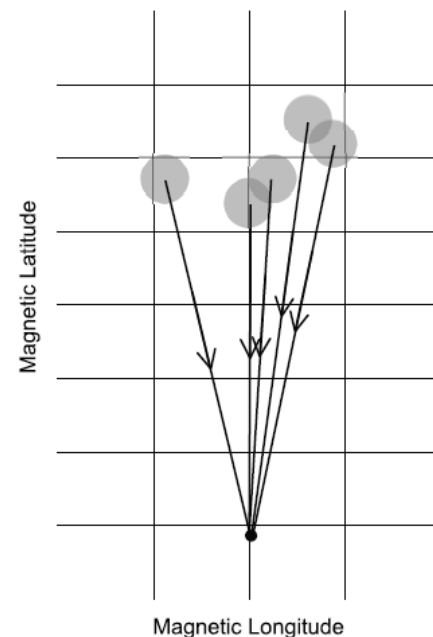
Nomura et al. (JGR, 2011)



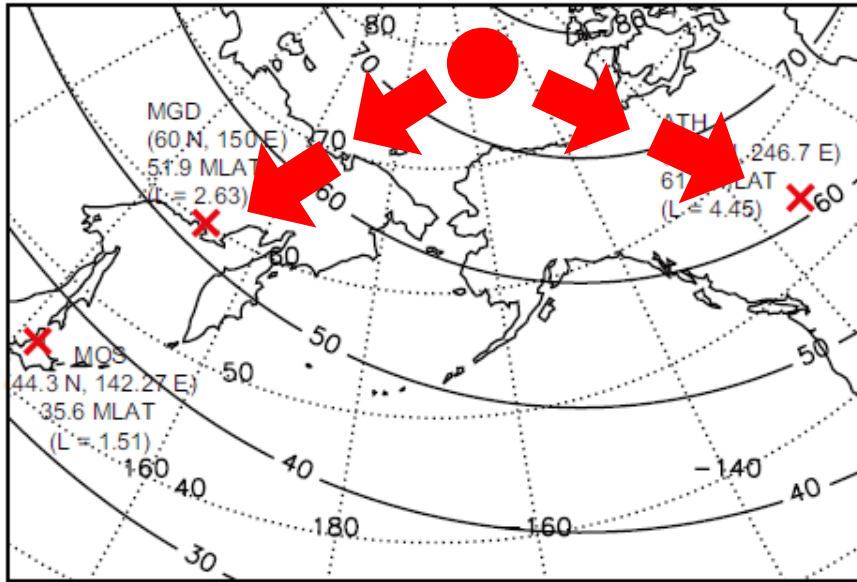
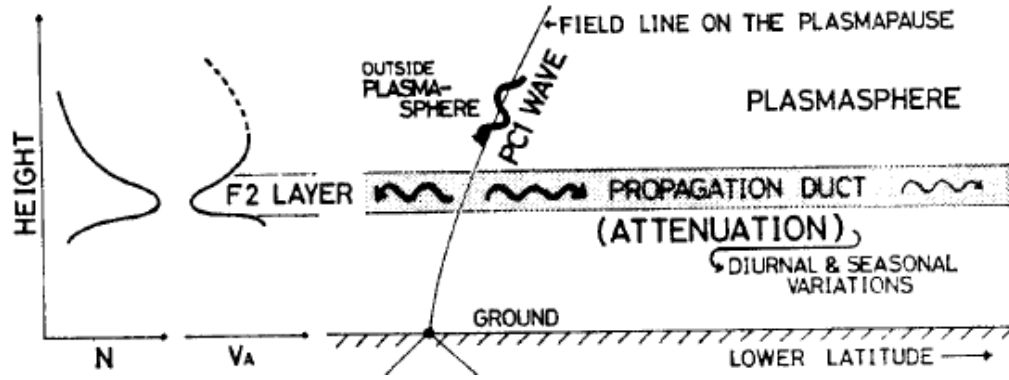
(a) Monotonic FDA event



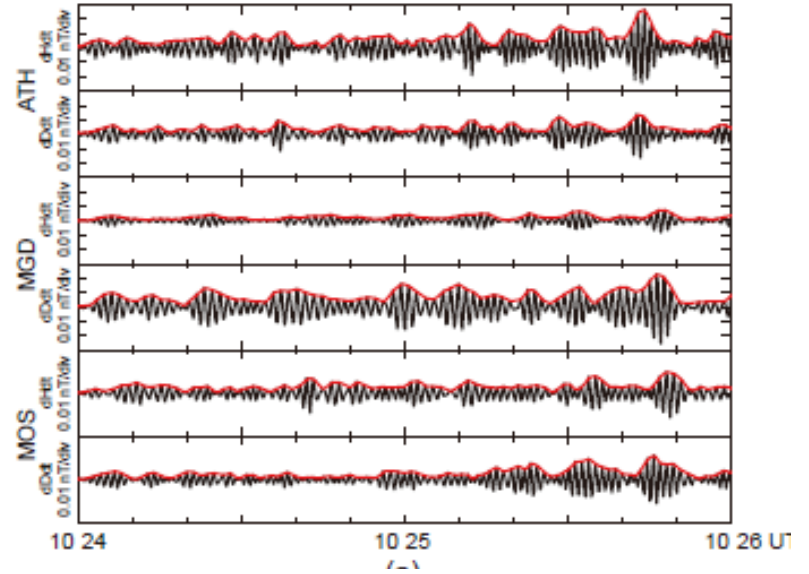
(b) Patchy FDA event



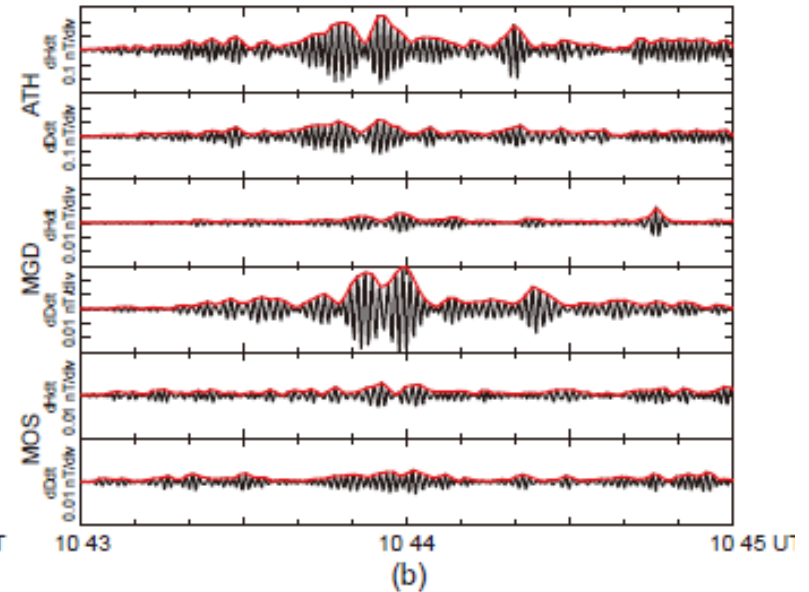
Pc1 Pearl Structure by ionospheric beating



April 8, 2010



April 8, 2010

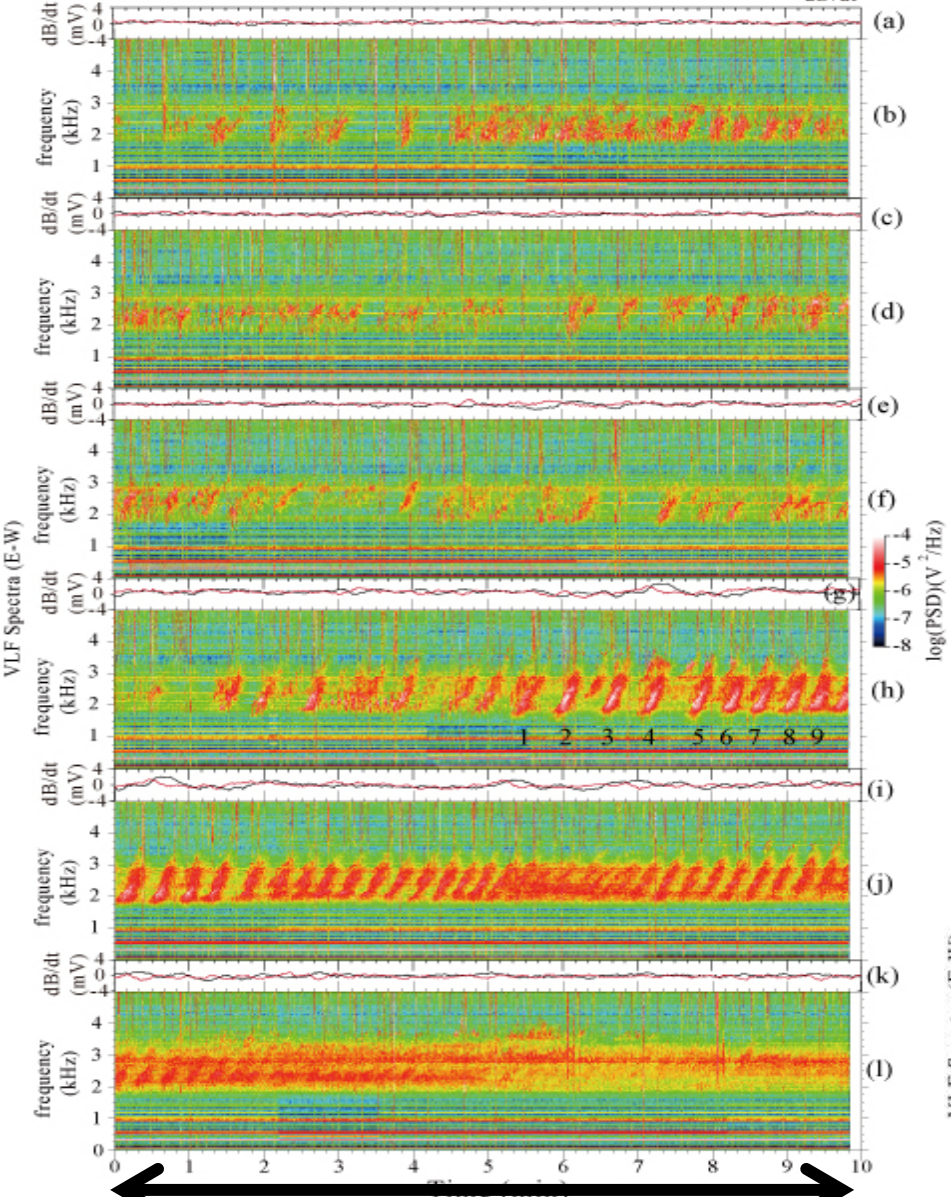


Beating during duct propagation in the ionosphere.

Jun et al. (EPS, in press, 2014)

QP emission / SC-triggered emission

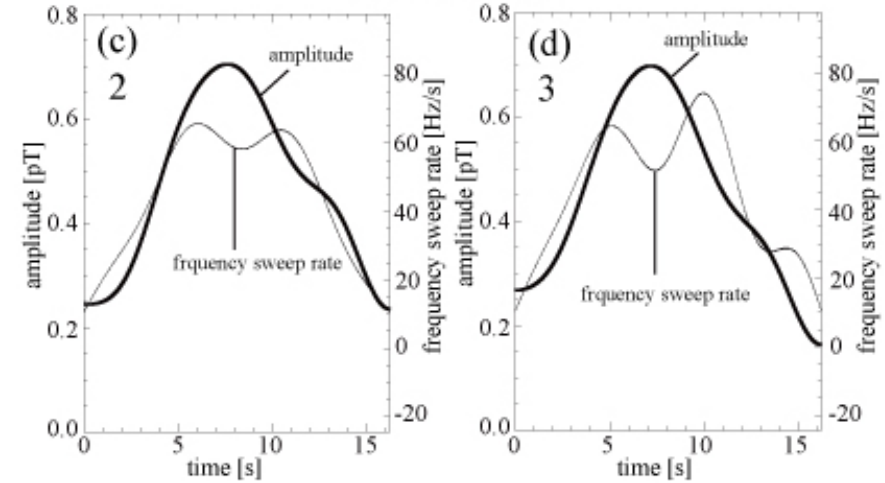
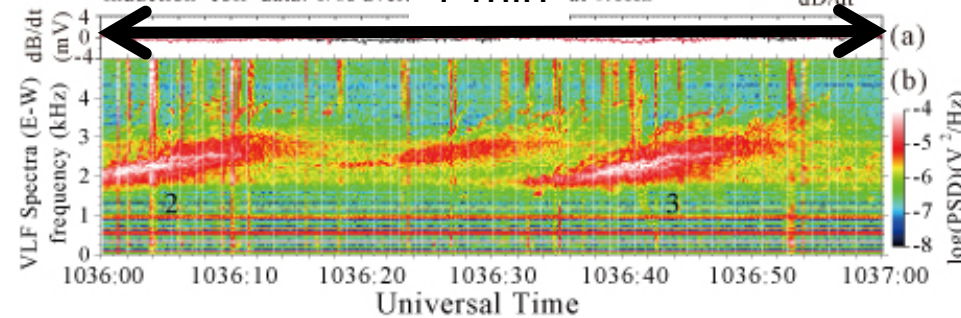
Feb.17, 2012 Athabasca 10:00:00-11:00:00UT
 induction coil data: 1s averages, 81 mV/nT at 0.1Hz



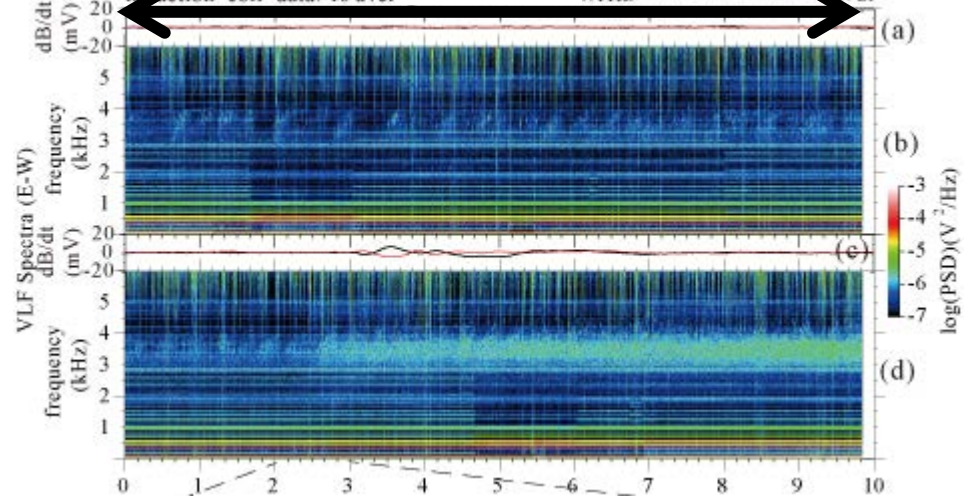
10-min

Shiokawa et al. (JGR, in press, 2014)

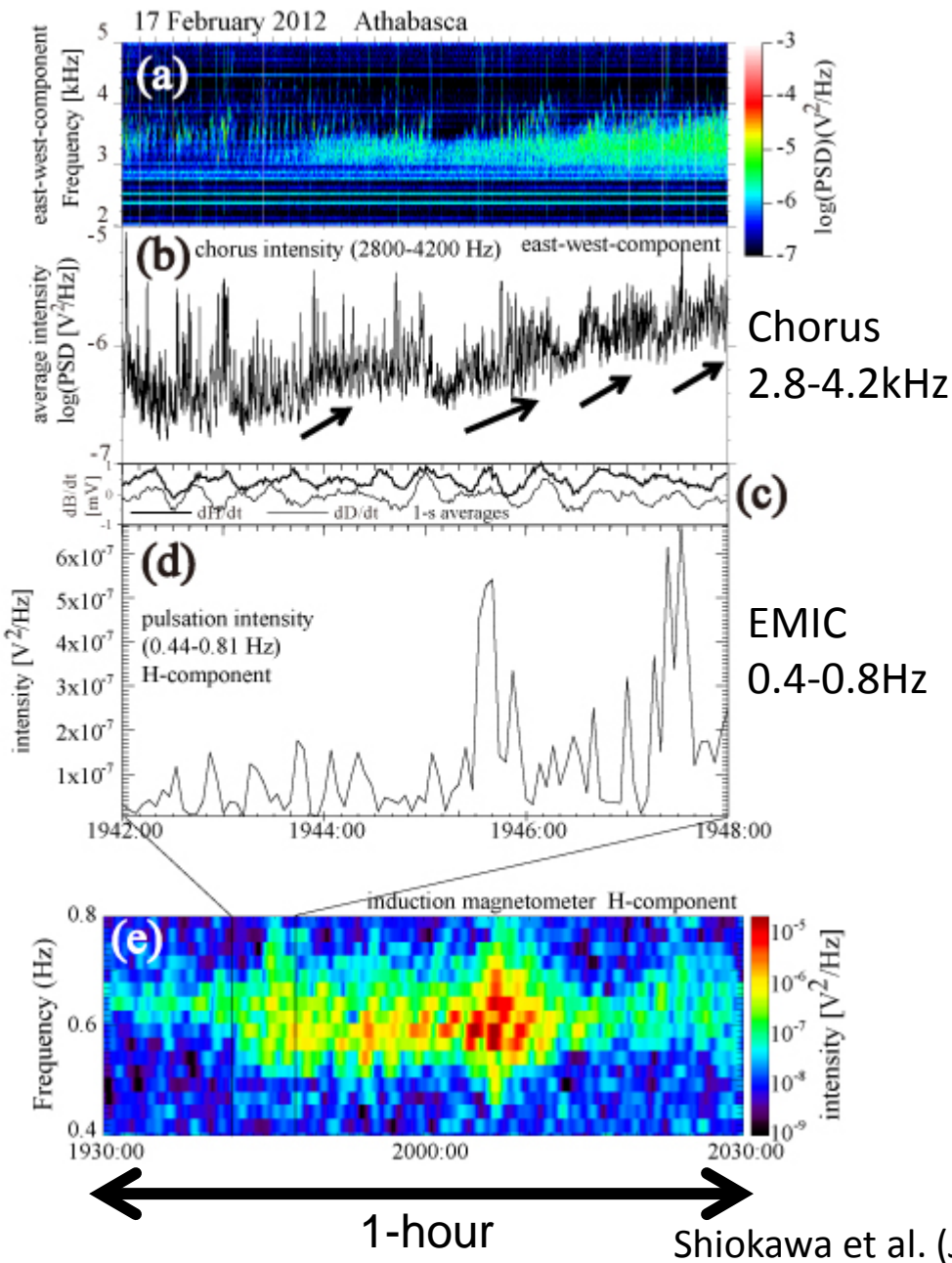
Feb.17, 2012 Athaba
 induction coil data: 1/8s aver: 1-min at 0.1Hz



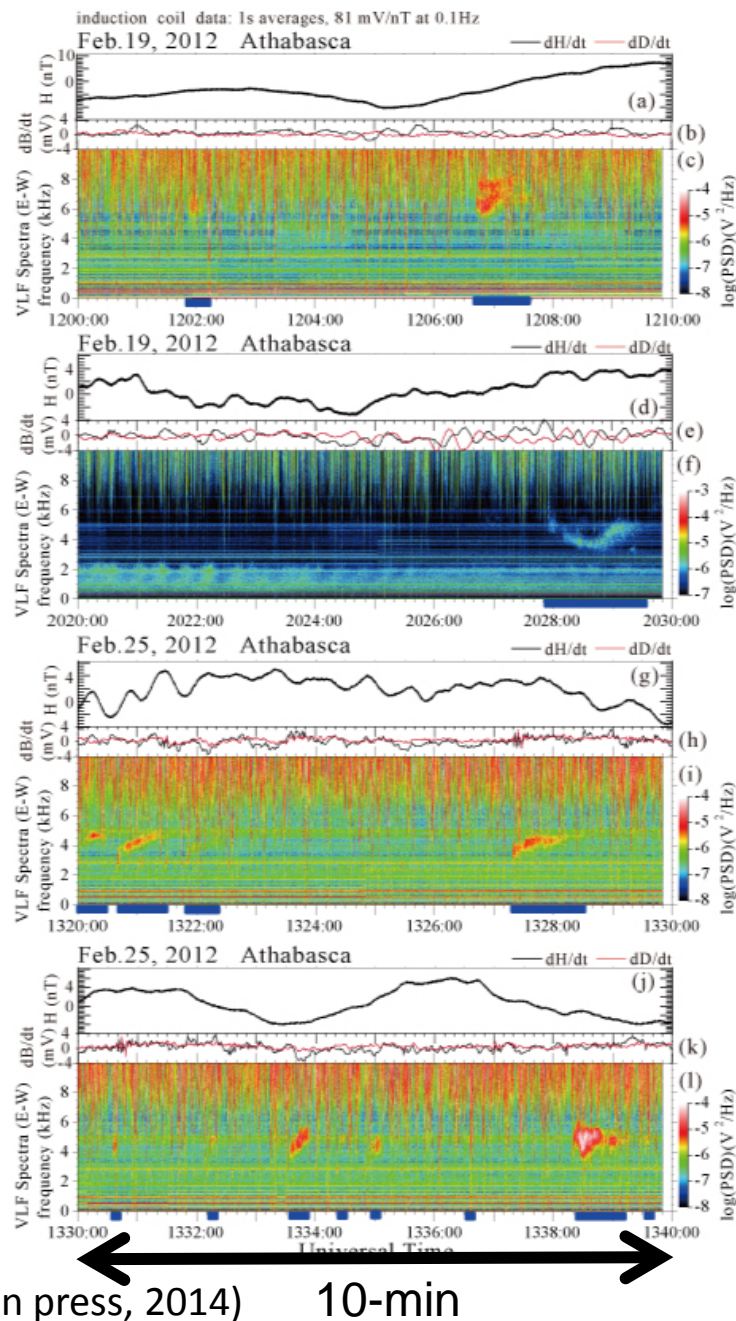
Feb.18, 2012 Athal
 induction coil data: 1s aver: 10-min 0-1650:00UT



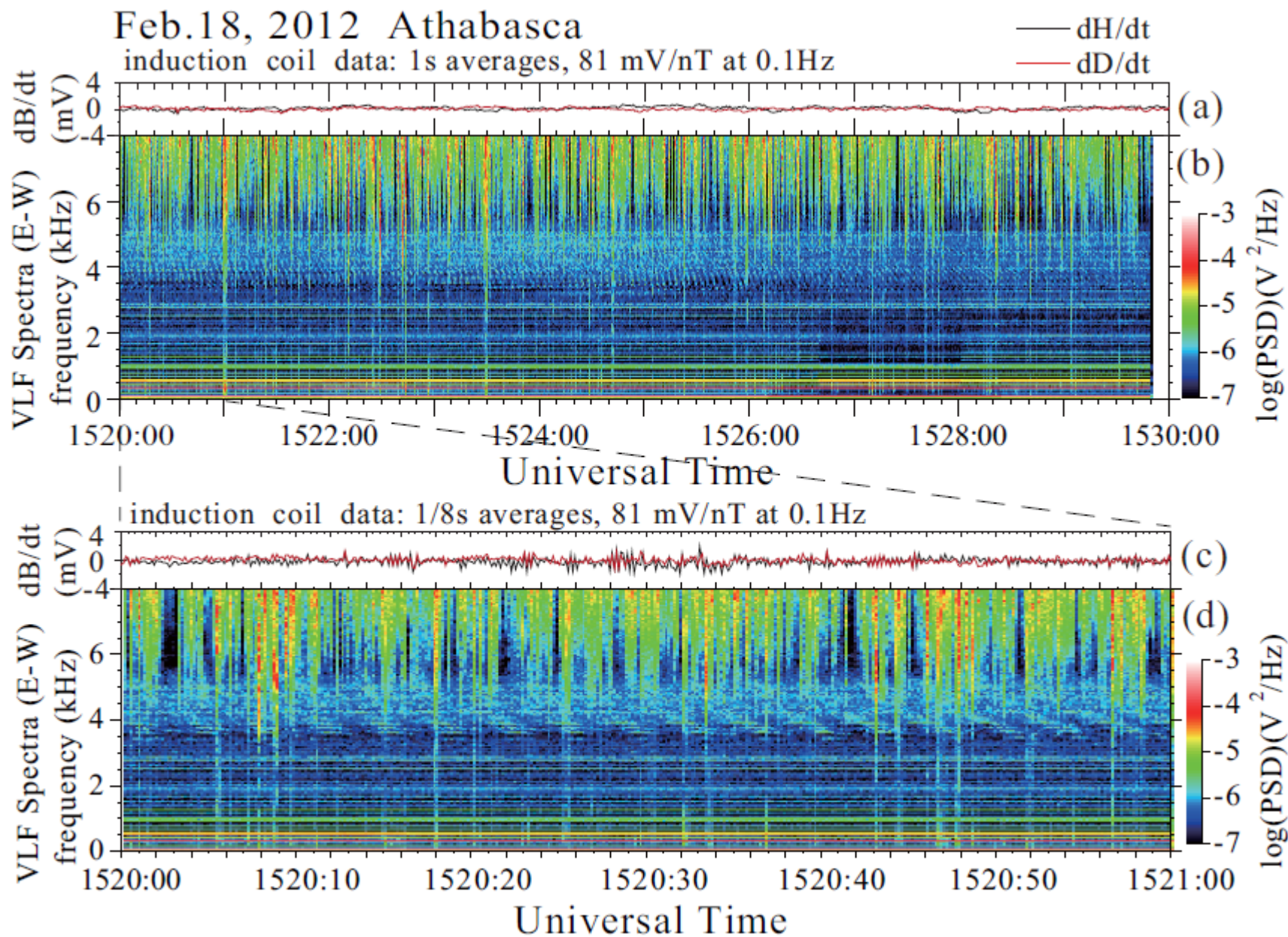
Chorus and EMIC correlation

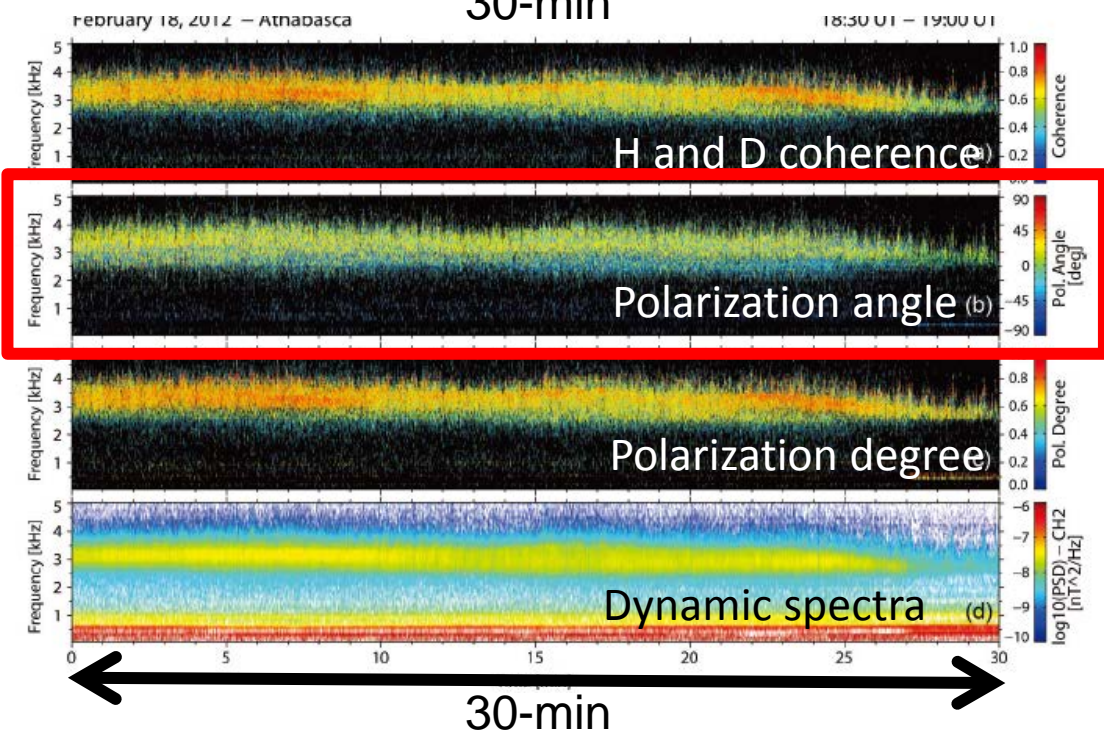
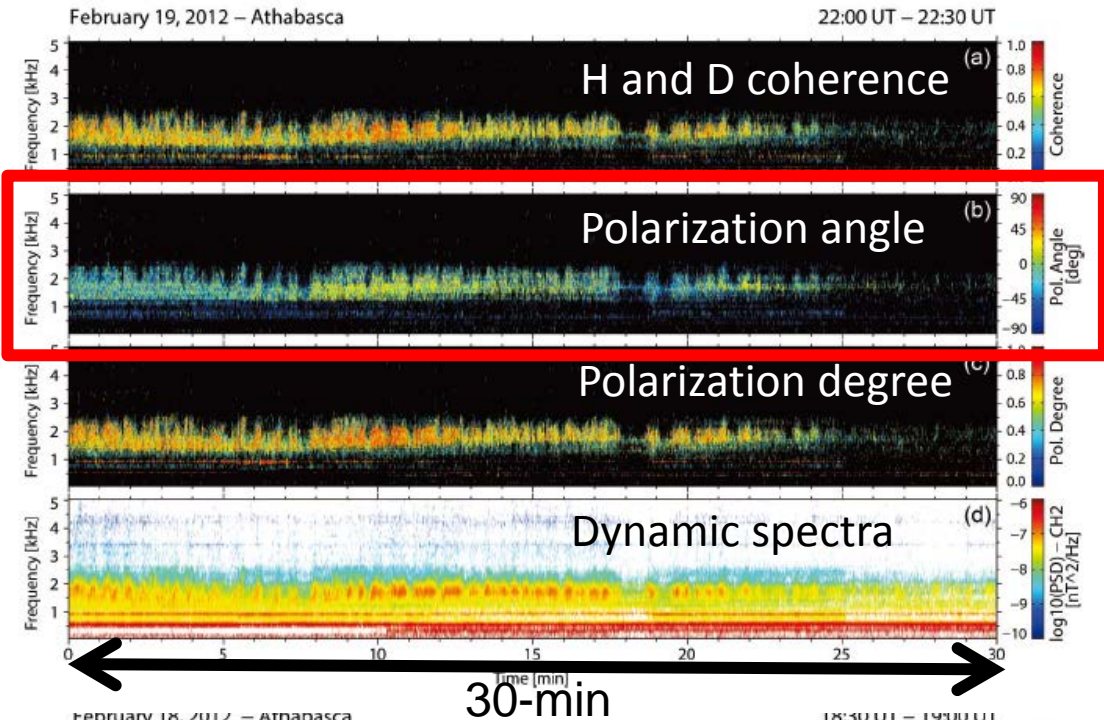


Bursty patch emissions

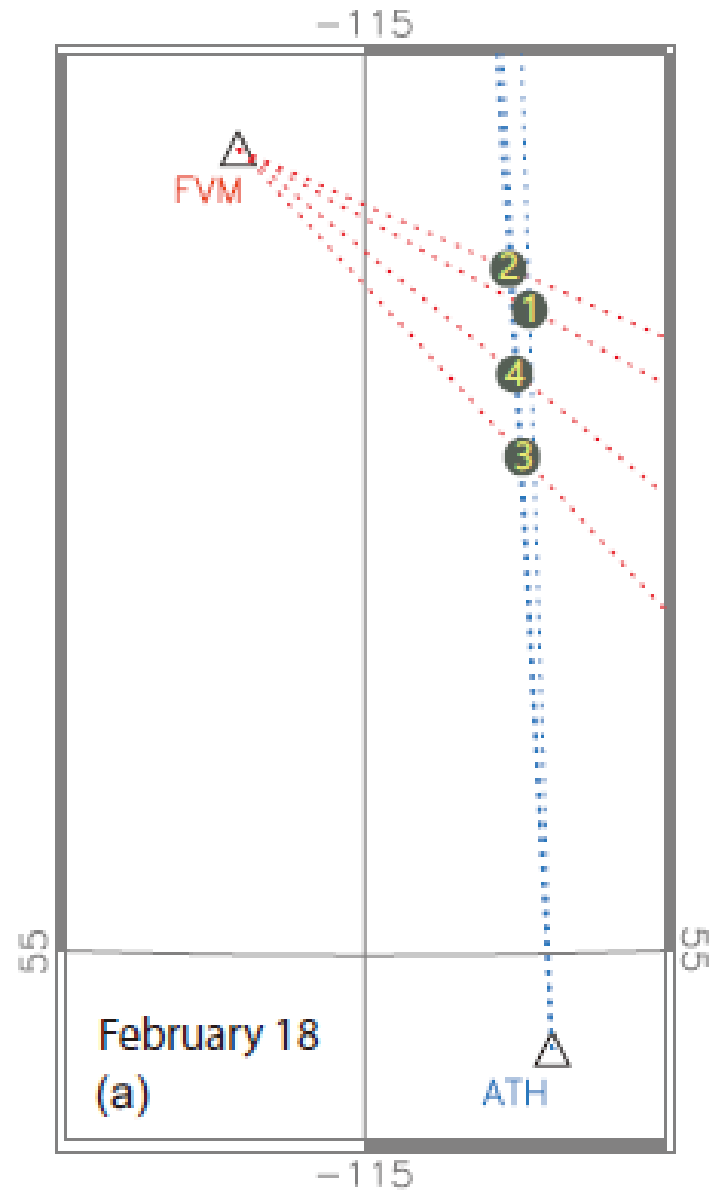


Falling tone chorus with sweep rate change from 0.7kHz/s to 0.05 kHz/s in 10 min





Chorus polarization analysis



Martinez et al. (EPS, in press, 2014)

 Observable from ground

relativistic electron
orbits (black arrows)

storm time ring
current **ions** &
electrons

EMIC
waves

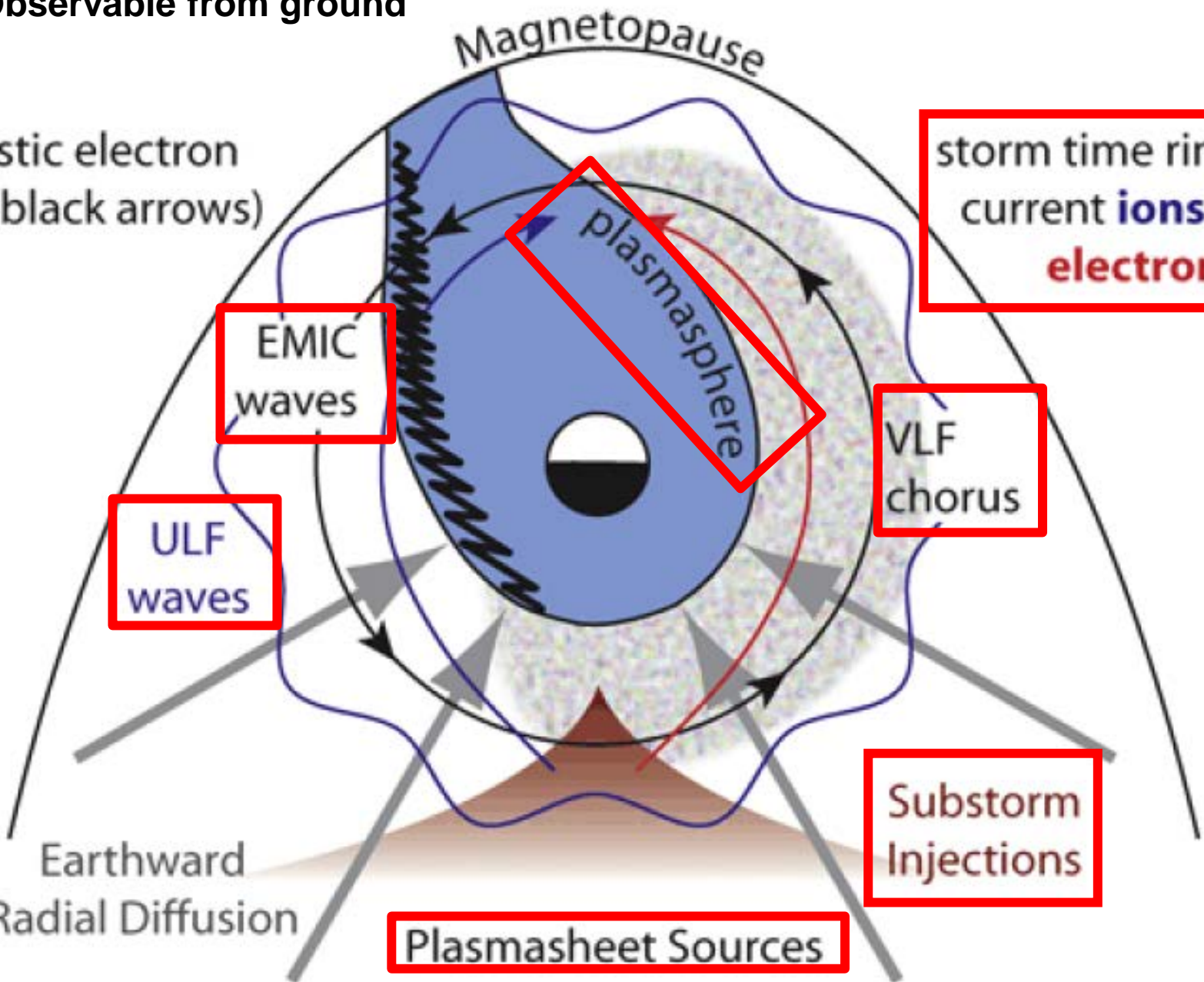
ULF
waves

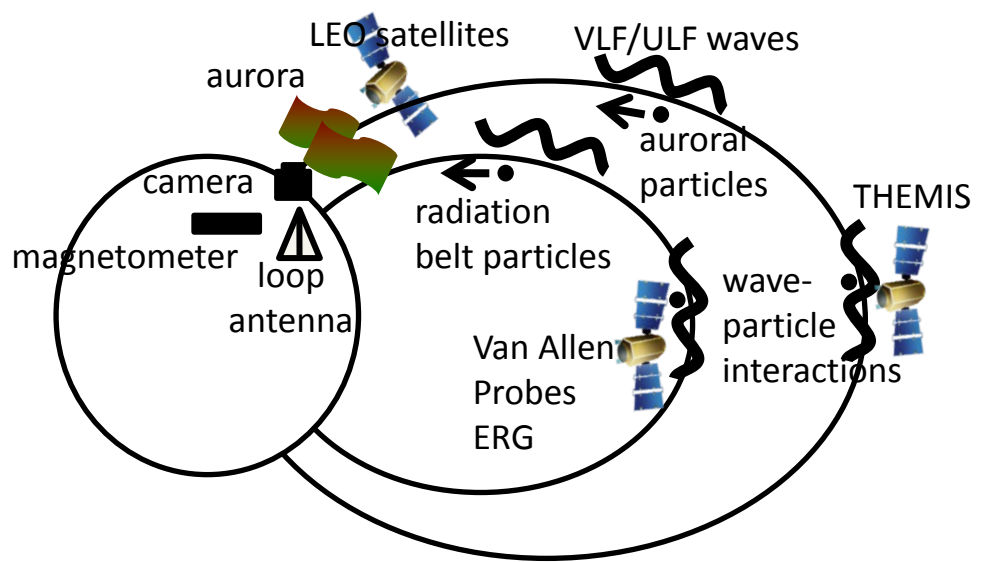
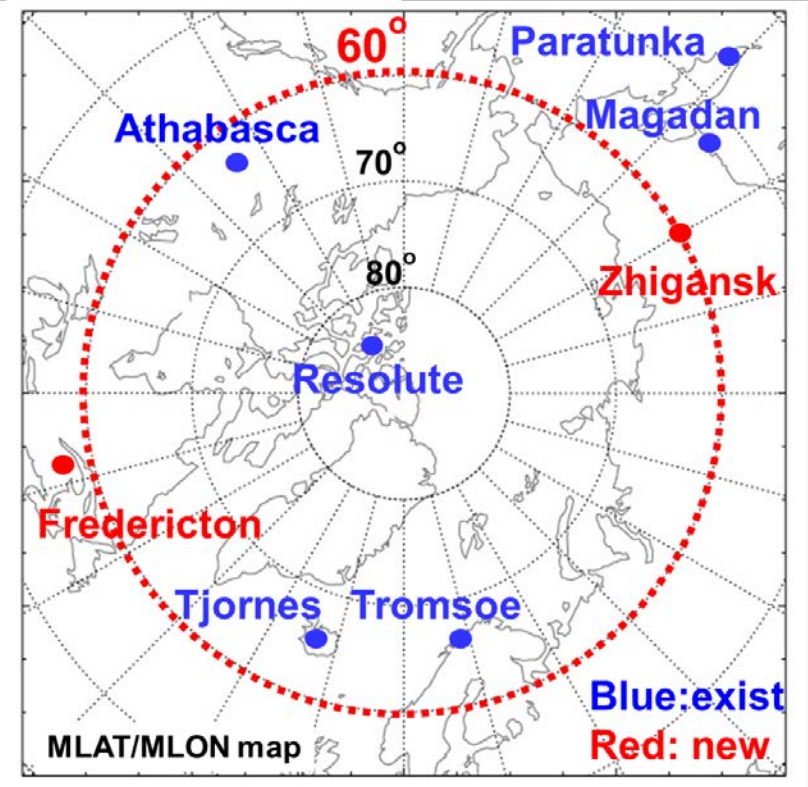
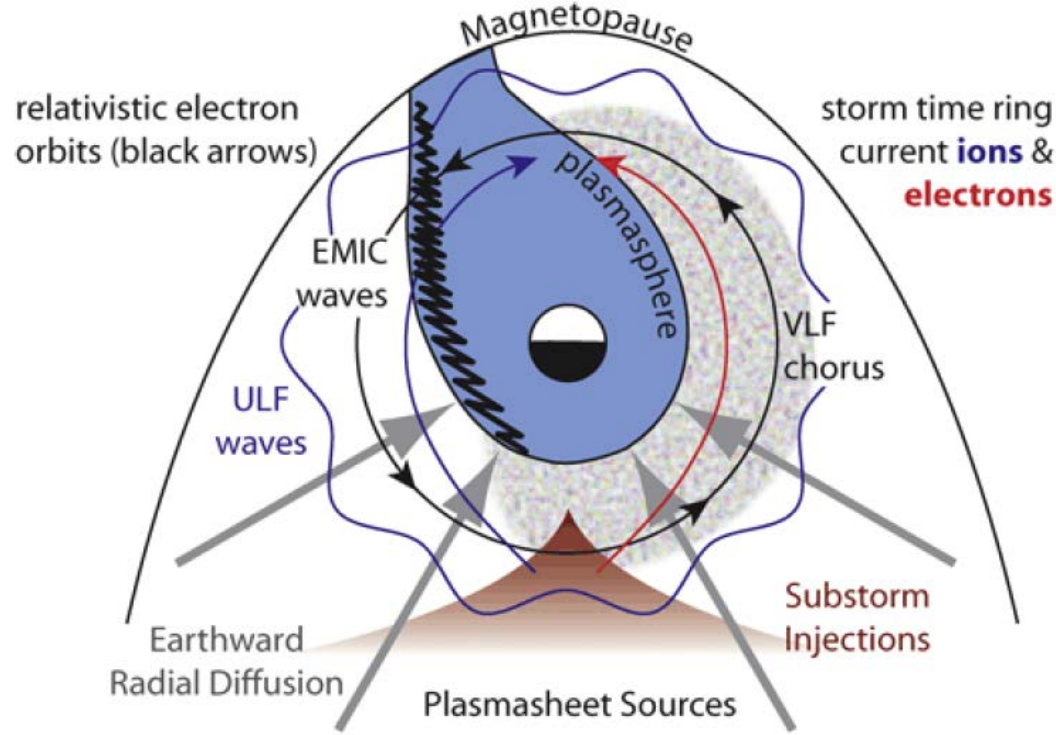
VLF
chorus

Earthward
Radial Diffusion

Substorm
Injections

Plasmasheet Sources





Summary

Aurora/Airglow Imagers

- MSTID motion** → monitoring of penetrating electric field
- auroral fragmentation** into patches → ballooning/interchange?
- SAR arc** → plasmasphere-ring current interaction

Induction magnetometers

- isolated proton aurora** → monitoring of wave-particle interaction
- Pc1 polarization** characteristics (dependence on frequency and distance from the source)
- Pc1 pearl** structure (amplitude modulation of EMIC waves)
 - Ionospheric beating versus magnetospheric processes

ELF/VLF receivers

- temporal variation of **QP** emission and **falling tone** emission
- positive correlation of **frequency sweep rate and intensity** for QP
- correlation between **EMIC and chorus** wave intensity
- bursty patch** emission and **ssc-triggered** emission
- frequency and temporal dependence of **chorus polarization**
- correlation with pulsating aurora – **30-40s switching** of interaction