



# Birth and evolution of magnetosheath mirror modes as seen by the global hybrid-Vlasov simulation Vlasiator

**S. Hoilijoki**<sup>1,2</sup> , B. Walsh<sup>3</sup> , Y. Kempf<sup>1,2</sup> , S. von Alfthan<sup>1</sup> , O. Hannuksela<sup>1,2</sup> and M. Palmroth<sup>1</sup>

<sup>1</sup>Finnish Meteorological Institute, Helsinki, Finland

<sup>2</sup>University of Helsinki, Helsinki, Finland

<sup>3</sup>Space Sciences Laboratory, University of California, Berkeley, California, USA

Geospace revisited, Rhodes, Greece, 2014



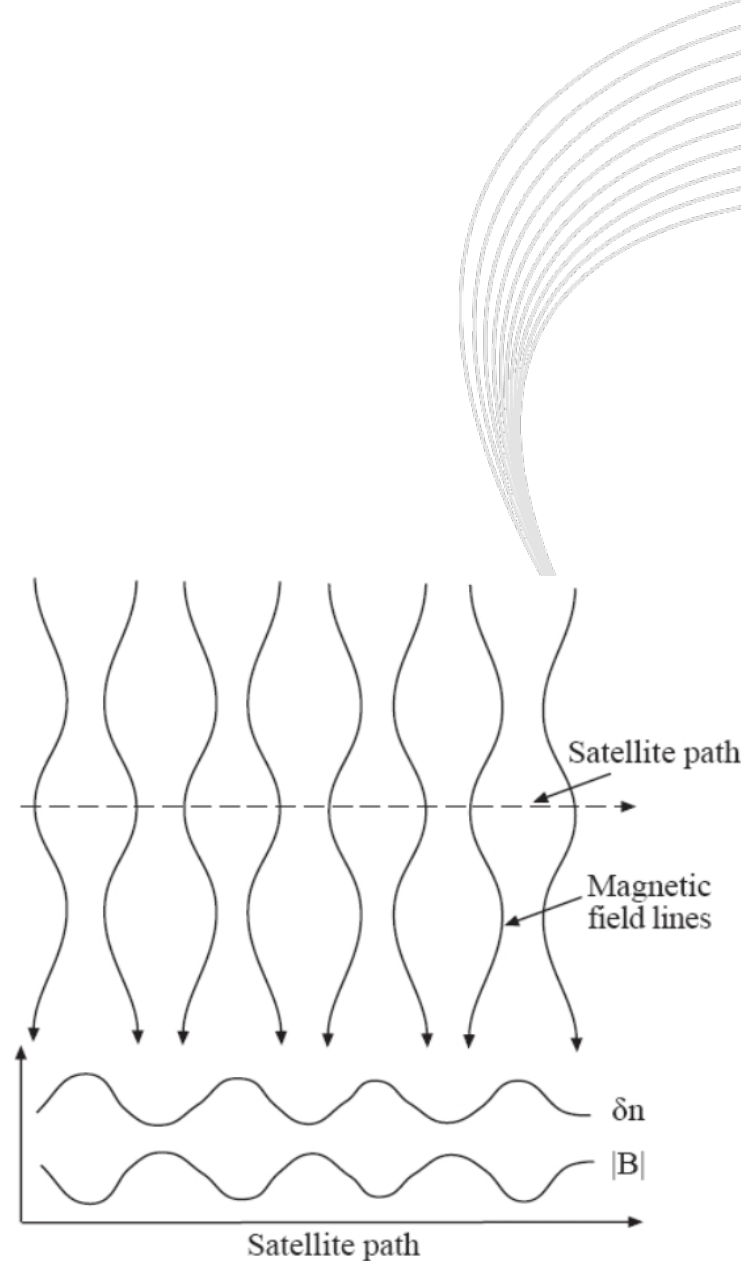
# Outline

- Mirror modes
- Vlasiator
- Mirror modes in Vlasiator
- Comparison with THEMIS observations
- Conclusions



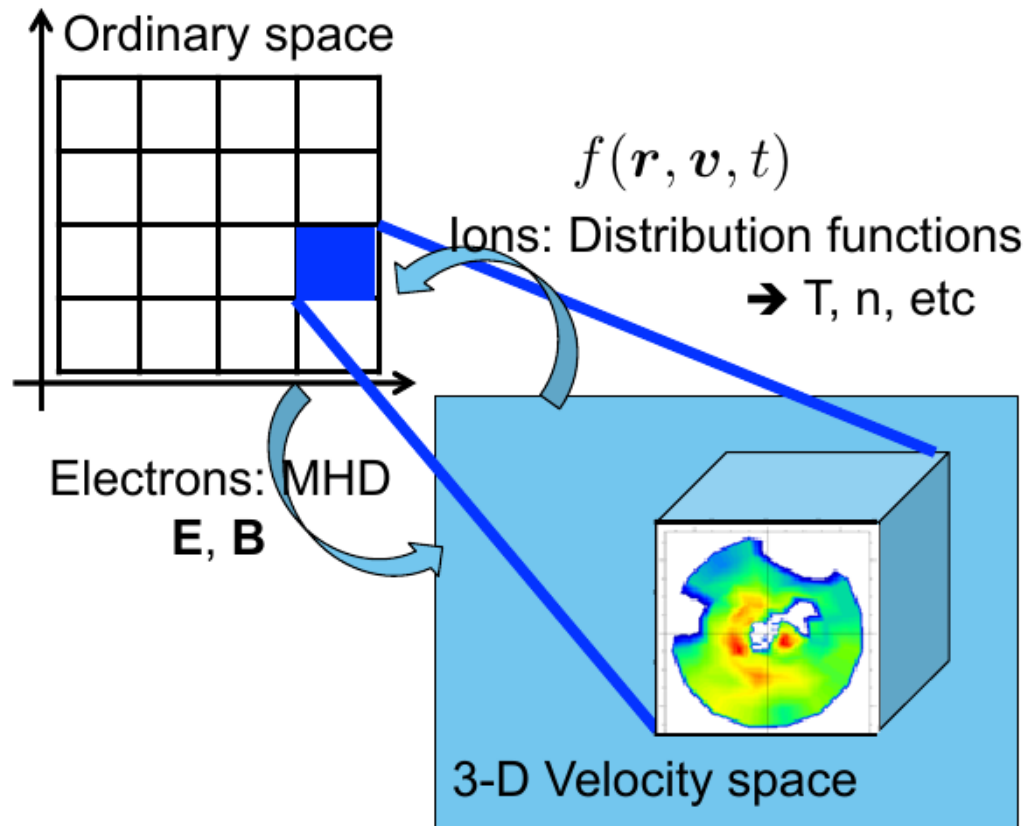
# Mirror modes

- Physical properties:
  - Large amplitude, low frequency waves
  - Anti-correlation between magnetic field strength and plasma density
  - Arise from temperature anisotropy in region fulfilling the condition:
    - $T_{\perp}/T_{\parallel} > 1 + 1/\beta_{\perp}$  (Hasegawa, 1969)
  - Almost linearly polarized
- Criteria for observations:
  - Variance matrix eigenvalues fulfill:
    - $\lambda_{\max}/\lambda_{\text{int}} \gg \lambda_{\text{int}}/\lambda_{\min} \sim 1$  (Wilson III et al, 2009)
    - $\lambda_{\max}/\lambda_{\text{int}} > 1.5$  and  $\lambda_{\min}/\lambda_{\text{int}} > 0.3$  (Soucek et al, 2008)
  - Angle between ambient magnetic field and maximum variance direction less than  $30^{\circ}$  (Soucek et al, 2008)
  - Amplitude with respect to background  $> 10\%$



# Vlasiator scheme in a nutshell

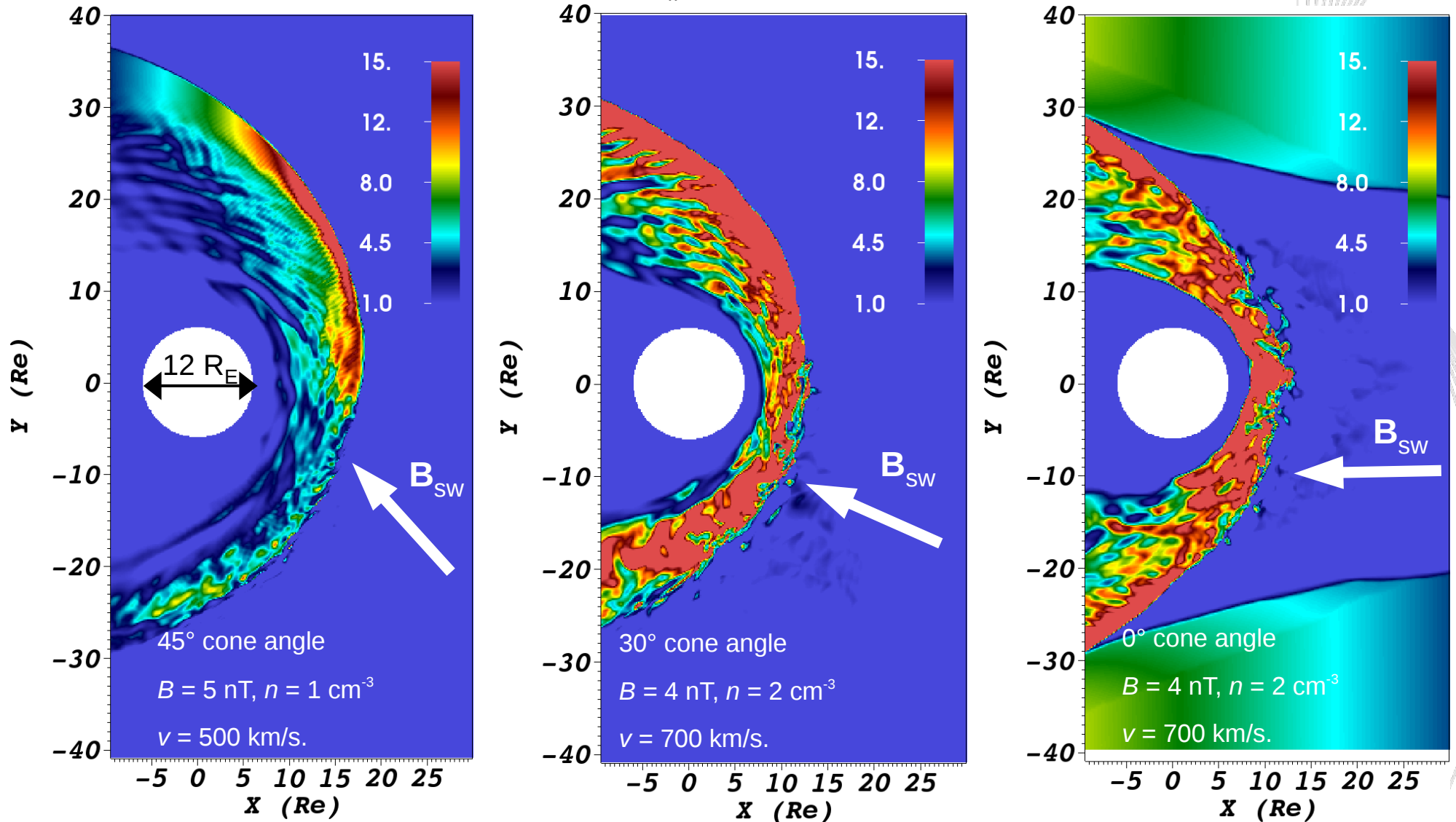
- Compute EM-fields in ordinary space
- Each ordinary space cell contains a 3D velocity space
- In velocity space, propagate distribution function with Vlasov equation
  - Couple back to ordinary space to compute EM-fields





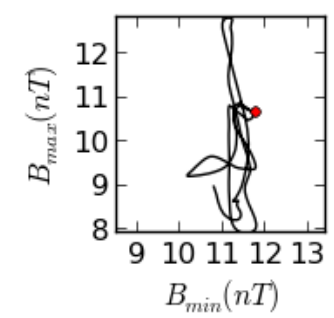
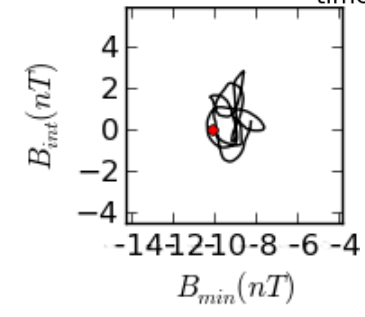
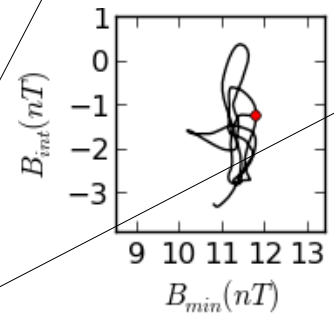
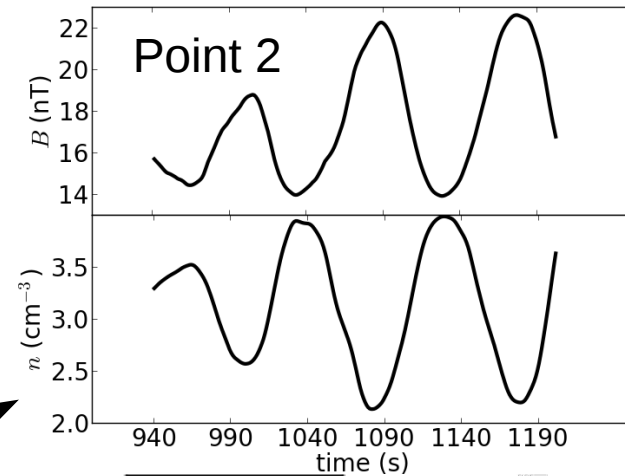
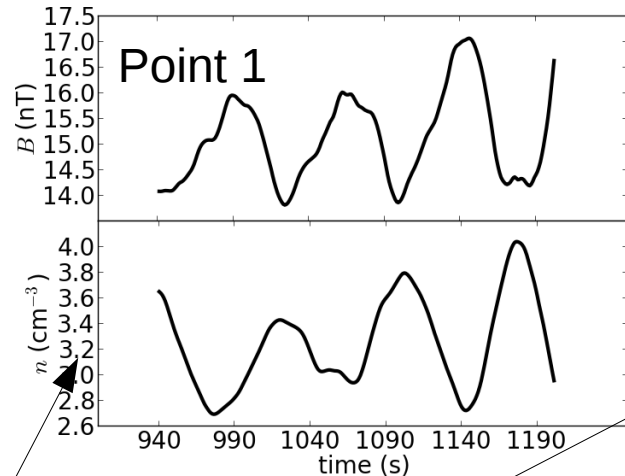
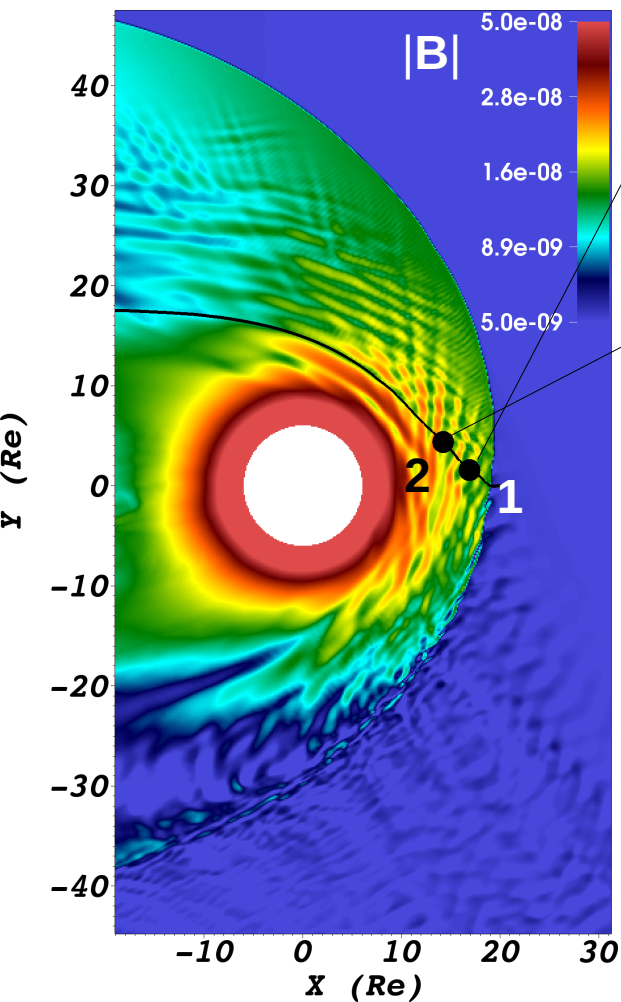
# Runs

Color: mirror instability condition  $\beta_{\perp} (T_{\perp}/T_{\parallel}-1) > 1$  at  $t = 1000$  s

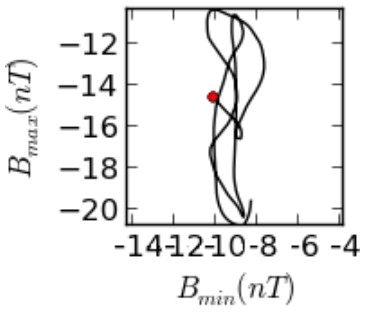




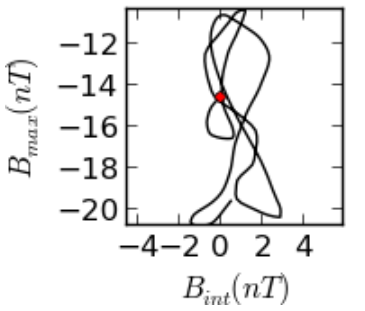
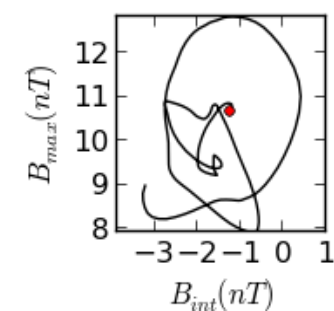
# Wave analysis



$\lambda_{\max} / \lambda_{\text{int}} = 2.6$   
 $\lambda_{\text{int}} / \lambda_{\min} = 4.8$   
 $\lambda_{\min} / \lambda_{\text{int}} = 0.2$



$\lambda_{\max} / \lambda_{\text{int}} = 10.8$   
 $\lambda_{\text{int}} / \lambda_{\min} = 2.6$   
 $\lambda_{\min} / \lambda_{\text{int}} = 0.4$





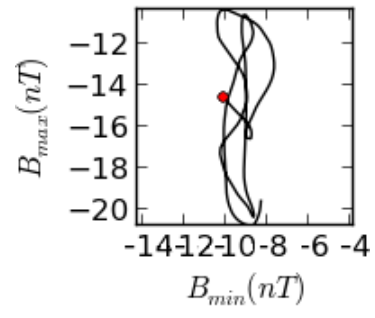
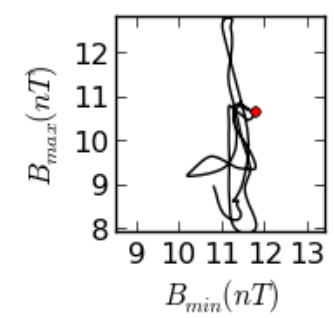
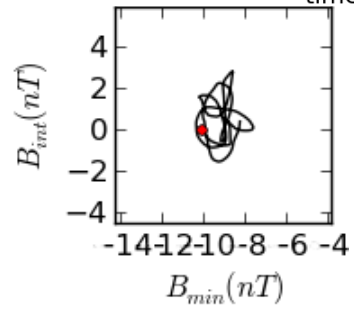
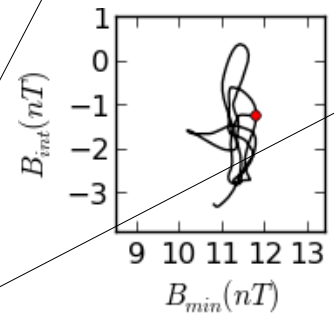
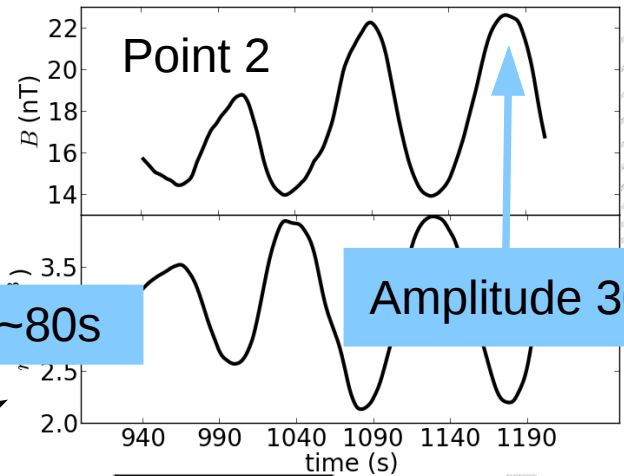
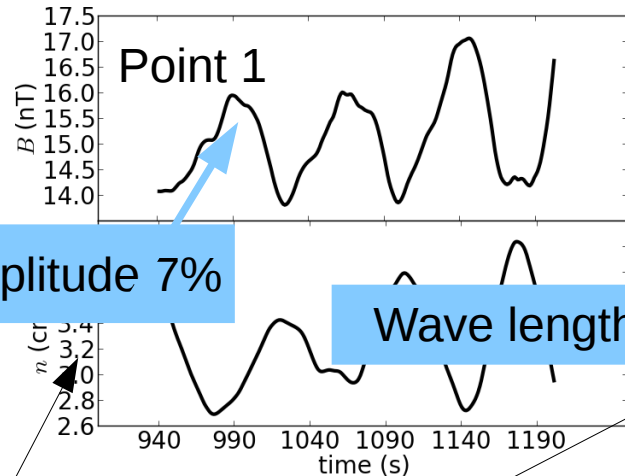
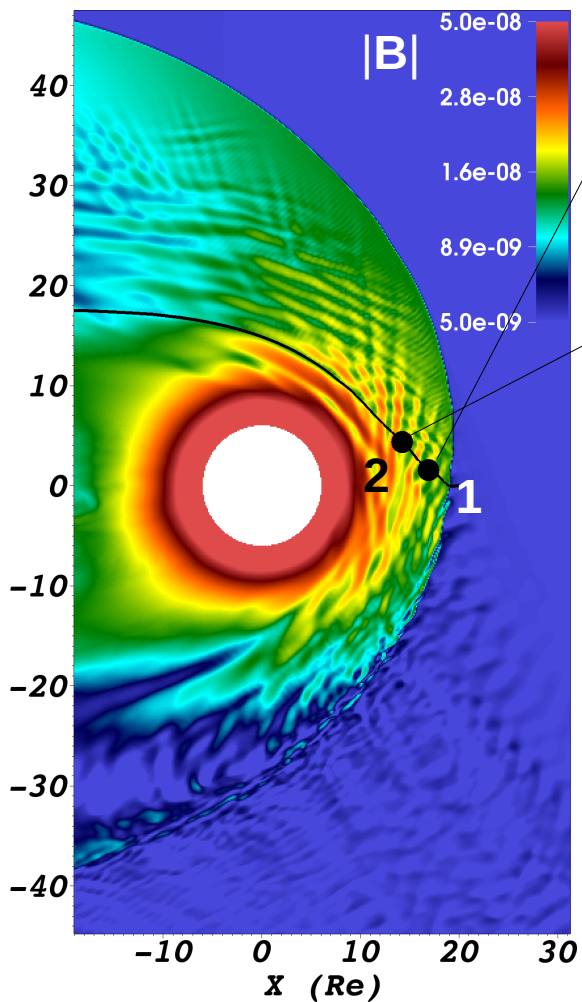


# Wave analysis

Amplitude 7%

Wave length ~80s

Amplitude 30%



$$\lambda_{\max} / \lambda_{\text{int}} = 2.6$$

$$\lambda_{\text{int}} / \lambda_{\min} = 4.8$$

$$\lambda_{\min} / \lambda_{\text{int}} = 0.2$$

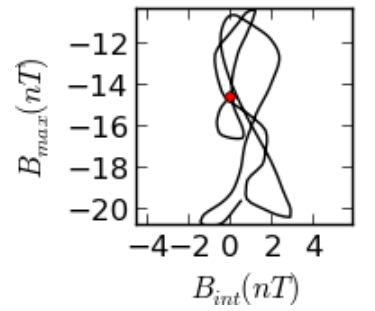
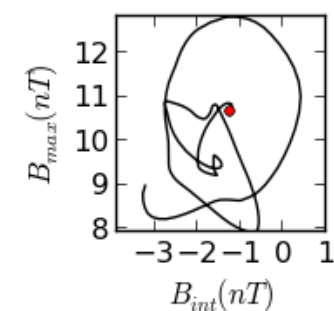
$$\lambda_{\max} / \lambda_{\text{int}} = 10.8$$

$$\lambda_{\text{int}} / \lambda_{\min} = 2.6$$

$$\lambda_{\min} / \lambda_{\text{int}} = 0.4$$

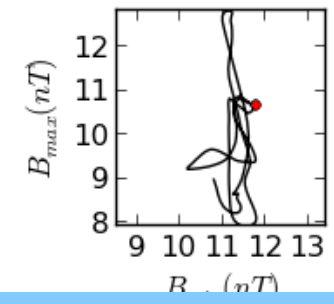
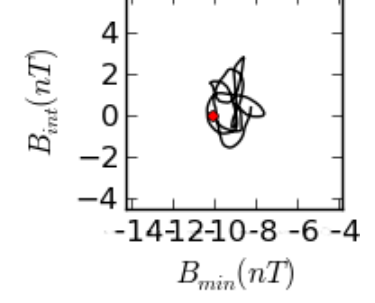
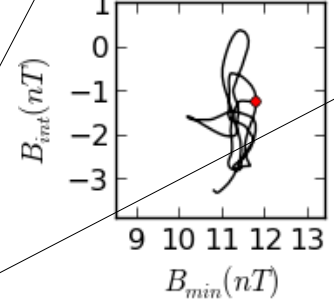
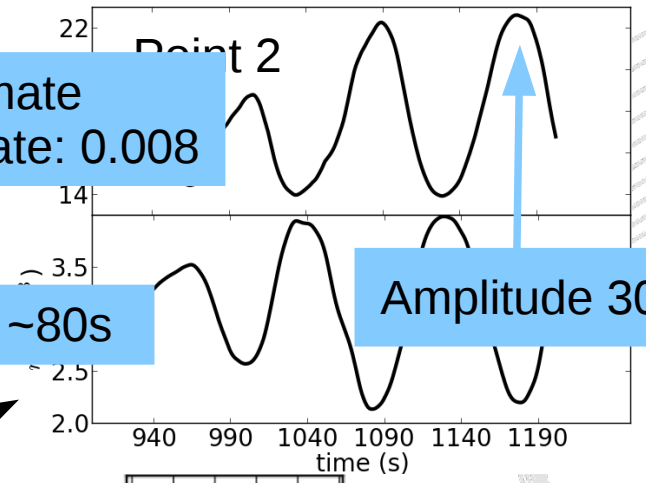
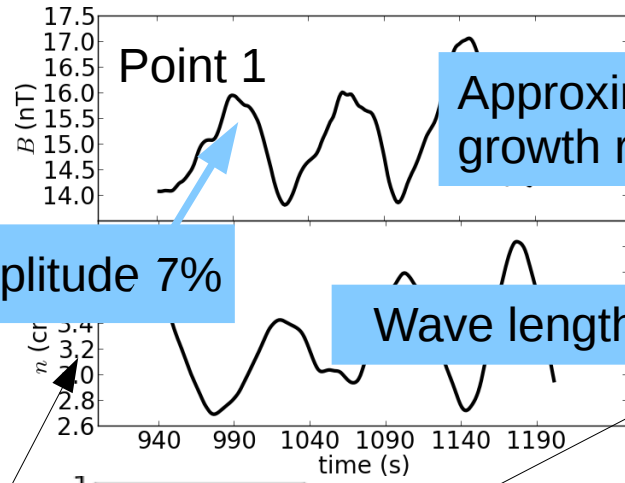
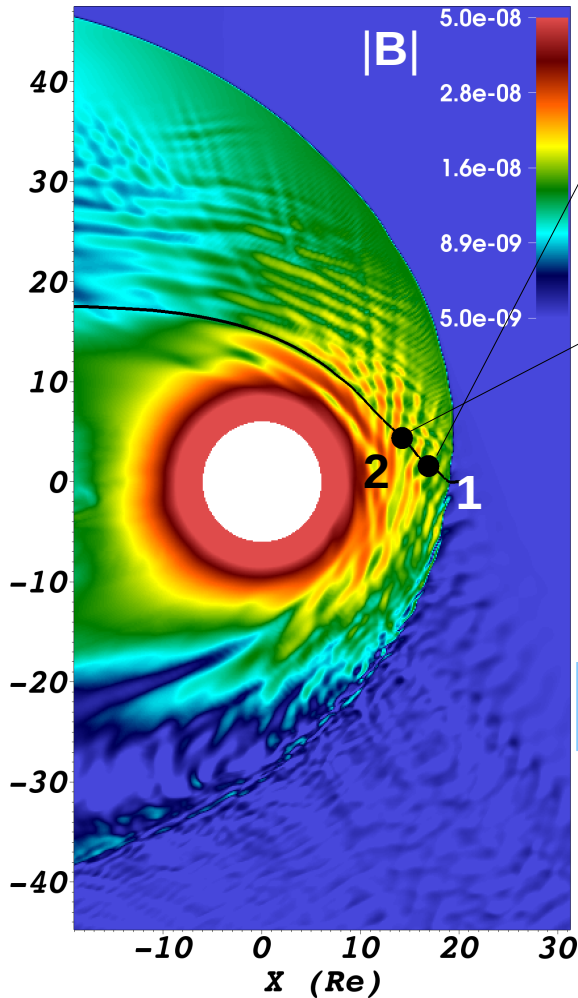
Criteria:

- $\lambda_{\max} / \lambda_{\text{int}} \gg \lambda_{\text{int}} / \lambda_{\min} \sim 1$
- $\lambda_{\max} / \lambda_{\text{int}} > 1.5$
- $\lambda_{\min} / \lambda_{\text{int}} > 0.3$

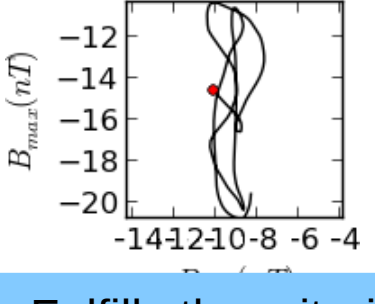




# Wave analysis



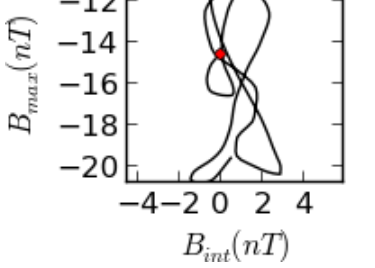
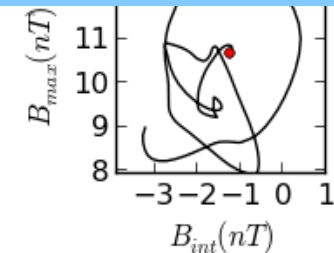
$\lambda_{max} / \lambda_{int} = 2.6$   
 $\lambda_{int} / \lambda_{min} = 4.8$   
 $\lambda_{min} / \lambda_{int} = 0.2$



$\lambda_{max} / \lambda_{int} = 10.8$   
 $\lambda_{int} / \lambda_{min} = 2.6$   
 $\lambda_{min} / \lambda_{int} = 0.4$

Does not fulfill the criteria

Fulfills the criteria



**Criteria:**  
 $\lambda_{max} / \lambda_{int} \gg \lambda_{int} / \lambda_{min} \sim 1$   
 $\lambda_{max} / \lambda_{int} > 1.5$   
 $\lambda_{min} / \lambda_{int} > 0.3$





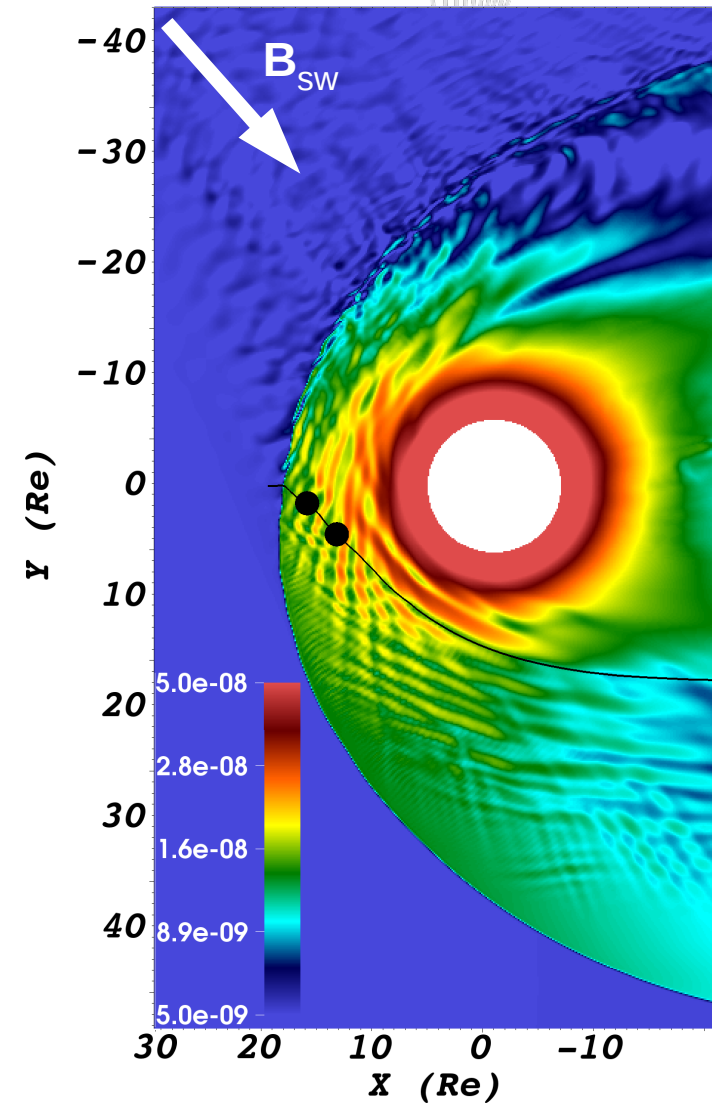
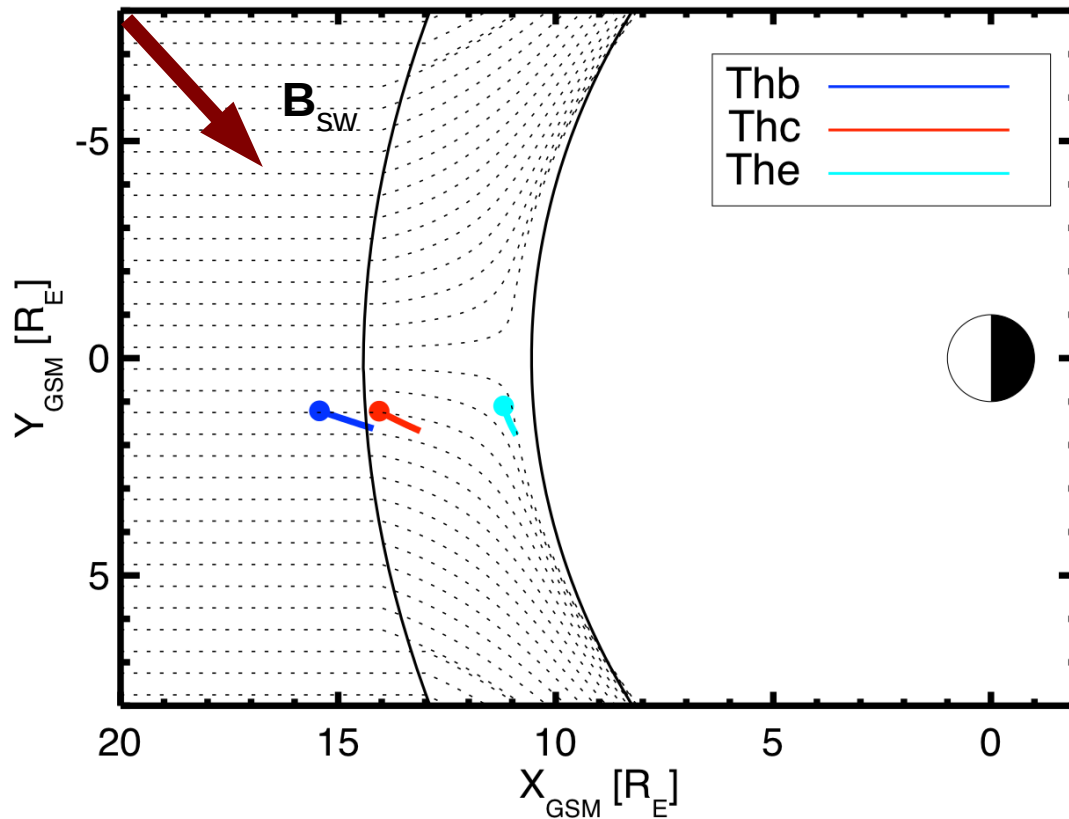
# Observations

**Themis B** in solar wind

**Themis C** in magnetosheath behind bow shock

**Themis E** in magnetosheath close to magnetopause

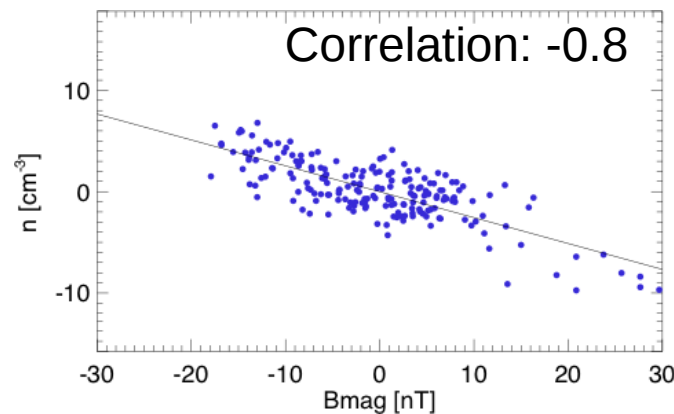
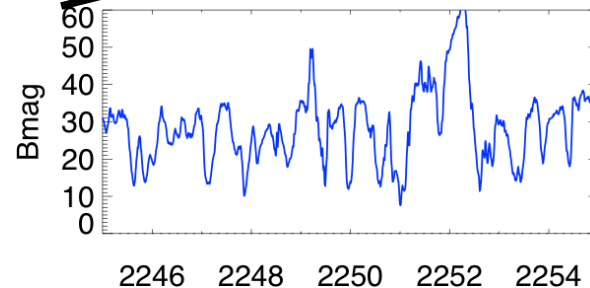
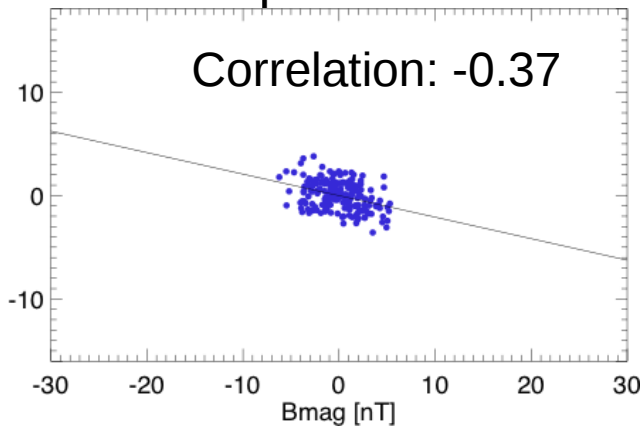
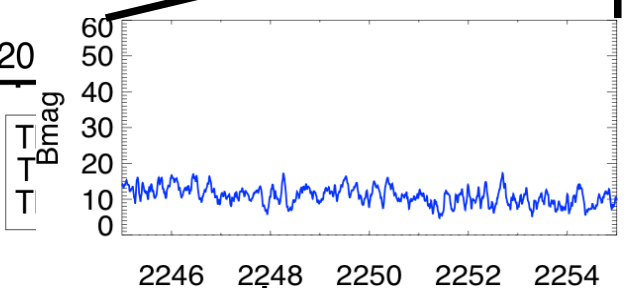
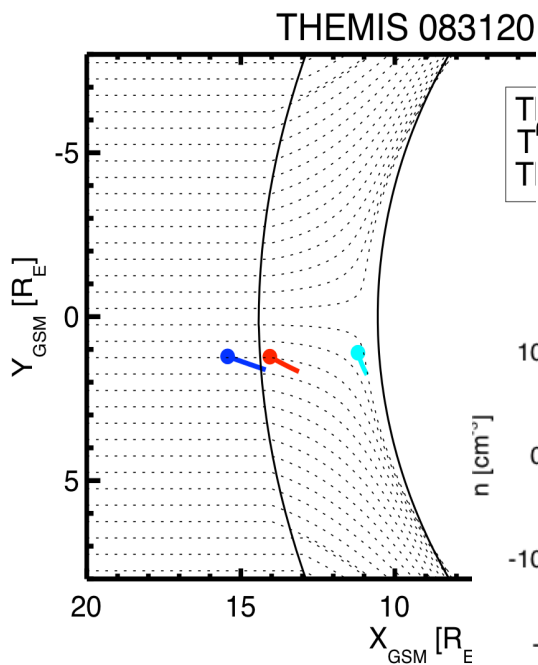
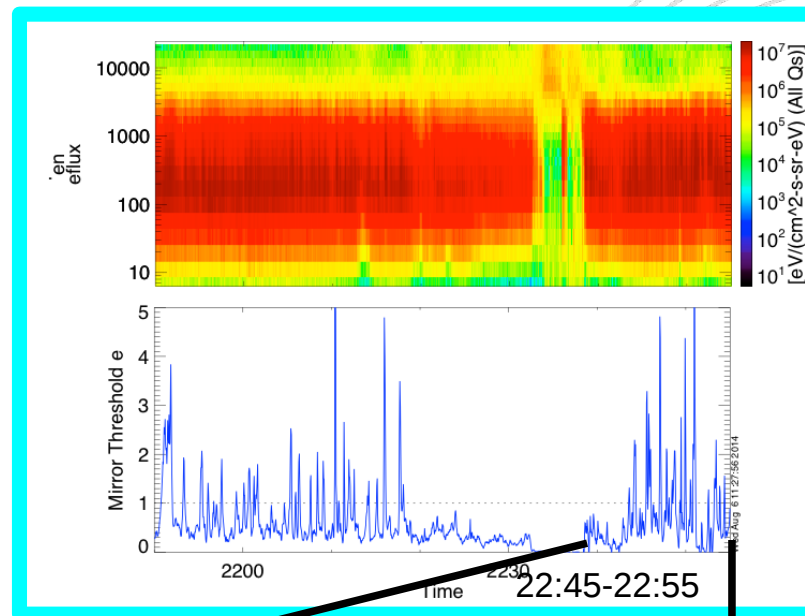
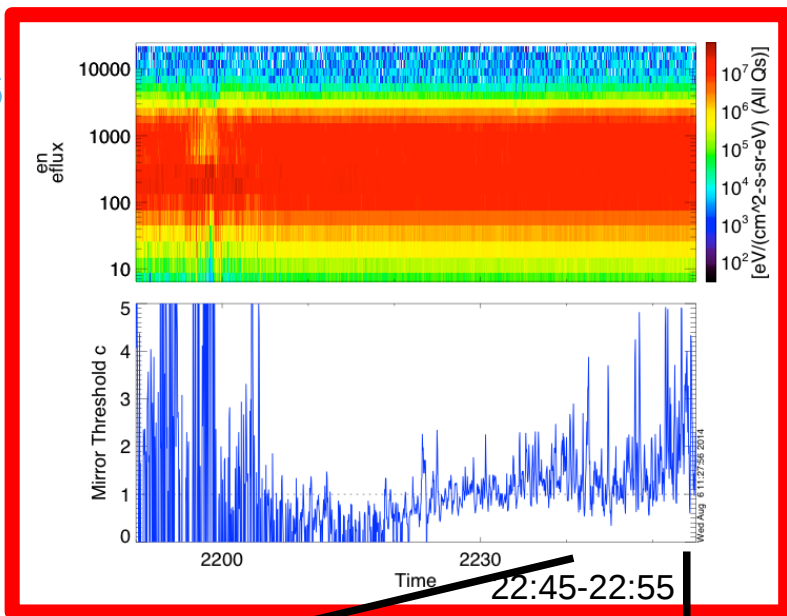
Outward Parker spiral THEMIS 08312008





# Mirror modes in THEMIS observations

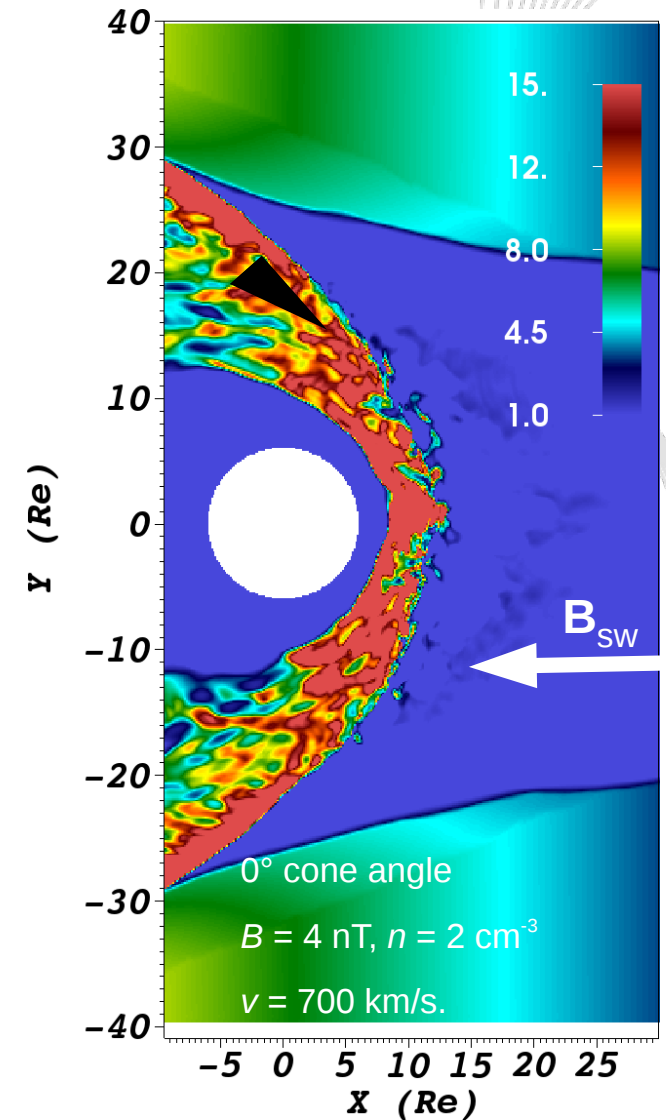
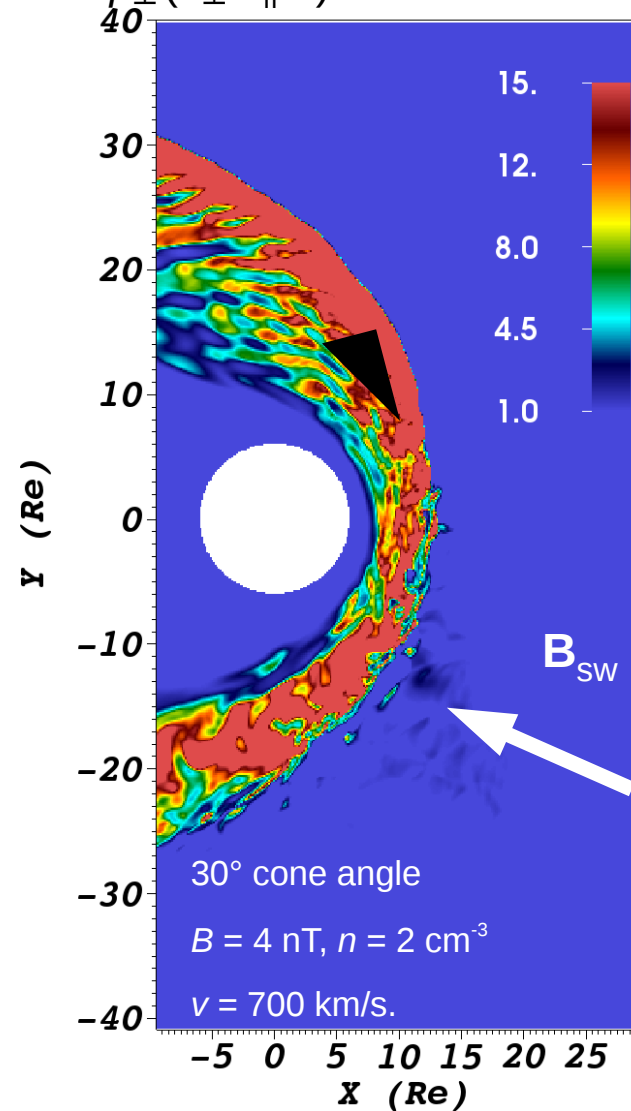
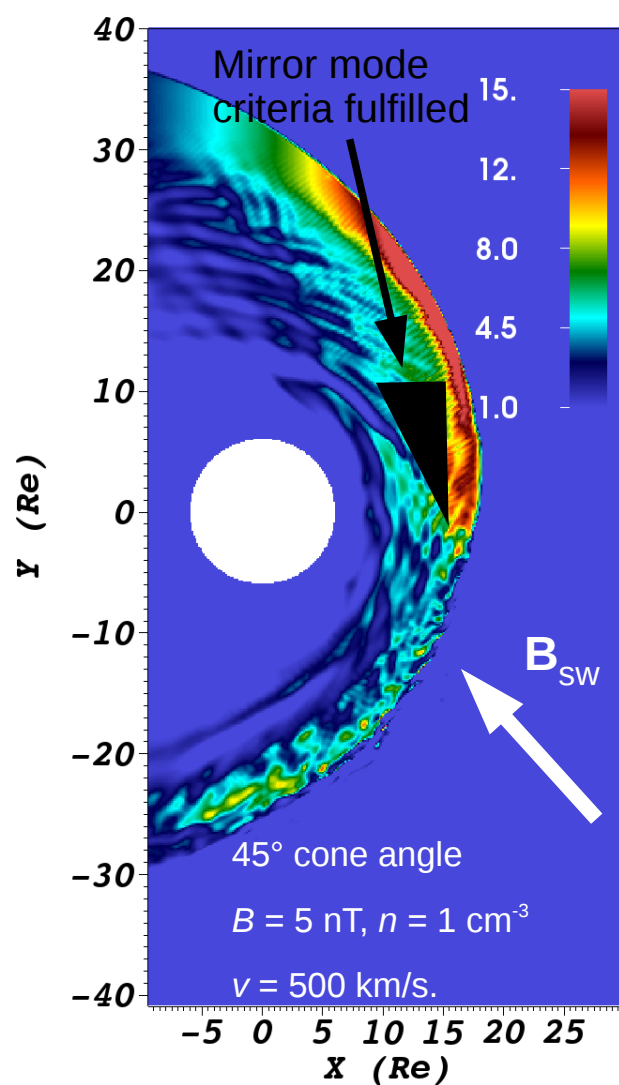
Approximate growth rate:  $\sim 0.003$





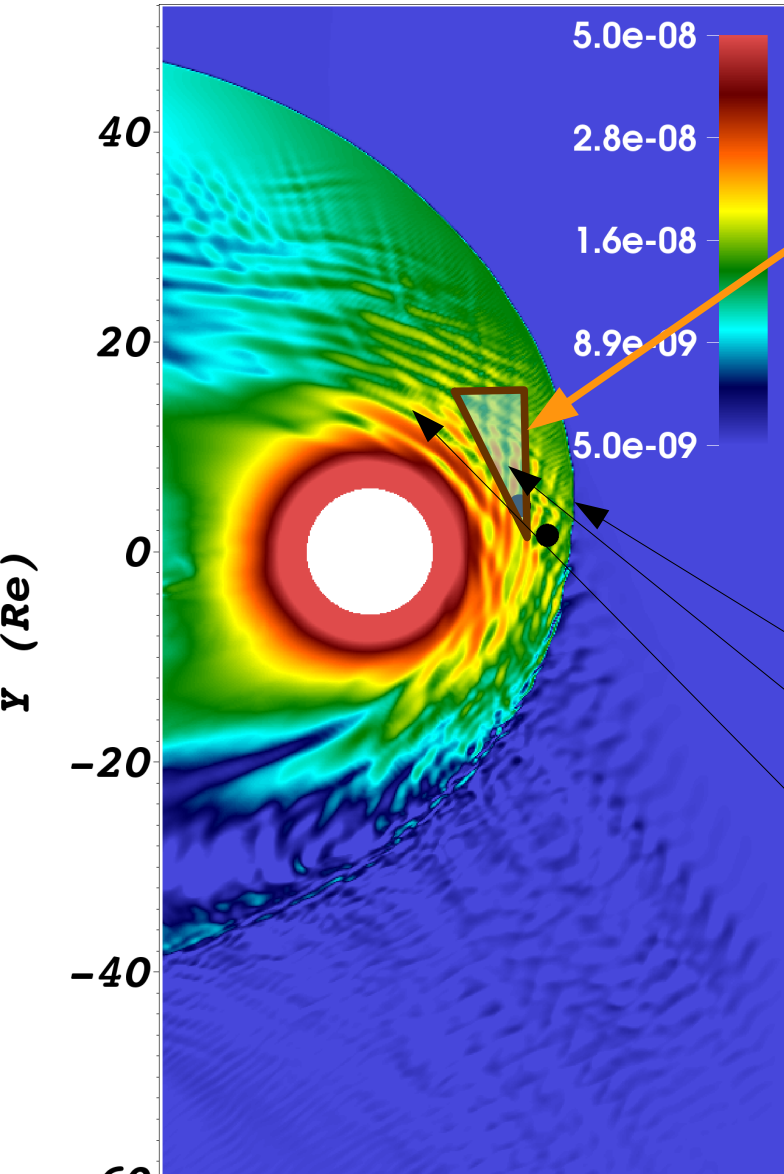
# Wave location and IMF cone angle

Color: mirror instability condition  $\beta_{\perp} (T_{\perp}/T_{\parallel}-1) > 1$  at  $t = 1000$  s





# Conclusions



- 3 runs analyzed ( $45^\circ$ ,  $30^\circ$ ,  $0^\circ$ )
- Vlasov fluctuations fulfill observational criteria:
  - Linear polarization
  - Amplitude
  - Anti-correlation between B and n
- Vlasov results in quantitative agreement with THEMIS observations
- Wave evolution:
  - Formation near the ULF wave boundary
  - Linear polarization further away from bow shock
  - Elliptical polarization further in flanks



Thank you for your attention!

[vlasiator.fmi.fi](http://vlasiator.fmi.fi)

