

# Spectral shapes of whistler-mode chorus emissions

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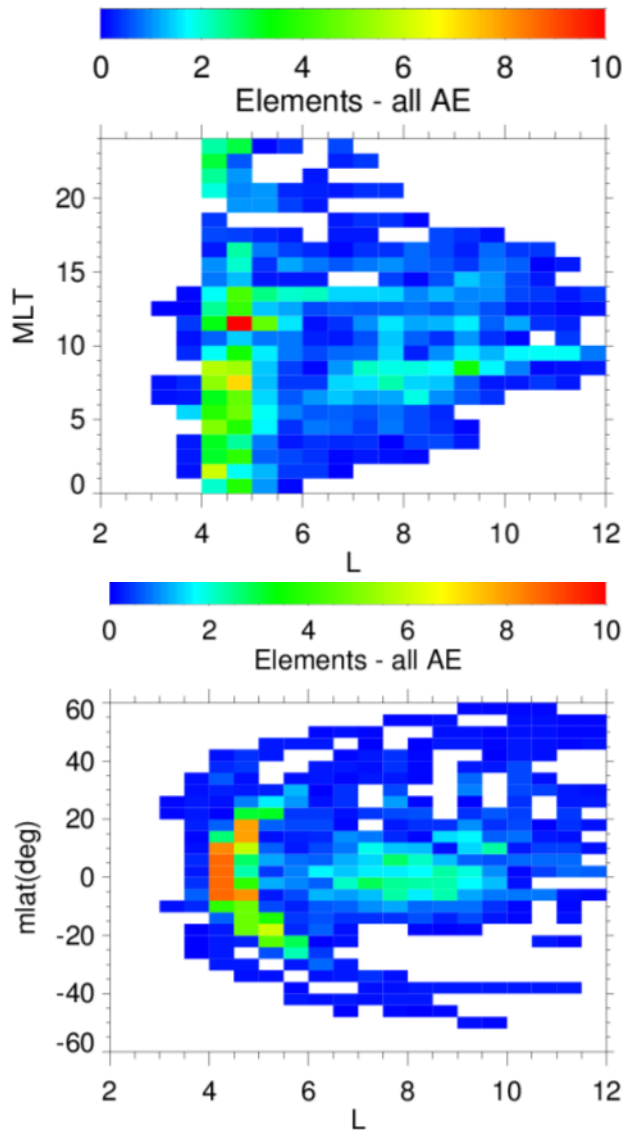
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# STAFF-SA and WBD instruments (Cluster spacecraft)



Hours of detected chorus emissions with discrete structure.

## Data set

- 11 years of Cluster measurements (2001-2012)
- list of visually selected banded emissions from the STAFF-SA instrument measurements (fulfilling a set of conditions) – almost 6000 hours
- list of visually selected chorus emission with the discrete structure from the WBD instrument measurements – almost 1300 hours

## STAFF-SA

- frequency coverage up to 4 kHz
- time resolution – 4 seconds (0.125s)
- 3 components of **B**; 2 of **E** (EFW)

## WBD

- frequency coverage up to 9.5 kHz (continuous waveforms - the instrument mode used for this analysis)
- time resolution – 37 $\mu$ s
- 1 component of **B** or **E**

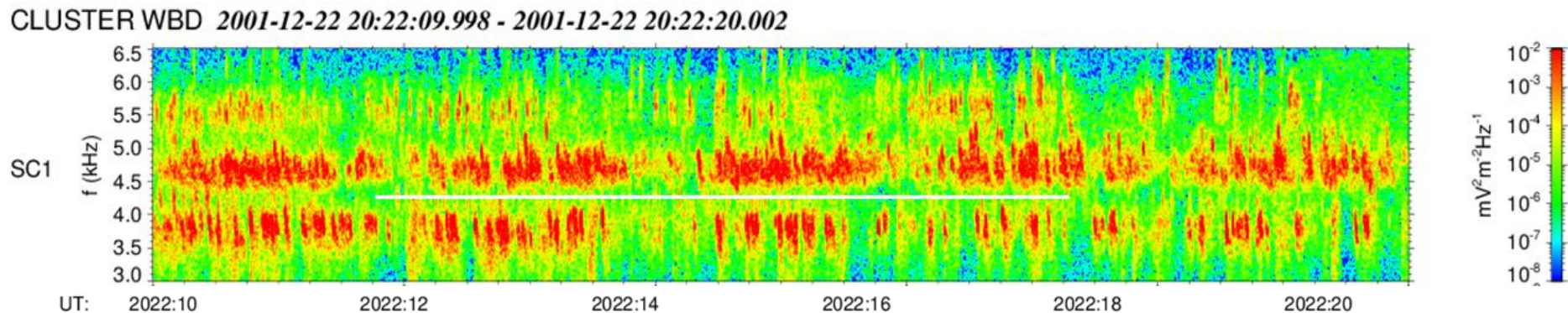
# Distribution of chorus spectral shapes

## Number of frequency bands

- one frequency band (12 %)
- **two frequency bands** (83 %)
- multi-banded emissions (three or more frequency bands) (5 %)

## Banded emissions consisting of two frequency bands

- combination of individual wave packets and hiss (54 %)
- frequency bands contain hiss (25%)
- **frequency bands contain wave packets** (21 %)



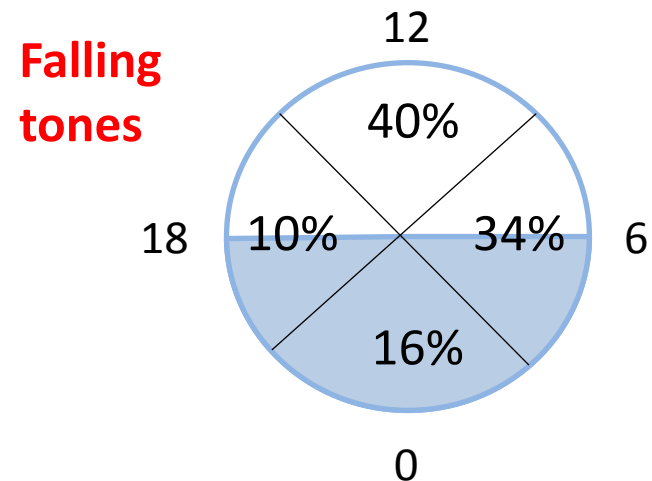
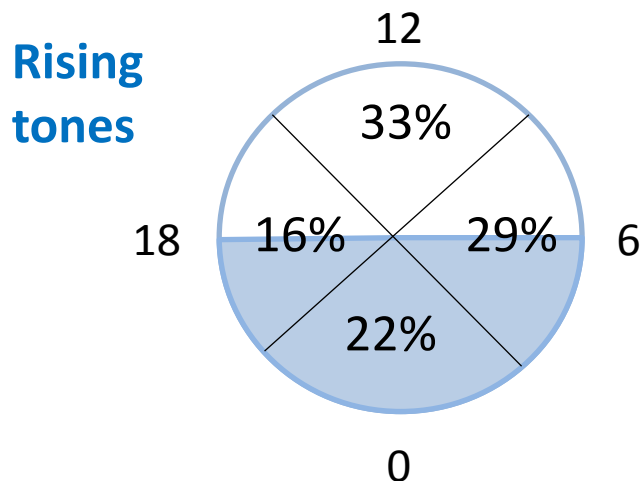
An example of time-frequency spectrograms of power spectral density (PSD) of electric field fluctuations measured by Cluster 1 on December 22, 2001. The white line corresponds to the one half of the electron cyclotron frequency. The spacecraft position is:  $\lambda_m = 2.2^\circ$ ,  $R = 4.4 R_E$  and  $\text{MLT} = 4.6$

# Distribution of chorus spectral shapes, cont.

Different types of spectral shapes of the **discrete wave packets**:

- **Rising tones (RT) –  $df/dt > 0$  (51 %)**
- **Falling tones (FT) –  $df/dt < 0$  (22%)**
- Broadband vertical wave packets  $df/dt$  (BL)  $\rightarrow \infty$  (5 %)
- **Hooks (H) (9 %)**
- Noisy elements (BP) (11%)
- Other types (mainly partly composed of combination of previous types) – (17 %) - rising and falling tones observed simultaneously (4 %) ...

## Distribution of RT and FT in the MLT sectors

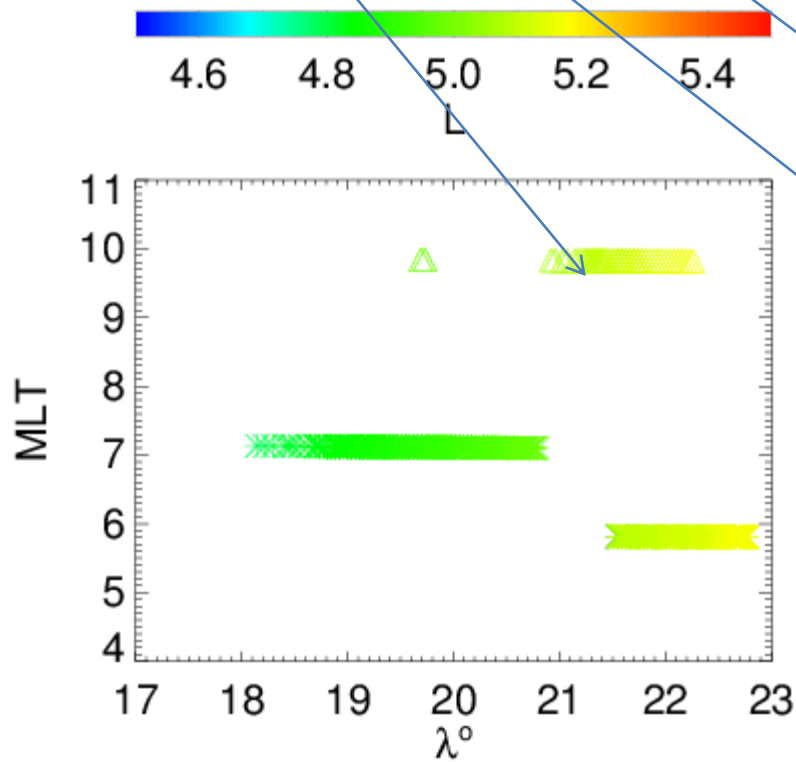


# Example events

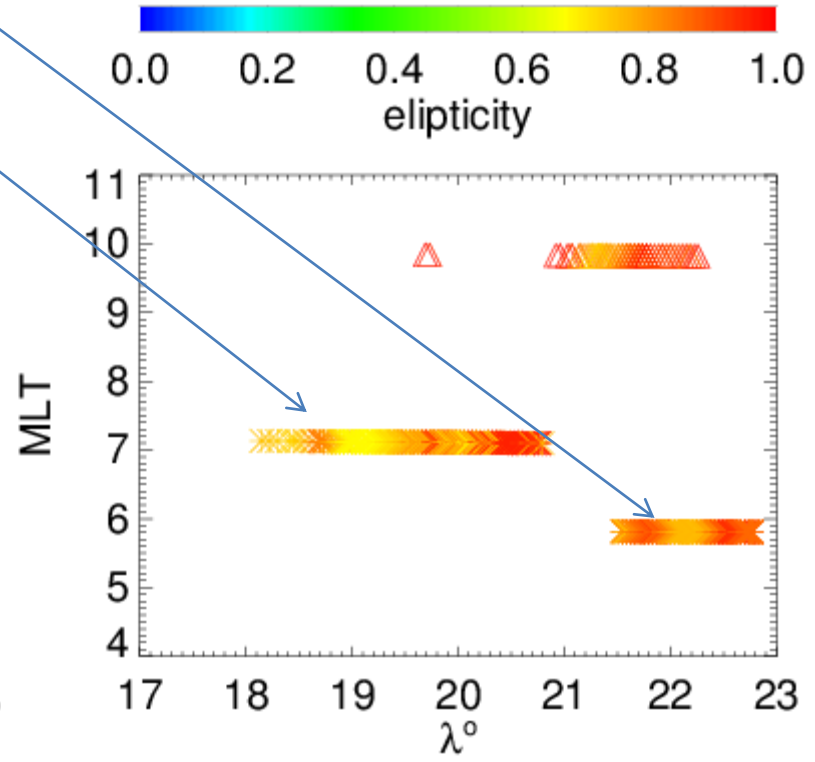
Rising tones : 2001-12-03

Falling tones: 2001-11-14

Hooks: 2001-11-07



Position of the 3 events as a function of  $\lambda$  and MLT and  $L$ .



Ellipticity of the 3 events.

# Rising tones (2001-12-03)

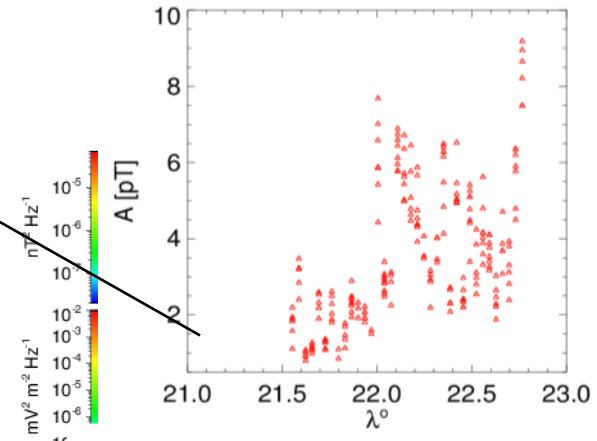
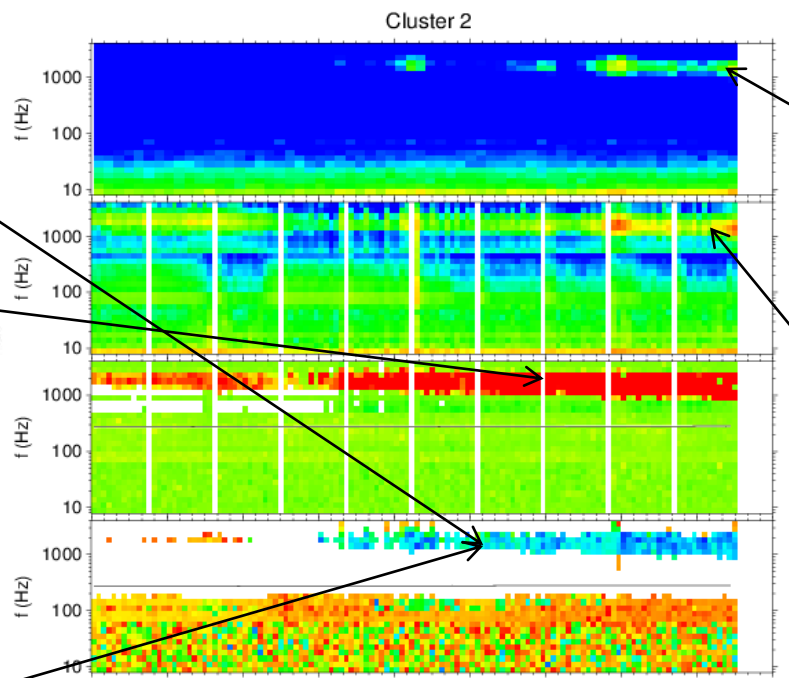
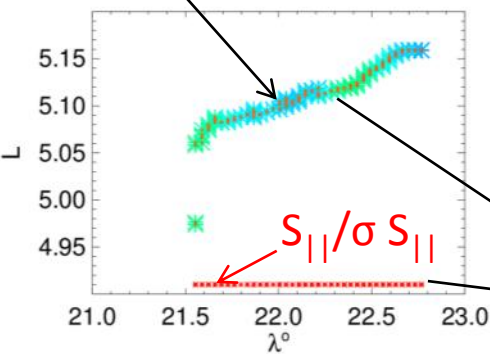
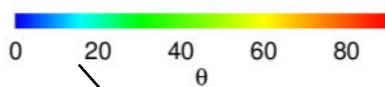
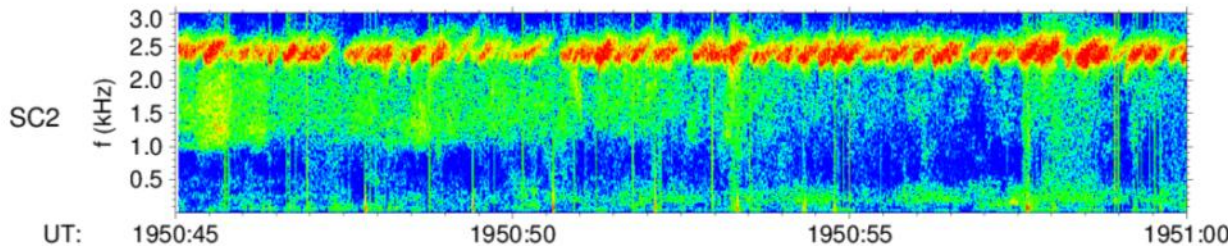
time: 19:58 – 20:04

AE index: 379-416 [nT]

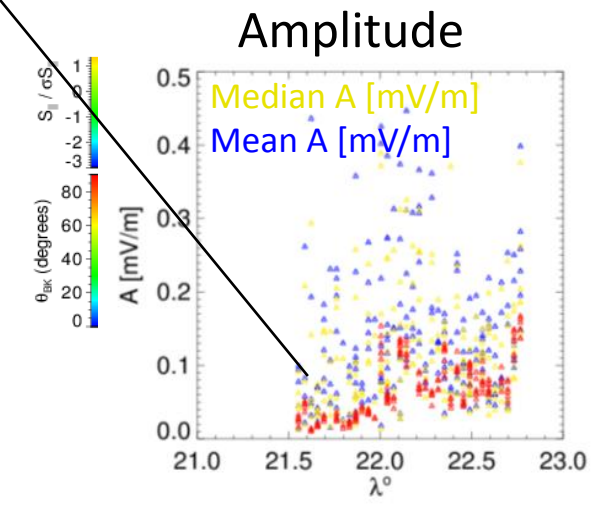
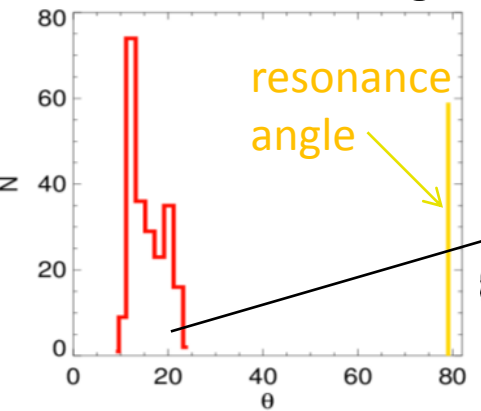
Spacecraft: C1, C2 and C4

Source position: close to the geomagnetic equator

CLUSTER WBD 2001-12-03 19:50:45.041 - 2001-12-03 19:51:00.001

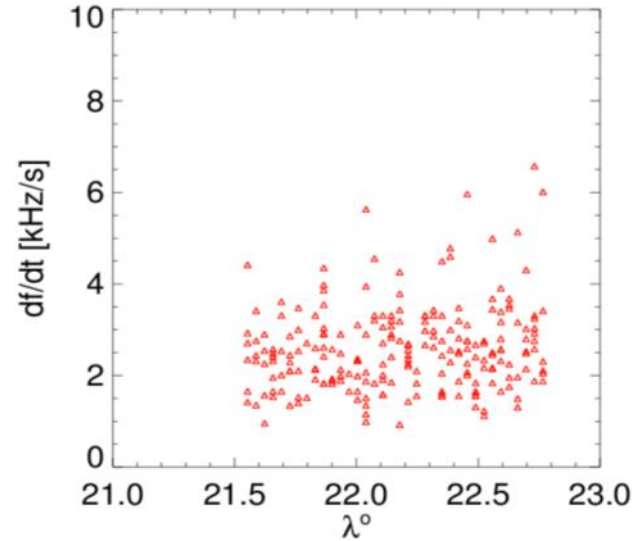
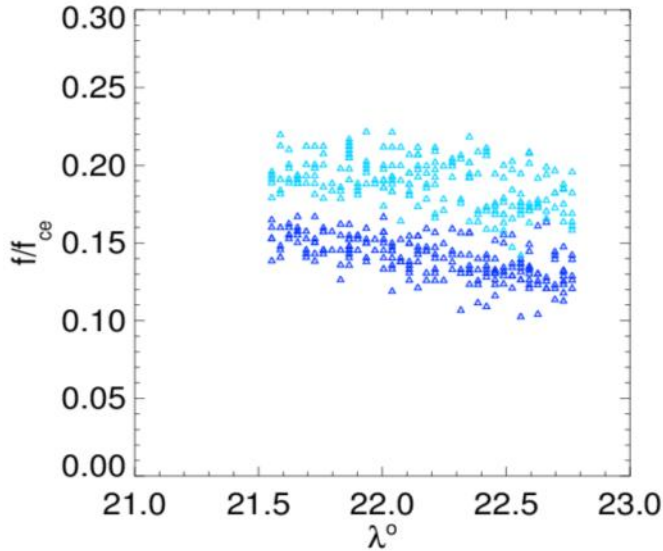


Wave normal angle



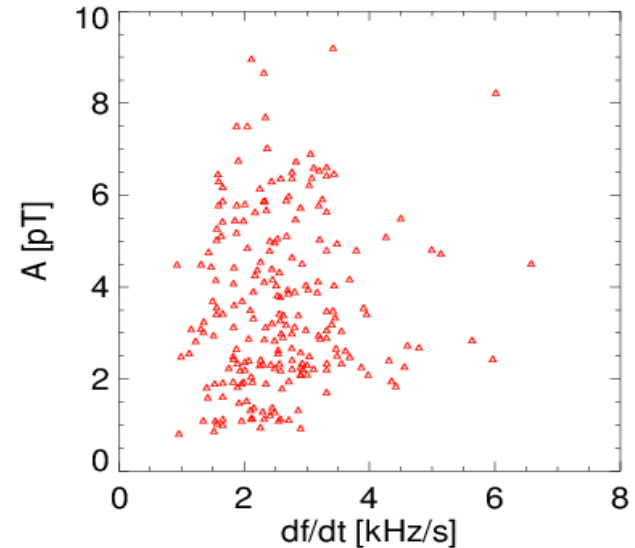
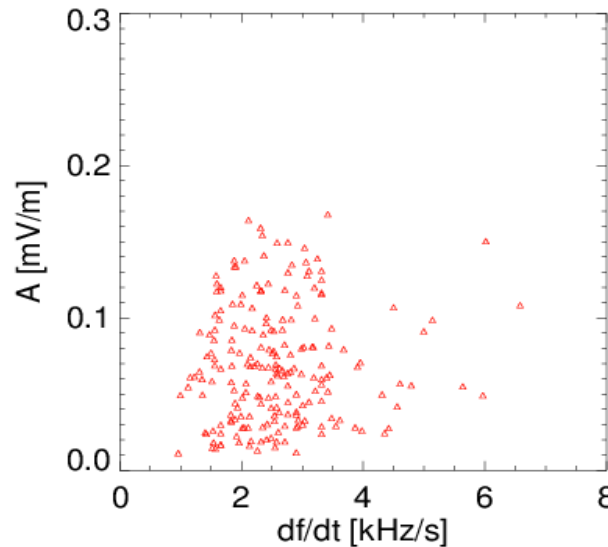
UT:	1955	1956	1957	1958	1959	2000	2001	2002	2003	2004
R (R <sub>s</sub> ):	4.40	4.40	4.41	4.42	4.42	4.43	4.44	4.45	4.46	4.46
t (deg):	20.34	20.89	21.44	21.99	22.54	23.08	23.62	24.16	24.70	25.24
ILT (h):	5.77	5.77	5.76	5.76	5.75	5.75	5.75	5.74	5.74	5.74

# Rising tones (2001-12-03), cont.



The frequency sweep rate ( $df/dt$ ) of detected elements.

Starting and ending frequencies of each individual wave packet, normalized to the local  $f_{ce}$ , as a function of  $\lambda$ .



Amplitude as a function of  $df/dt$ .

# Falling tones (2001-11-14)

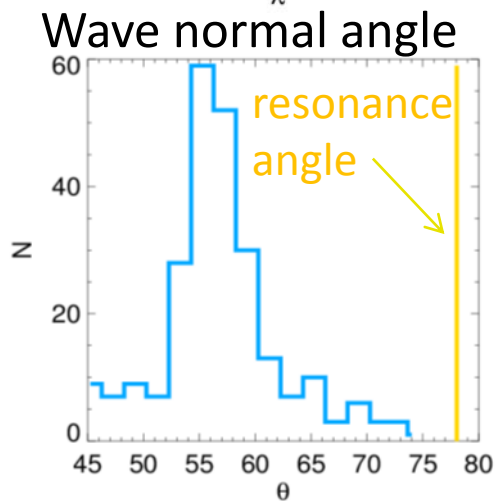
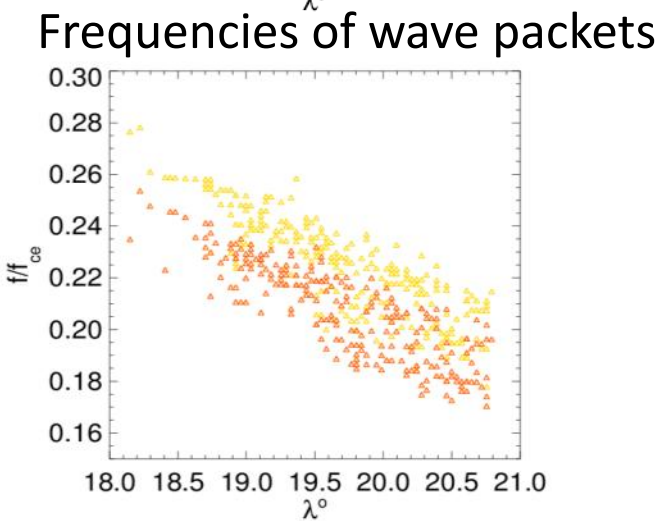
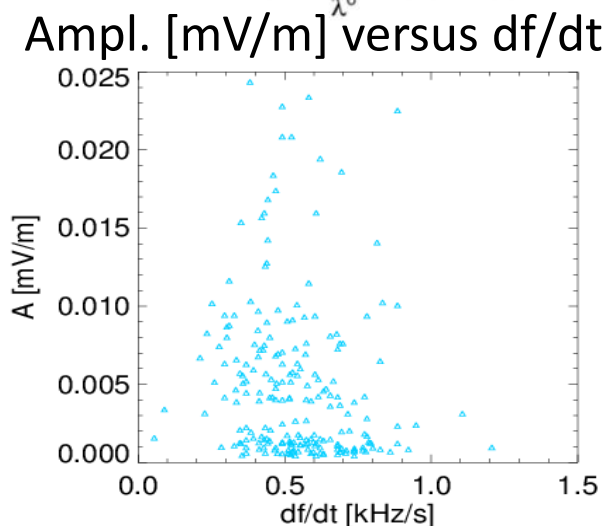
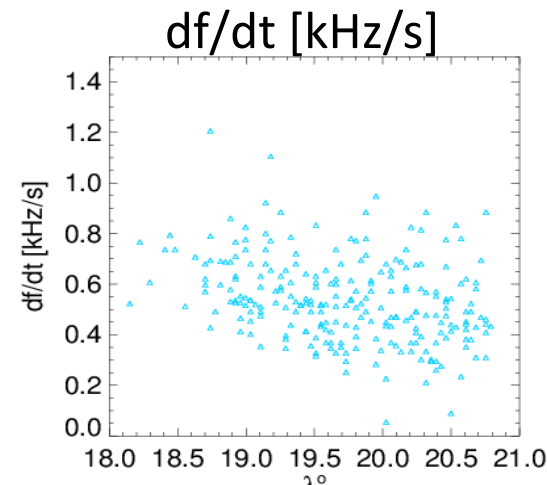
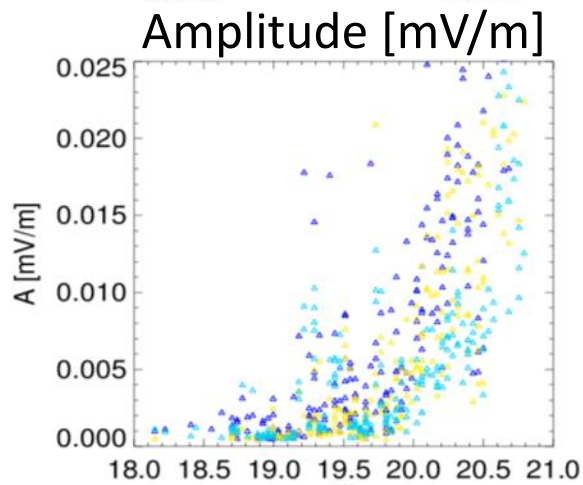
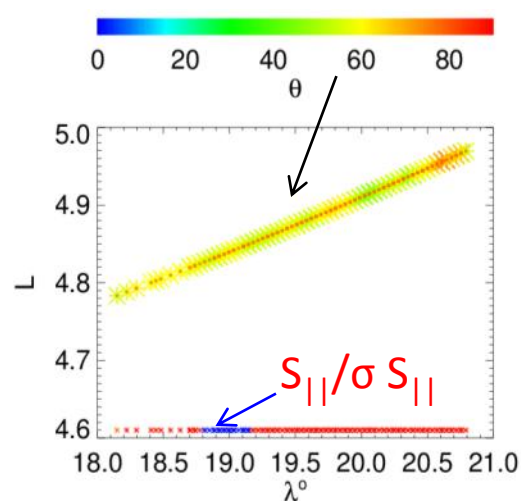
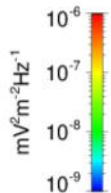
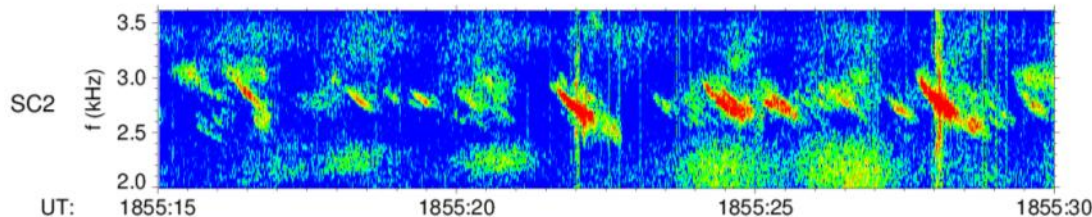
CLUSTER WBD 2001-11-14 18:55:15.029 - 2001-11-14 18:55:30.003

time: 18:53 – 18:58

AE index: 43 [nT]

Spacecraft: C2

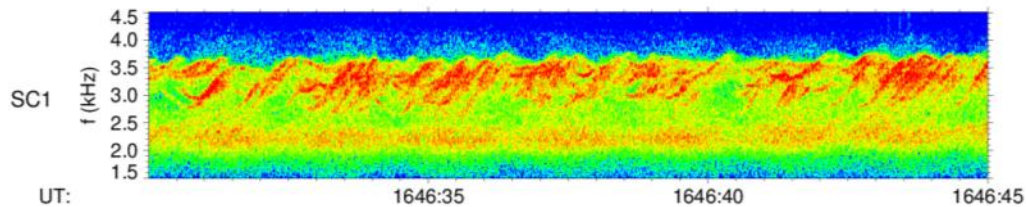
Source: close to the geom. eq.





# Hooks (2001-10-07)

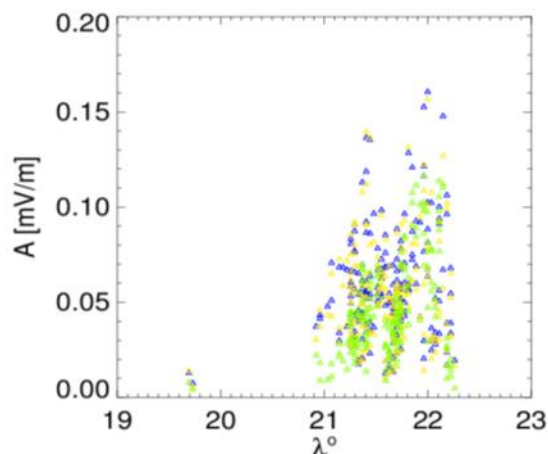
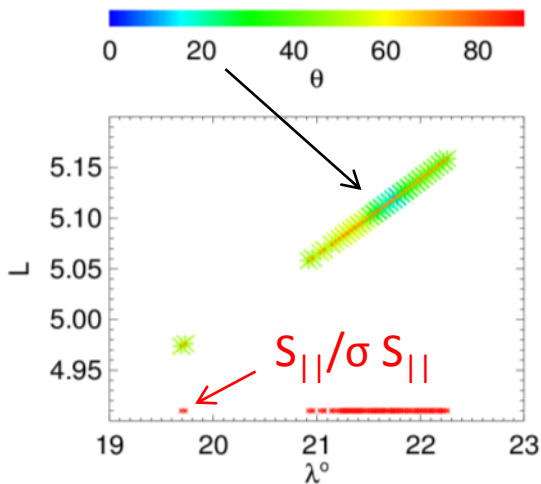
CLUSTER WBD 2001-10-07 16:46:30.014 - 2001-10-07 16:46:45.003



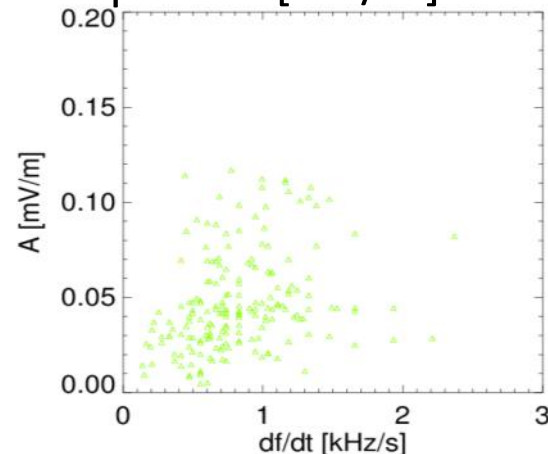
time: 16:42 – 16:49

AE index: 54 [nT]

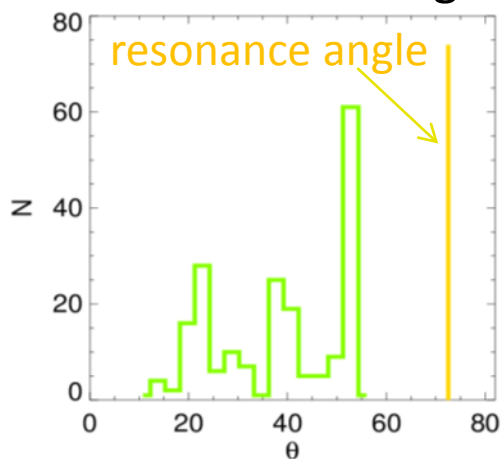
Spacecraft: C1



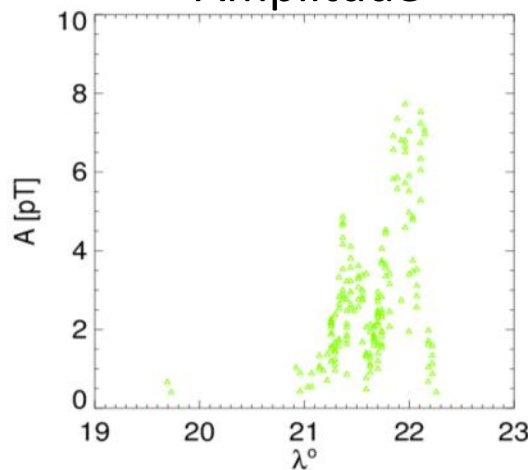
Amplitude [mV/m] versus  $df/dt$



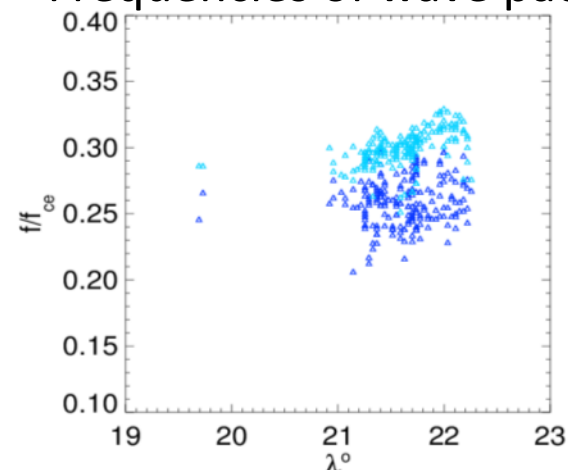
Wave normal angle



Amplitude



Frequencies of wave packets



# Results

Spectral shapes of chorus emissions have been systematically analyzed using the data of the WBD instrument on-board Cluster

The corresponding integrated amplitudes of the electric field fluctuations, the magnetic field fluctuations, the polarization, and propagation properties have been obtained from the STAFF-SA multicomponent measurements in appropriate time and frequency intervals

Rising tones were observed in 51% of cases, almost twice more often than falling tones, and five times more often than hooks or noisy elements.

Different types of spectral shapes were observed simultaneously in more than 10% of events

Examples of rising tones show quasi-parallel propagation, examples of falling tones and hooks give higher wave vector angles.

These examples also show increasing average amplitudes as a function of the magnetic latitude

*Thank you for your attention*